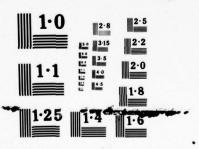


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friction terms with novel mesh-area (latitude and depth) dependent eddy-viscosity and

bottom-friction coefficients. The well-known astronomical tide-generating forces are modified by

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reffects due to solid earth tides and ocean-tidal loading. New averaged finite differences in time are used to enhance stability characteristics and to facilitate the hydrodynamical interpolation of empirical data. This unique interpolation is accomplished by a controlled adjustment of the bottom-friction coefficient and by redefining a more physical shoreline.

Extensive computer experiments were conducted to study the characteristics of the novel friction laws and hydrodynamical interpolation methods. The computed M_2 tide data along with all (specially labeled) empirical constants are tabulated in map form for four typical 30° by 50° ocean areas. A complete tabulation and discussion of the computed M_2 tide will be published in Part II of this report. It is estimated that the tabulated tidal charts permit a prediction of the M_2 -tide elevation of the ocean surface over the geoidal level with an accuracy of better than 5 cm anywhere in the open ocean and with somewhat less accuracy near rough shorelines. With the forthcoming construction of the lesser S_2 , N_2 , and S_2 ; S_1 , S_2 , and S_3 , and S_4 , and S_5 tidal constituents, the total tide-prediction error can be kept below the 10-cm bound posed by applied researchers of today.

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FOREWORD

This report gives a comprehensive account of a five-year research effort devoted to developing a global oceantide model that permits a prediction of the tidal disturbances of the geoidal sea surface with an accuracy of better than 10 cm anytime and anywhere in the world oceans. This extraordinary accuracy is required in many military and civilian applications to geodesy, geophysics, oceanography, meteorology, astronomy, and space and missile technology. In particular, in the combined NASA-DoD mission to map the geoid at sea by satellite altimetry up to 10-cm accuracy, altimeter measurements of the geoidal ocean surface must be corrected for tidal disturbances of equivalent accuracy. After numerous and elaborate computer experiments, it is concluded that the tide model arrived at meets those requirements.

It is the author's most pleasant obligation to acknowledge the sustained and generous support of the Head of the center's Strategic Systems Department, Mr. R. A. Niemann. This research work would have been impossible without the continued sponsorship of the Head of the Department's Astronautics and Geodesy Division, Mr. R. J. Anderle, and his Assistant, Dr. C. Oesterwinter, who placed the challenging problem before the author and followed its slowly progressing solution with patience and confidence.

In the current year, the ongoing project on ocean tides and currents is supported by the center's Independent Research Fund and by a grant from the National Geodetic Survey of the Department of Commerce/NOAA/NOS.* The author wishes to take this special opportunity to thank his colleagues, Dr. C. J. Cohen and Dr. B. Zondek, for many critical, stimulating, and beneficial discussions and suggestions. Grateful acknowledgment must be extended to the author's programmer, Mr. L. Szeto, for his competent and effective preparation of the involved computer programs and his untiring cooperation in carrying out numerous and sometimes extensive program changes, many of which could not be described in this report.

The date of completion was June 15, 1978.

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^{*}National Oceanographic and Atmospheric Administration (NOAA) National Ocean Survey (NOS)

ABSTRACT

A new hydrodynamical interpolation technique is developed and tested to construct a model of global ocean tides with the support of empirical tidal constants collected around the world. The discrete tide model features a 1° by 1° spherically graded grid system in connection with a hydrodynamically defined bathymetry recognizing barrier effects of large boundary and bottom anomalies. The Laplace tidal equations are augmented by turbulent friction terms with novel mesh-area (latitude and depth) dependent eddy-viscosity and bottom-friction coefficients. The well-known astronomical tide-generating forces are modified by simplified versions of secondary effects due to solid earth tides and ocean-tidal loading. Central staggered finite differences in space and new averaged finite differences in time are used to enhance decay, dispersion, and stability characteristics and to facilitate the hydrodynamical interpolation of empirical data. This unique interpolation is accomplished by a controlled adjustment of the bottom-friction coefficient and by redefining a more physical shoreline through a monitored in- or out-flow allowed across the artificial mathematical coastline.

Extensive computer experiments were conducted to study the characteristics of the novel friction laws and hydrodynamical interpolation methods and to determine by trial-and-error optimum friction, interpolation, and finite-differencing parameters. The computed M_2 -tide data along with all (specially labeled) empirical constants are tabulated in geographical map form for four typical 30° by 50° ocean areas. A complete tabulation and discussion of the computed M_2 tide will be published in Part II of this report. It is estimated that the tabulated tidal charts permit a prediction of the M_2 -tide elevation of the ocean surface over the geoidal rest level with an accuracy of better than 5 cm anywhere in the open ocean and with somewhat less accuracy near rough shorelines. With the forthcoming construction of the lesser S_2 , N_2 , and K_2 ; K_1 , O_1 , P_1 , and O_1 ; and O_1 ; and O_2 is a tidal constituents, the total tide-prediction error can be kept below the 10-cm bound posed by applied researchers of today.

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1. INTRODUCTION

The restless sea with its surfing waves, swirling currents, and swelling tides has fascinated, worried, and fright-ened mankind from the beginning. The mighty erosive and devastating power of ocean tides flooding shorelands and seaports always troubled land developers, harbor engineers, and seamen. The observed periodic feature of tides very early led laymen tidalists to record, analyze, and predict high and low tides in coastal waters and to warn and prepare people to take appropriate precautions. Indeed, quite satisfactory tide tables were constructed by simple rules of thumb. Nowadays, in seashore areas, newspapers, radio and television broadcasts, and special bulletins publish daily, monthly, and longer-range tide predictions of interest to many users.

Up to recent years, practical interest in ocean tides was essentially confined to coastal waters. With the advancement of science and technology, the need for extremely accurate tide predictions in all the world oceans has become an urgent problem. In fact, ocean tides represent fluctuating loads on the solid earth, which cause tilting of the crust and disturbances of the earth's stress and gravity fields. The precise knowledge of these loads permits researchers to determine important hydrodynamical parameters of the oceans and elastic parameters of the solid earth. Similarly, interactions between ocean tides and atmosphere, celestial bodies, and artificial satellites and missiles can be studied with high accuracy.

Recently, the National Aeronautics and Space Administration (NASA) and the Department of Defense (DoD) joined in the common mission to map the geoid at sea by satellite altimetry up to about 10-cm accuracy. However, the geoidal (rest) sea level is hydrodynamically disturbed by ocean tides and other currents due to pelagic density variations and atmospheric surface forces. Such hydrodynamical undulations of the sea surface topography reach amplitudes of more than 100 cm in open oceans and 200 cm in coastal waters. Thus, ocean tides and currents need to be determined with an accuracy compatible with the desired geoid accuracy in order to provide effective corrections for altimeter measurements.

In brief, contemporary researchers of today in geophysics, geodesy, oceanography, meteorology, astronomy, and space technology all pose the same challenging *question*:

We need to know, for any given time and place of the world oceans, the tidal elevation of the sea surface over its geoidal level within 10-cm accuracy.

In the past two decades, ocean tidalists have devoted a considerable effort to answer this extraordinary question by:

- a. Empirical methods; that is, by means of vast numbers of tidal recordings taken at continental, island, and deep-sea stations around the world.
- b. Theoretical methods; i.e., by tapping the wealth of novel mathematical tools designed by computer-oriented numerical analysts for the integration of appropriate hydrodynamical equations governing tidal currents with theoretical and/or empirical boundary and initial conditions.

Considerable progress in the qualitative mapping of ocean tides has indeed been made. Many realistic features of oceanic tides have been discovered and interpreted. This rich and extensive work will be briefly discussed in Section 2, Reviews and Previews. However, in spite of the massive empirical and theoretical effort, the satisfactory answer to the question posed above remained elusive. The constructed tidal charts vary over large ocean areas from investigator to investigator, and tide predictions fall considerably short of the desired accuracy. Nevertheless, the

published tidal maps, for instance, those by Pekeris and Accad (1969),* Zahel (1970, 1973, 1975), and Estes (1975, 1977), do suggest that the desired mapping of ocean tides might succeed with somewhat more refined models. A similar situation was recognized (Döös et al., 1974; and Reid et al., 1975) by numerical oceanographers and meteorologists modeling even more involved general ocean currents and atmospheric circulations. Obviously, the successful study of such turbulent fluid motions depends most of all on the realistic representation of the dissipative forces acting in the fluid interior and along its boundary. The special field of ocean tides, which are distinguished by their simplifying periodic property and by their widely observed boundary values, may be considered a proving ground for other models of turbulent fluid motions.

In the following treatise, a new combined hydrodynamical and empirical technique will be introduced and tested to construct, at first, the principal semidiurnal moon (M₂) tide. This procedure represents a modified version of the well-known hydrodynamical numerical method developed and tested by Hansen (1966) and his coworkers Friedrich (1966), Brettschneider (1967), Trepka (1967), and Zahel (1970, 1973, 1975), as well as by Estes (1975, 1977). Among other modifications, the most significant improvements of the results were attributed to:

- a. A unique "hydrodynamical interpolation" of more than 2 000 empirical tidal data into the original purely hydrodynamical tide model. This interpolation is accomplished by a controlled local adjustment of the bottom-friction coefficient and by allowing a monitored in- or out-flow across the mathematical ocean boundary and, thus, redefining implicitly a more physical boundary.
- b. A graded 1° by 1° grid system, in connection with a hydrodynamically defined bathymetry recognizing the important barrier effects of narrow ocean ridges.
- c. New averaged finite differences in time enhancing decay, dispersion, and stability characteristics and facilitating the implicit hydrodynamical interpolation of empirical data (a).
- d. A novel mesh-area (mesh-size and depth) dependent eddy viscosity specifying a more realistic eddy dissipation.

The complete results for the constructed M_2 tide will be published in table and map form in Part II of this report (Schwiderski, 1978 b). Similar charts for the S_2 , N_2 , and K_2 ; K_1 , O_1 , P_1 , and Q_1 ; and, possibly, Mf, Mm, and Ssa tides (Table 1) will be constructed and published as additional parts of this report. A separate tabulation (Schwiderski, 1978 a) of the new hydrodynamically defined ocean bathmetry is in preparation. All tidal and bathymetry data will be available in tape form at NSWC, Dahlgren, Virginia.

^{*}A list of references is included at the end of this report.

Table 1. Constants of Major Tidal Modes

Tidal Mode	K^a (m)	$\sigma^{\rm b}(10^{-4}/{\rm sec})$	$\chi^{c}(\text{deg})$
	Semidiurnal Species		
M ₂ = Principal Lunar	0.242 334	1.405 19	$2h_0^d - 2s_0$
S ₂ = Principal Solar	0.113 033	1.454 44	0
N ₂ = Elliptical Lunar	0.046 398	1.378 80	$2h_0 - 3s_0 + p_0$
K ₂ = Declination Luni-Solar	0.030 704	1.458 42	2h ₀
	Diurnal Species		
K ₁ = Declination Luni-Solar	0.141 565	0.729 21	h ₀ + 90
O ₁ = Principal Lunar	0.100 514	0.675 98	$h_0 - 2s_0 - 90$
P ₁ = Principal Solar	0.046 843	0.725 23	$-h_0 - 90$
Q ₁ = Elliptical Lunar	0.019 256	0.649 59	$h_0 - 3s_0 + p_0 - 90$
	Long-Period Species		
Mf = Fortnightly Lunar	0.041 742	0.053 234	2s ₀
Mm = Monthly Lunar	0.022 026	0.026 392	$s_0 - p_0$
Ssa = Semiannual Solar	0.019 446	0.003 982	2h ₀

 a_K = amplitude of the partial tide.

$$h_0 = 279.69668 + 36000.768930485T + 3.03 \cdot 10^{-4}T^2$$
,

$$s_0 = 270.434358 + 481267.88314137T - 0.001133T^2 + 1.9 \cdot 10^{-6} T^3$$
,

$$P_0 = 334.329653 + 4069.0340329575T - 0.010325T^2 - 1.2 \cdot 10^{-5} T^3$$
,

where

 $T = [27\ 392.500\ 528\ +\ 1.000\ 000\ 035\ 6D]/36\ 525,$

$$D = d + 365(y - 1975) + Int [(y - 1975)/4]$$

d = day number of year (d = 1 for January 1).

 $y \ge 1975 = year number$,

and

Int [x] = integral part of x.

 $b\sigma = frequency of the partial tide.$

 $^{^{}c}\chi$ = astronomical argument of the partial tide.

 $^{^{}d}(h_{0}, s_{0}, p_{0}) = mean longitudes of sun, moon, and lunar perigee at Greenwich midnight;$

2. REVIEW AND PREVIEW

Very early, scientists linked the periodic phenomenon of ocean tides to the apparent motions of the moon and sun. The first physical explanation of tides was given by Newton (1687), who applied his newly discovered theory of gravitation and introduced the still important "equilibrium theory" of tides (Sect. 3.D): under the attraction of the moon and the sun, the sea surface is assumed to attain instantaneously the shape of a level surface in hydrostatic equilibrium. This theory was perfected after Newton in 1738 (Cantor, 1901) in a contest conducted by the Paris Academy of Sciences. Among the leading contestants for the best mathematical and physical description of ocean tides were such distinguished scientists as D. Bernoulli, Euler, and Maclaurin. The hydrostatic equilibrium theory explained in fact the periodicity and some other observed simple features of ocean tides, but failed entirely to describe all hydrodynamical properties that were clearly observed.

Although Newton recognized the distinct hydrodynamical nature of ocean tides, it was Laplace (1775) who formulated the first hydrodynamical equations of oceanic tidal motions (Sect. 3.F). The Laplace tidal equations (LTEs) consider an inviscid and incompressible fluid subject to kinetic, potential, and Coriolis forces generated by the primary astronomical tide-producing potential, which is directly proportional to the equilibrium tide. Numerous solutions of these equations were constructed by analytic and semi-analytic methods for idealized ocean basins, estuaries, and channels. Realistic results were sought by many well-known researchers, such as Laplace (1775), Airy (1842), Ferrel (1874), Thomson (Lord Kelvin, 1879), Hough (1897), Poincaré (1910), Lamb (1915), Proudman (1915), Proudman and Doodson (1927), Platzman (1971; 1972 a, b; 1974), and Miles (1974).

Most of this rich and elucidating work has been collected and analyzed in books and reviews (see References). The conclusions clarified many realistic properties of ocean tides including the existence of so-called amphidromic points around which oceanic tidal waves rotate with amazing wave speeds (Sec. 6.B). Nevertheless, the accurate and detailed prediction of ocean tides remained unsolved. More realistic physical and geometrical models and advanced mathematical techniques were obviously needed to compute the desired ocean tides.

The theoretical models as well as the empirical predictions of ocean tides were greatly aided by Thomson (Lord Kelvin, 1868), who, following an earlier suggestion by Laplace, introduced the method of harmonic analysis into tidal studies. The astronomical tide-generating potential (equilibrium tide) is expanded into an almost periodic (with a nonharmonic frequency spectrum) series of harmonic tidal components (Sect. 3.D). Later, G. H. Darwin (1883) improved the decomposition up to 39 terms, and Doodson (1921) carried it on to 400 frequency modes by introducing his ingenious Doodson numbers. A novel nonharmonic and, subsequently, harmonic decomposition of the equilibrium tide was derived by Cartwright and Tayler (1970).

The harmonic expansion of the tide-generating potential is evidently of fundamental significance for the mathematical analysis and prediction of ocean tides. If one assumes that ocean tides are governed by linear or almost linear equations of motion (e.g., LTEs, Sect. 3.F), then the oceanic tide is also representable as a linear superposition of harmonic modes. Every harmonic component of the equilibrium tide generates, through the ocean's response, a similar component of the oceanic tide of identical frequency differing only in two time-independent constants: the amplitudes and phases. Once these harmonic constants are determined for a given geographical location either by observation or by theoretical computations, then the time-dependent tide can be predicted by simple linear superposition. Moreover, every oceanic tidal mode is hydrodynamically independent of all others and can be constructed without any knowledge of the others.

However, some nonlinear oceanic responses in coastal areas, where the sea surface varies rapidly, may be necessary in a realistic tide model (Sect. 5.F). There, one must remember that any nonlinearity generates higher harmonic frequencies and causes involved interactions between different tidal modes and also other ocean currents. In this connection, one may compare the introductory remarks to the British Admiralty (1977) tide tables.



Over many years, massive tidal records (some longer than half a century with more than 10 million hourly readings) have been made with about 10 000 tide gauges placed around the world at continental and island stations. Most of these data have been harmonically analyzed for many ocean tide components. The extracted harmonic constants have been tabulated and are now available at, for instance, the National Ocean Survey (1942), the International Hydrographic Bureau (1966), and the British Admiralty (1977).

In recent years, Eyrie (1968), Filloux (1968), and Snodgrass (1968) developed sensitive deep-sea tide (pressure) gauges that can be placed almost anywhere on the ocean floor to record tidal variations for up to one year. Munk and Cartwright (1966) introduced the so-called "ocean response method" in order to analyze relatively short-time records with sufficient accuracy. Though this method is basically nonharmonic, it too yields as its end product the most desired harmonic constants. The local time series of the ocean tide is considered as a convolution of the tide-generating potential and a transfer function (complex response weights) which is, a priori, not known and, hence, must be determined in an optimum sense (Cartwright, 1968, 1969; and Cartwright et al., 1969). By using analyzed coastal reference stations, satisfactory results are being claimed for rather short recording times of two weeks and even one week. Deep-sea measurements have been published, for instance, by Munk et al. (1970), Irish et al. (1971), Nowroozi (1972), Mofjeld (1975), Pearson (1975 a, b), and Zettler et al. (1975).

By simple inspection of numerous observed ocean-tide data and experienced intuition, cotidal (equi-phase) and some corange (equi-amplitude) maps have been constructed for marginal seas and partial or worldwide oceans by Whewell (1833, 1848), Berghaus (1845), Harris (1904), Sterneck (1920, 1921), Prüfer (1939), Dietrich (1944 a, b), Villain (1952), Bogdanov (1961 a, b), Luther and Wunsch (1974), and others. All of these maps give a qualitative overview of the phenomena of ocean tides.

A more systematic empirical method to map large-scale ocean tides was proposed and demonstrated by Kuo et al. (1970 a, b), Jachens and Kuo (1973), and Kuo and Jachens (1977). In their inversion method, modern optimization techniques are utilized to fit low-degree polynomials simultaneously to observed ocean-tide data and to measured oceanic tidal disturbances of the solid-earth gravity, which are correlated through well-known interaction equations. While this approach is definitely more reliable than purely intuitive attempts, it appears to be constrained by rather low-degree polynomials which can hardly accommodate the various tidal undulations known from observations (Sect. 6.B). The method also offers no physical insight into the tidal phenomena and yields no information on tidal currents, which is needed in certain applications. Aside from those shortcomings, the method seems to yield promising results in limited ocean regions.

With the advent of electronic computers, hydrodynamical numerical methods dominated the theoretical scene to map worldwide ocean tides. The modern, large-scale numerical investigation of ocean tides was pioneered by Hansen (1948, 1949, 1966), who realized that the complexity of natural ocean basins renders any analytic treatment intractable. Hansen was soon followed by a long line of researchers, such as Gohin (1961), Accad and Pekeris (1963), Ueno (1963, 1964), Bogdanov et al. (1964), Bogdanov and Magarek (1967, 1969), Brettschneider (1967), Tiron (1967), Trepka (1967), Pekeris and Accad (1969), Marchuk et al. (1969), Zahel (1970, 1973, 1975), Hendershott (1972, 1975), Gordeyev et al. (1973), Marchuk et al. (1973), and Estes (1975, 1977). Their probing computer experiments of steadily increasing sophistication were paralleled and supported by an even longer list of numerical analysts modeling other fluid motions, in particular closely related general ocean currents and atmospheric circulations. Among the many investigators, only very few may be named in this connection because of their important contributions with direct relevance to ocean tides: Bryan (1963), Smagorinsky (1963), Friedrich (1966, 1970), Crowly (1968, 1970), Leith (1968), Cox (1970), O'Brien (1971), and Holland and Hirschman (1972), as well as Döös et al. (1974), and Reid et al. (1975).

Hansen (1948, 1949) began his large-scale hydrodynamical numerical investigations of ocean tides with the classical LTEs (Sect. 3.F). These equations are derived from the Euler-Lagrange equations of inviscid and incompressible fluid motions by integration over the ocean depth, invoking the hydrostatic-pressure assumption and neglecting all nonlinear effects and any motion in depth (Sect. 3.E). Most of those simplifying conditions could be justified by long-wave theory. The basic LTEs are the starting point for all ocean tide investigators. The quality of their work depends essentially on their more or less realistic modeling of the following hydrodynamical and mathematical components characterizing true tidal motions.

(a) <u>Lateral Eddy Dissipation</u>. To improve his results, Hansen (1966) recognized the fundamentally turbulent nature of tidal currents and augmented the LTEs by lateral eddy-dissipation terms (Sect. 3.C). Such interior friction forces were first introduced into fluid mechanics by Boussinesq (1877), who replaced the unknown turbulent Reynolds stress tensor by the known laminar stress tensor with a constant eddy viscosity at one's disposal. In oceanography, this simple concept was brought into practice by Ekman (1902) and was later effectively used, for instance, by Munk (1950) and Bryan (1963).

Ignoring the successes of Boussinesq's unique substitution, which serves to control the supercritical instability of turbulent flow in place of the Reynolds stresses, many researchers still regard this approximation as controversial, artificial, fictitious, and phony. Apparently, this difficulty originates from the observation that the "constant" eddy viscosity varies enormously in form and magnitude from investigator to investigator. While the kinematic molecular viscosity of water is of the order 10^{-2} cm²/sec, oceanic eddy viscosities of the order 10^{0} to 10^{11} cm²/sec have been quoted (Cox, 1970; Zahel, 1970). Nevertheless, the huge variation of the eddy viscosity is physically plausible, and some empirical evidence for the lower value has been discussed by Munk (1966).

As was first realized by Prandtl (1925), the eddy viscosity is a characteristic quantity of the turbulent motion under investigation and not just a fluid coastant such as the molecular viscosity. In fact, a prescribed eddy viscosity actually specifies the "macroscopical" (averaged) features of the "microscopically undefined" turbulent flow. Consequently, the quality of the modeling of any turbulent motion is directly linked to the investigator's realistic choice of the eddy viscosity.

The physically sound meaning of Boussinesq's approximation of the Reynolds stresses was first illustrated by Prandtl's (1925) celebrated "momentum austausch (exchange, mixing) theory" of turbulent flow. With the help of constant "mixing-length parameters" (see, e.g., Schlichting, 1968), the eddy viscosity (momentum austausch coefficient) is found to be dependent on the averaged flow velocity; i.e., the eddy dissipation appears as a quadratic phenomenon. In meteorology and oceanography, velocity-dependent eddy viscosities were derived on physical grounds by Smagorinsky (1963), Leith (1968), Crowley (1968, 1970), O'Brien (1971), and others.

Considering the complexity of the world oceans, it is attractive to retain the simplifying linearity of the tidal equations. Since averaged tidal velocities can be expected to be generally slow, it appears more consistent with the overall Stokes slow-motion assumption underlying the basic LTEs to employ a velocity-independent eddy viscosity (compare bottom friction below). Also, the refined studies of Zahel (1970, 1973, 1975), Marchuk et al. (1973), and Estes (1975, 1977) do not exhibit any need for nonlinear eddy-dissipation terms. However, the author's frustrating early experiments with a worldwide, constant eddy viscosity in an ocean of depths varying from 10 to 7 000 m (Sect. 5.B) clearly indicated that the lateral eddy viscosity must depend linearly on the ocean depth (Sect. 3.C). Eddy viscosities depending on 1/2 and 3/2 powers of the depth were also tried, but they failed to yield equally good results.

A renewed look at the phenomenon of turbulent motion led to the introduction of a lateral eddy viscosity that depends linearly on a horizontal (cell or mesh size) and a vertical (cell size or depth) mixing length (Eq. 6).

This physically plausible law explains the enormous variations of eddy viscosities mentioned above. The fact that the eddy viscosity must depend in some way on the mesh size (resolution) in discrete models of ocean currents was earlier noticed by Cox (1970). Friedrich (1970), Holland and Hirschman (1972), and Zahel (1975). It may be mentioned, that the author's extensive computer experiments with the novel law of eddy dissipation resulted in rather drastic improvements of the ocean-tide model.

(b) <u>Bottom-Friction Law.</u> In addition to eddy dissipation, Hansen (1966) supplemented the LTEs also by quadratic bottom-friction terms. In fluid mechanics, the physical significance of the quadratic law of wall friction was first pointed out by Boussinesq (1896). Later, G. I. Taylor (1918) applied it advantageously in his study of tidal currents in the Irish Sea.

In order to retain the linearity of the tidal equations for their adopted time-independent numerical procedure, Pekeris and Accad (1969) used the linear law of bottom friction with a coefficient depending inversely on the ocean depth. These authors also called attention to the earlier work by Grace (1931), who applied both the quadratic and the linear laws of bottom friction to the problem of tides in the Gulf of Suez. Grace experienced a slight-preference for the linear law. Similarly to eddy dissipation (see above), it appears that the linear law of bottom friction is more consistent with all other assumed linearizations of the equations of slow (Stokes-like) averaged tidal motion (see, e.g., Schlichting (1968)). Furthermore, the convergence characterisites of the time-stepping computations carried out and displayed by Estes (1975) with the quadratic law of bottom friction fail to exhibit any nonlinear symptoms.

Other bottom-friction laws were considered by Johns (1966) and McGregor (1972), as well as by Kagan (1971 a, b). In Section 3.B, the linear law of bottom friction will be selected as preferable to the quadratic law at least over all open-ocean areas. In contrast to Pekeris and Accad (1969), the bottom-friction coefficient was found more realistic without any dependence on the ocean depth, which varies in the present model from 10 to 7 000 m. Analogous to the eddy viscosity (see (a)), the bottom-friction coefficient (as a vertical eddy viscosity) is assumed on physical arguments to depend on the bottom cell area; that is, on the two horizontal mesh sizes of the discrete tidal model considered (Eq. 4). Certain nonlinear bottom-friction effects will be introduced indirectly along coastal boundaries by the novel technique of hydrodynamical interpolation of empirical tide data described below and in Section 5.F.

(c) Primary and Secondary Tide-Generating Forces. Most numerical tidalists use only the moon's or sun's tide-generating potential as the primary driving force of ocean tides. For obvious mathematical advantages (see above), only one harmonic (mostly the largest M₂) component is investigated (Sect. 3.D). Recently recognized significant interactions between the solid earth and ocean tides motivated Farrell (1972 a, b; 1973) to suggest that accurate models of ocean tides should include tidal deformations of the solid earth, tidal loading of the ocean, and their associated perturbations of the primary tide-generating potential. Farrell went on to derive the mathematical expressions of those secondary tide-modifying forces, which should be incorporated in the forcing terms of the tidal equations.

Because of the elastic properties of the solid earth, it is sufficient for ocean-tide models to include only second-order Love number approximations (Sect. 3.D) of the terrestrial tide and its associated gravity disturbance. This simple static approximation sets both solid-earth tidal effects directly proportional to the equilibrium tide. Hence, any linear (or almost linear) ocean-tide model with homogeneous boundary data but without those second-dary driving forces can be corrected, a posteriori, by applying a uniform factor to all amplitudes. No correction of the phases is required.

In order to express the responses of the solid earth and of the gravitational field to the ocean's tidal load, Farrell derived a Green's function expanded in spherical harmonics. This representation changes the character of

the equations of oceanic tidal motion from its differential form to a much more involved integro-differential form, which must be solved by some iteration process. Farrell's suggestion was picked up by Hendershott (1972, 1975), who attributed rather drastic effects to oceanic tidal loading although his numerical iteration procedure failed to converge. The same secondary forcing functions were also used by Estes (1977) in connection with Hansen's hydrodynamical numerical technique (see (g) and Sect. 5.D) without detecting any divergence difficulties. Estes' results did not confirm the drastic effects of oceanic tidal loading as claimed by Hendershott. The computed effects on amplitudes and phases are in magnitude considerably below those modifications that are needed and that could be achieved by varying eddy-dissipation and bottom-friction parameters (Sect's. 3.B and 3.C). Nevertheless, Estes' important computations definitely supported Farrell's original contention that an accurate description of ocean tides should include ocean loading effects.

In a critical discussion of Farrell's representation of oceanic tidal loading effects, Pekeris (in a presentation given in 1977 at the Office of Naval Research) noticed that the series expansion of the Green's function is only conditionally (not absolutely) convergent and possesses only a generalized derivative of Dirac's δ -function type. To avoid all complications of this representation, Pekeris suggested a simple static treatment of all effects due to oceanic tidal loading. Analogous to Newton's equilibrium theory and to the static approximation of the solid-earth tidal effects (see above), the solid earth and the gravity field are assumed to respond instantaneously to the ocean's tidal load; i.e., the responses are directly proportional to the ocean tide.

In Section 3.D, Pekeris' straightforward solution of the rather involved problem will be incorporated in the present model. The consistent, approximate treatment of both effects due to the solid-earth tide and to the ocean's tidal load can be considered as quite satisfactory, particularly for the present ocean tide model, which will be supported by numerous empirical tide data that naturally include all solid-earth and ocean responses. In agreement with Estes' results, the present computations displayed no startling changes traceable to oceanic tidal loading.

(d) Boundary and Initial Values. In order to specify a unique integral of the LTEs, Hansen (1948, 1949) supported his tide model with empirical tide data as boundary values known around the ocean basin considered. This boundary condition is easily implemented numerically by combining all equations of oceanic tidal motion into a single elliptic, second-order differential equation for the complex tidal amplitude, provided a linear law of bottom friction and no eddy-dissipation terms are considered. Although this attractive elliptic boundary-value problem has been studied by several investigators, it must be emphasized that it leaves the fluid flow across the ocean boundaries without any control. Such an unchecked violation of the stiff condition of conservation of mass is obviously somewhat hazardous. For example, it is clearly one major cause of the divergence problems encountered by Hendershott (1975; compare also Farrell, 1972 b).

Instead of empirical tidal data, Accad and Pekeris (1963) and other authors used the purely mathematical boundary condition of no-flow across the ocean shorelines. When higher-order eddy-dissipation terms are included, two sets of boundary data are required to specify a solution. Hansen (1966), Zahel (1970, 1972, 1975), Estes (1975, 1977), and others used the mathematical boundary conditions of no-flow across and free-slip along the ocean boundaries. Some authors (e.g., Marchuk et al., 1969) replaced the free-slip condition by the no-slip condition, which is appropriately used to specify laminar flows. However, in turbulent flows with very thin boundary layers and large eddy viscosities, the free-slip condition is preferred. Moreover, this condition is consistent with the single-layer-ocean assumption of almost depth-independent horizontal velocities. The author's computer experiments with both the free-slip and the no-slip conditions exhibited a slight preference for the first one.

Some authors used a mixture of either empirical or mathematical boundary values. For example, Tiron (1967) used observed tidal elevations wherever available and the condition of to-flow across the coastline everywhere else. In any case, whenever empirical tide boundary values were incorporated, all eddy-dissipation terms had to be neglected and the condition of conservation of mass had to be violated without any possible control.

In the present tide model (Sect. 5.E), the mathematical condition of no-flow across and free-slip along the ocean boundaries will be strictly maintained wherever no empirical tide data are known (mainly at the coast of Antarctica and the Arctic Sea). To improve the quality of the tide model, known empirical tidal elevations will be incorporated at some 2 000 shore points by a unique "hydrodynamical interpolation" technique without neglecting important eddy-dissipation terms and permitting only a "monitored small" break of the no-flow boundary condition.

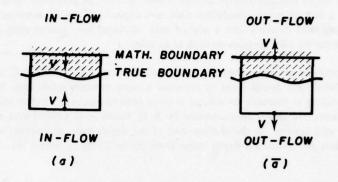
First attempts to use the empirical tide data in place of any mass-conserving condition failed to produce realistic velocity distributions; i.e., unacceptably large periodic mass flows in and out of the oceans were encountered. A second attempt to keep the empirical tide data, while enforcing strictly the no-flow boundary condition, by adjusting the bottom-friction coefficient (see (b) and Sect. 5.F) point-wise near the shore to modify the velocity components, also failed to yield realistic results.

After some evaluation of this disappointing situation, it was felt that quite satisfactory results could be obtained by a certain measured compromise. The pointwise adjustment of the bottom-friction coefficient and the resulting modification of the velocity field had to be controlled by two uniform parameters (Sect. 5.F), which remained to be determined by trial-and-error computations to yield the best results while satisfying the mass-conservation condition as closely as possible. The remaining smaller violations of the no-flow boundary condition could then be explained, for instance, by the relatively large geometric inaccuracies of the shoreline of the model (e.g., shape, size, and depth of boundary mesh cells (see (f), Sect. 5.F, and Figure 1)). Naturally, only monitored strictly periodic in- and out-flows could be permitted over any boundary segment, so that the total water mass of the ocean remains conserved. The novel technique thus developed was found to bring about the most spectacular improvements of the tide model desired.

When the complete time-dependent problem is considered, its integration must be started from some meaningful, initial oceanic state, say, at rest. The computational process can be interrupted at any time for qualitative inspection and then restarted with or without parameter changes. The integration can be stopped when the time-periodic state imposed by the harmonic forcing function is reached with satisfactory accuracy. It may be pointed out that the time-dependent integration procedure was chosen in Section 5.D as preferable because its stability analysis reveals important dispersion, decay, and stability characteristics that determine the quality of the tide model. Moreover, the experience gained with tidal time-periodic currents can be directly utilized for the modeling of more involved general ocean currents and atmospheric circulations.

(e) <u>Polar Singularities and Ice Caps</u>. In spherical coordinates, the equations of oceanic tidal motion (COTEs, Sect. 3.F) become singular at the North Pole. Most authors avoid this difficulty by terminating the ocean basin at some fictitious northern boundary. In order to check his numerical method against an analytic solution in a circular arctic ocean basin, Zahel (1970) derived an analytic integral of the equations of motion without eddy-dissipation terms. In Section 4, a unique second-order analytic solution of the complete equations of tidal motion will be constructed. This special integral will be matched with the numerical solution which starts at 4° colatitude. It may be pointed out that the use of the analytic North Pole solution allows for a more uniform grading (see (f) and Sect. 5.A) of the 1° by 1° grid system needed for the definition of the finite difference analog. As will be shown, a less uniform grid system requires considerably more computer time.

In all tide models, it is tacitly assumed that the polar ice covers have very little effect on oceanic tidal currents. This assumption is probably justified within the desired accuracy. It is certainly true at least near the North Pole, where, according to the analytic solution (Sect. 4), all semidiurnal and diurnal tides vanish with the second or first power of the distance from the Pole. Furthermore, the present tide model is aided by numerous empirical tide data collected at coastal stations of the Arctic Ocean and by some data obtained at Antarctica, which naturally contain ice-cap effects.



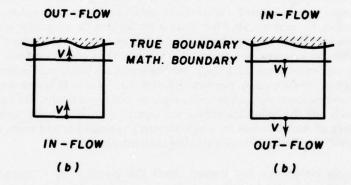


Figure 1. Boundary Cell In- and Out-Flow Illustration:
(a) and (a), also (b) and (b) Are Half-Periods Apart
(Shaded region is land area.)

(f) <u>The Gridded Ocean Basin and Bathymetry</u>. The definition of a finite-difference analog in place of the equations of oceanic tidal motion requires a mesh discretization of the ocean area under investigation. While one would like to utilize a grid system as finely detailed as possible to achieve a high resolution of the tidal currents, one is, of course, quite limited by the memory capacity of the computer facility available and by the known bathymetric charts of the ocean basins. For the worldwide oceans, essentially 1° by 1° averaged depth data were compiled by Dishon (1964), Dishon and Heezen (1968), and, independently, S. M. Smith et al. (1966). The latter collection has been somewhat revised by Gates and Nelson (1975), and both sets are available in tape form.

In spite of the specially collected depth data, all investigators prepared their own bathymetric charts corresponding to the ocean grid system chosen. All data are quite drastically curtailed by artificial limits on both the

shallow and deep ends of the depth range. Moreover, most computations were carried out on rather coarse grids of 2° by 2° to 6° by 6°. Obviously, in such an ocean geometry, much of the absolutely necessary resolution of, for example, narrow ocean ridges and continental shelves is lost. In fact, all published numerically constructed tidal maps suffer from such a lack of proper resolution. One may expect a 1° by 1° seamount or island to have little effect on the surrounding tidal currents, but a row of such obstacles over several mesh cells represents a hydrodynamical barrier separating the tidal motions on both of its sides.

Based on the experience of earlier researchers, the author (Sect. 5.B) utilizes a 1° by 1° grid system that is "graded" toward the North and South Poles to maintain a more uniform mesh area. As was demonstrated by Zahel (1970), such a grading is desirable to achieve a more uniform accuracy, higher stability, larger time steps, and shorter computer times. The depth data compiled by S. M. Smith et al. (1966) were revised and linearly interpolated to fit the new grid system. At the shallow end of the depth scale, the oceans were cut off at the 50-m depth level. The new data set retained a depth range from 50 to 7 700 m, where the upper limit remained untouched.

The author's exploratory computations of the M₂ tide used the revised S. M. Smith et al. (1966) depth data and constructed a Preliminary M₂ Tide (Schwiderski, 1976) without any empirical tidal elevations. The overall accuracy attained was encouraging. Over most ocean areas, satisfactory agreement with numerous observations at island and continental stations was achieved. Copies of this model have been requested by many research institutions for various applications. For example, Mr. Clyde Goad at the National Oceanographic and Atmospheric Administration's (NOAA) National Geodetic Survey computed the M₂-tide effects on the moon's orbit and found, for the first time, good agreement with observations and other theories. Yet, a considerable lack of accuracy, especially in the phase constants, was encountered particularly over narrow ocean ridges, such as the Aleutian, Carribbean, Marianas, and Sunda Ridges, paralleling deep trenches. Although a 1° by 1° grid system was employed, the results showed no satisfactory improvement over computations made by earlier researchers. For example, the Pacific tide was found sweeping freely into the Bering Sea without any separation effect expected from the dividing Aleutian Ridge. This striking failure of resolution may be seen in all purely computed cotidal maps, such as the well-scaled map of Zahel (1970), which also displays the corresponding observed values.

A reinspection of the bathymetric data revealed clearly that even a 1° by 1° grid scheme falls far short in representing a narrow ocean ridge in a hydrodynamically proper fashion. The defect is particularly compounded when the narrow ridge parallels a deep trench. The reason for this deficiency is obviously the purely "hydrostatic" character of the averaging principle employed by S. M. Smith et al. (1966) in order to assign a depth value at the center of a mesh cell that is supposed to be representative for the entire cell. For instance, if the area of a mesh cell is (by subjective sight) more than half land, then it is called a "land cell" and the cell is given (for the present purpose) the depth value "zero." In the alternative case, the cell is declared "oceanic" and a depth value is assigned that conserves the estimated actual water mass. Because of this "hydrostatic" principle, cells were found that contained elongated islands crossing even several cells, but every cell was declared oceanic. Moreover, an oceanic trench portion of the cell with some 7 000-m true depth produced an average depth of more than 3 500 m. Clearly, for ocean current models the entire cell represents an impassable wall and the depth value should be "zero" instead of 3 500 m.

In order to eliminate the shortcomings of the bathymetric tables compiled by S. M. Smith et al. (1966) and to implement the necessary "hydrodynamical" averaging principle recognized above (Sect. 5.B), a renewed worldwide revision of the depth data was undertaken. More than 3 000 depth data belonging to continental, island, and submerged-ridge cells were modified. In addition, the minimum and maximum cutoff-depth values were set at 10 and 7 000 m, respectively. The maximum limit affected only few isolated holes with unnecessarily adverse consequences on the stability (Sect. 5.G) of the tide model. Exploratory computations with the new ocean-bottom relief im-

mediately reflected the improved resolution. The improvement of the results was even more spectacular when empirical tidal elevations (see (d) and Sect. 5.F) were hydrodynamically interpolated.

Naturally, it must be kept in mind that the decision to declare a boundary mesh cell as oceanic and the subsequent assignment of an averaged depth value is (even with an empirically founded judgment) entirely subjective and, hence, subject to some error. While isolated inaccuracies of this sort were generally found ineffective to alter the overall quality of the tide description, errors in a series could not be tolerated in a satisfactory tide model. Serious consequences of this uncertainty are greatly reduced in the present model, which is supported by numerous tide data observed at boundary stations. In fact, in the hydrodynamical interpolation procedure (see (d) and Sect. 5.F), it was established necessary to allow some monitored periodic in- and out-flow across the mathematically fixed land side of a boundary cell. The unknown, geographically true shoreline may be farther outside or inside the mathematical cell and, thus, really require some small, periodic out- or in-flow (see Figure 1), respectively. Evidently, the permissible in- or out-flow is enhanced at river estuaries and at narrow entrances of border seas (e.g., the Mediterranean Sea), which are excluded from consideration as meshwise disconnected from the main oceanic body. The selected (Sect. 5.C) and incorporated empirical tide data, however, do reflect the true situation more closely.

(g) The Finite-Difference Analog. On the gridded ocean basin (see (f) and Sect. 5.A), the differential equations of oceanic tidal motions can be converted into a discrete, finite-difference analog for a computerized solution. Since the beginning of the computer age, numerous finite-differencing schemes with various characterisites have been developed and tested by numerical analysts working in applied mathematics, fluid mechanics, oceanography, meteorology, etc. The application of those techniques with or without possible variations to the present ocean-tide problem depends essentially on the numerical tidalist's empirically founded modeling skill and imagination, limited only by the available computer capacity. Since the forcing function (see (c) and Sect. 3.D) of the basic tide model is time-periodic, the integration at hand may be accomplished as a boundary-value problem (BVP) or as an initial-value problem (IVP).

In the chosen BVP approach, one removes the complex time-periodic factor from all three dependent variables; i.e., tidal elevation and east and north velocity components, provided the basic tide model is strictly* linear (see (a), (b), and Sect. 4). In the remaining complex (time-independent) differential equations for the complex amplitudes of the independent variables, all spacial derivatives may then be replaced by central finite differences corresponding to the grid system. As was first pointed out by Richardson (1922), the differencing scheme may be chosen "regular," "semistaggered," or (fully) "staggered;" that is, the three dependent variables may be tabulated (computed) at collocated, semidislocated, or fully dislocated points, respectively, as shown in Figure 2.

The resulting three difference equations may be combined into two equations for the complex velocity amplitudes by eliminating the complex tidal amplitude, which is directly computable from the velocity via the continuity equation. With sufficient velocity boundary conditions (no-flow and, if necessary, free-slip or no-slip; see (d)), one winds up with an implicit, complex system of linear equations that can be solved by modern computerized methods. If higher-order, eddy-dissipation terms are included in the basic tide model, the linear system to be solved becomes quite involved. Such a system has not yet been tried. However, if eddy dissipation is completely neglected (see (a) and Sect. 3.C), the linear system to be solved is considerably simpler and has been successfully inverted by Accad and Pekeris (1963) and Pekeris and Accad (1969).

The same three complex-difference equations derived above for the three dependent variables can also be combined into a single difference equation for the complex tidal amplitude, provided all eddy-dissipation terms are

^{*}If nonlinear properties (e.g., square law of bottom friction) are modeled, infinite Fourier series could, in principle, be used, but those lead to an involved infinite system of equations.

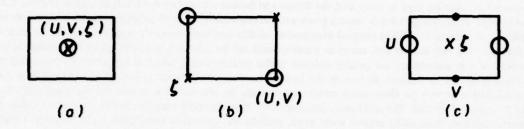


Figure 2. Finite-Difference Schemes in Space: (a) Regular, (b) Semi-staggered, (c) (Fully) Staggered (o U-Point, • V-Point, x ζ-Point)

neglected. This alternative approach to the linear BVP is obviously advantageous whenever empirical tidal-boundary values are introduced in place of the purely mathematical velocity-boundary conditions imposed in the first approach. Of course, the obtained simpler-difference analog must still be integrated by standard computer programs. This oversimplified model attracted the attention of several investigators, such as Hansen (1948, 1949), Bogdanov and Magarek (1967, 1969), and Hendershott (1972).

Since the basic tide model is time-dependent, it may be integrated as an IVP starting from some specified initial state. In addition to the advantages already pointed out under (d) above, this most natural approach leaves unrestricted room for full-scale, realistic modeling of oceanic tidal currents, including linear or nonlinear eddy dissipation (a) and bottom friction (b) as well as mathematical boundary data together with hydrodynamically interpolated empirical tidal data (d). In fact, in Section 6.A it will be argued that it is the discrete, time-dependent, finite-difference analog that reflects the physically realistic properties of oceanic tidal currents much more perfectly and ponderably than the continuous differential model. The differential equations can be used advantageously to define various finite-difference formulations, but it is neither possible nor desirable to seek convergence of the discrete solution to the continuous integral as is the objective in laminar-flow problems.

The finite-difference scheme of the BVP discussed above remains the same for the IVB in space variables. In the differential equations of oceanic tidal motion, derivatives in space variables may be replaced by central finite differences employing a Richardson regular, semistaggered, or (fully) staggered scheme. The resulting equations contain first-order derivatives in time, which can be inverted by integration over one or two time steps along the degenerated characteristic* of the system, i.e., parallel to the t-axis. The remaining (not explicitly integrable) integrals are then replaced by some average integration rule (e.g., one-point tangential, two-point trapezoidal, three-point Simpson). Depending on the chosen time-integration formula, the obtained finite-difference analog must be solved by a computer applying either an explicit or implicit time-stepping procedure that starts from an initial state and incorporates mathematical (no-flow, free-slip, or no-slip) and empirical boundary data (see (d)).

^{*}It may be suggested here that an integration along the ruled lines of the characteristic cone of the system might be preferable; but this alternative, or some mean of both possibilities, has never been tried in the modeling of ocean currents.

As was recognized by Richardson (1922), the time integration may utilize a "regular," a "step-over" ("leap-frog"), or an "intermediary" ("alternating") time-stepping scheme. In the "regular" scheme, all three (real) dependent variables (east and north velocity components and tidal height) are computed and tabulated at the same time points (Figure 3a). In Richardson's "step-over" (now often called "leap-frog") method, both velocity components are co-timed, but tidal elevations are computed and tabulated at intermediate points (Figure 3b). In the "intermediary" (or alternating) procedure, all three dependent variables are computed and tabulated coincidently, but both velocity components use computationally auxiliary intermediate time points in an alternating fashion (Figure 3c).

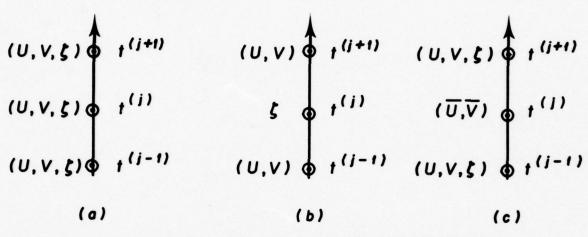


Figure 3. Finite-Difference Schemes in Time: (a) Regular, (b) Step-Over (Leap-Frog), (c) Intermediary (Alternating); (U, V, ζ) Computed and Tabulated, (U, V) Computed Not Tabulated

Although the step-over scheme is widely used in general ocean-current and atmospheric circulation studies, the regular and intermediary schemes seem to have been preferred in tidal models. The intermediary method has been applied in a rather involved explicit-implicit form by Marchuk et al. (1969), Gordeyev et al. (1973), and Marchuk et al. (1973) to tidal computations in border seas and in a global ocean model. In the now well-known hydrodynamical numerical method of Hansen (1966), the staggered scheme in space is combined with the regular scheme in time in an explicit and convenient fashlon. The method was successfully applied by Zahel (1970, 1973, 1975) and Estes (1975, 1977) to the tides in global oceans on 4°, 3°, 2°, and 1° square-grid systems.

The principles of the Hansen technique have been selected as most appropriate for the present investigation because of their simplicity and versatility. As will be shown in Section 5.E; the staggered scheme in space permits a very easy work-in of no-flow and free-slip (or no-slip) boundary data. The integration rule over one time step used by Hansen will be improved in Section 5.D by using mixed averages without losing the important explicit property of the finite-difference scheme. This modification resulted in an enhanced stability of the system. Moreover, it facilitated the implicit hydrodynamical interpolation of empirical tide data.

3. THE CONTINUOUS OCEAN TIDAL EQUATIONS (COTES)

A. THE NAVIER-STOKES EQUATIONS OF AVERAGED TURBULENT FLOW

Because of the enormous dimensions of the world oceans, their hydrodynamical tidal motion, generated by the attraction of the moon and/or sun, must be considered entirely turbulent. Accordingly, any comprehensive modeling of oceanic tidal currents should begin with the complete Navier-Stokes equations of averaged turbulent motions of a viscous, incompressible fluid including all unknown Reynolds stresses (see, e.g., Schlichting, 1968), which must provide the major friction stress to keep the intrinsic supercritical instability of the flow under control. In rotating spherical polar coordinates (λ, ϕ, r) , these equations may be written in the form (see, e.g., Whitaker, 1968)

$$u_{t} + \frac{u}{r\cos\phi} u_{\lambda} + \frac{v}{r}u_{\phi} + wu_{r} - \frac{uv}{r} - \frac{uv}{r}\tan\phi + \frac{uw}{r}$$

$$= \frac{1}{r\cos\phi} \left(q - \frac{p}{\rho} \right)_{\lambda} + 2\Omega\left(v\sin\phi - w\cos\phi \right) + f^{\lambda}, \tag{1a}$$

$$v_r + \frac{u}{r\cos\phi} v_{\lambda} + \frac{v}{r} v_{\phi} + w v_r + \frac{u^2}{r} \tan\phi + \frac{vw}{r}$$

$$= \frac{1}{r} \left(q - \frac{p}{\rho} \right)_{\phi} - 2\Omega u \sin\phi - \frac{1}{2} \Omega^2 r \sin 2\phi + f^{\phi}, \tag{1b}$$

$$w_t + \frac{u}{r \cos \phi} w_{\lambda} + \frac{v}{r} w_{\phi} + w w_r - \frac{1}{r} (u^2 + v^2)$$

$$= -\left(G + \frac{1}{\rho}P_r\right) + 2\Omega u \cos \phi + \Omega^2 r \cos^2 \phi + f^r, \tag{1c}$$

$$\frac{1}{r\cos\phi} \left[u_{\lambda} + (v\cos\phi)_{\phi} \right] + \frac{1}{r^2} \left(r^2 w \right)_r = 0, \tag{2}$$

$$\rho f^{\lambda} = \frac{1}{r \cos \phi} \tau_{\lambda}^{\lambda \lambda} + \frac{1}{r} \tau_{\phi}^{\lambda \phi} + \tau_{r}^{\lambda r} + \frac{3}{r} \tau^{\lambda r} - \frac{2}{r} \tau^{\lambda \phi} \tan \phi, \tag{3a}$$

$$\rho f^{\phi} = \frac{1}{r \cos \phi} \left[\tau_{\lambda}^{\lambda \phi} + \left(\tau^{\phi \phi} \cos \phi \right)_{\phi} \right] + \tau_{r}^{\phi r} + \frac{3}{r} \tau^{\phi r} + \frac{1}{r} \tau^{\lambda \lambda} \tan \phi, \tag{3b}$$

$$\rho f' = \frac{1}{r \cos \phi} \left[\tau_{\lambda}^{\lambda r} + (\tau^{\phi r} \cos \phi)_{\phi} \right] + \tau_{r}^{rr} + \frac{1}{r} (2\tau^{rr} - \tau^{\phi \phi} - \tau^{\lambda \lambda}). \tag{3c}$$

In these equations, all subscripts denote indicated partial derivatives, while all superscripts denote indicated components of the turbulent dissipation vector f or the Reynolds stress tensor τ . Furthermore, the following notations

= universal time in sec

= east longitude

 $= \frac{\pi}{2} - \theta = \text{north lati}$ = polar radius in m $-\theta$ = north latitude (θ = colatitude)

(u, v, w) = east. north, and radial velocities, respectively, in m/sec (averaged)= pressure (averaged) p

= total tide-generating potential

= dissipation vector f

= Reynolds stress tensor (to be specified)

= $0.72722 \times 10^{-4} \text{ sec}^{-1}$ = earth angular velocity

 $\approx 10^3 \text{ kg/m}^3 = \text{density of sea water}$

= 9.81 m/sec^2 = gravity acceleration

The equations of turbulent "mean" motion are obtained by a formal time-averaging procedure applied to the Navier-Stokes equations of viscous laminar flow that leads to the concepts of "averaged" velocities and pressure, as well as to the unknown Reynolds stress tensor, τ , containing the filtered out, fluctuating velocity residuals in quadratic form (see, e.g., Schlichting, 1968). A discussion of the physical meaning of these and following turbulence notions may be postponed to Section 6.A. The momentum equations (Eq's. 1) maintain the momentum balance between the familiar kinetic (inertial acceleration) forces on the left side and the forces of potential (tide, gravity, and pressure), Coriolis, centrifugal acceleration, and dissipation of the right. The continuity equation (Eq. 2) expresses the condition of conservation of mass. The Reynolds stress tensor, τ , and the total tide-generating potential, q, will be specified in Sections 3.C and 3.D.

B. BOTTOM AND SURFACE BOUNDARY CONDITIONS

The global oceans at hydrostatic rest may be described by

- (α) $r = R = 0.637 \times 10^7 \text{m}$ = spherical (geoidal rest) sea surface (all geoidal undulations are neglected without any significant loss of accuracy),
 - (β) $r = R H(\lambda, \phi)$ = sea-bottom relief, where
 - (γ) H = H(λ , ϕ) = realistic ocean depth in m (H(λ , ϕ) = 0 for land points, see Sect. 5.B), and
 - (δ) $p = P G\rho z$ = hydrostatic sea-pressure distribution with constant (arbitrary) sea-surface pressure P, where
 - (e) z = r R = new depth variable, so that z = 0 denotes r = R (see Figure 4).

Due to the time-dependent, tide-generating potential, q, acting on the ocean and solid earth, the hydrostatic conditions (α) – (ϵ) are altered to the following hydrodynamical boundary conditions of the sea surface and bottom (see Figure 4):

- (a) $z^s = \zeta^s = \zeta^s(\lambda, \phi, t) = \text{total sea surface tidal elevation over geoid } z = 0$;
- (b) $z^b = \zeta^b(\lambda, \phi, t) H(\lambda, \phi) = \text{sea-bottom tidal relief, where}$
- (c) $\zeta^b = \zeta^b(\lambda, \phi, t) = \text{total bottom tidal elevation over rest relief } z = -H(\lambda, \phi)$;
- (d) $\zeta = \zeta(\lambda, \phi, t) = \zeta^s \zeta^h$ = ocean tidal elevation (measured by bottom tidal pressure gauges) to be modeled:
- (e) $\xi^e = \xi^e(\lambda, \phi, t) = \text{earth tidal elevation to be specified (Sect. 3.D, Eq's. 9)};$
- (f) $\zeta^{eo} = \zeta^{eo}(\lambda, \phi, t) = \zeta^e \zeta^b = \text{earth dip-response to oceanic tidal load}, \zeta$ (Eq. 10);
- (g) $p^s = P = \text{constant surface pressure (no atmospheric pressure considered)};$
- (h) $w^s = \zeta_I^s = \text{surface radial velocity};$

- (i) $\tau^s = 0$ = no surface stress (free-slip, no wind force considered); (j) $(u^b, v^b, w^b) \cdot \nabla (H + z) \approx \xi_t^b$ = no flow across ocean bottom (∇ = gradient vector); (k) $\tau^b = \rho B(u^b, v^b)$ = bottom stress vector specified by invoking the linear law of bottom friction (Sect. 2(b)) with the coefficient

$$B = b\mu\cos\phi = \overline{b}L^2\mu\cos\phi \tag{4a}$$

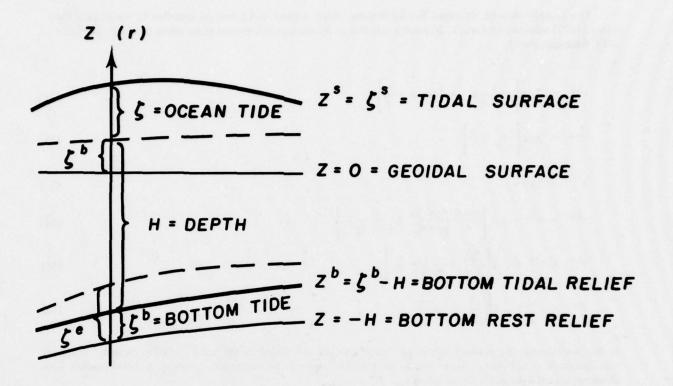


Figure 4. Earth-Ocean Tidal Interaction: $\zeta = \text{Ocean Tide}$, $\zeta^e = \text{Earth Tide}$, $\zeta^s = \text{Surface Tide}$, $\zeta^b = \text{Bottom Tide}$, $\zeta^{eo} = \zeta^e - \zeta^b = \text{Earth Dip-Response to Ocean Tide}$, and H = Ocean Depth

depending on the cell bottom area $L^2\mu\cos\phi$ (Sect. 6.A), where

- (1) L = chosen equatorial mesh size (Sect. 5.A),
- (m) μ = mesh "grading" parameter (Sect. 5.A), and
- (n) $b = \overline{b}L^2$ = uniform bottom friction parameter, which must be determined by trial-and-error computations for best results. The M₂ tide (Sect. 6.B and Part II) was computed with

$$b = 0.01 \text{ m/sec},$$
 (4b)

which is considerably below the value used by Pekeris and Accad (1969). Some implicit local adjustments of b will be allowed in Section 5.F to better accommodate observed tidal data.

(o) Lateral boundary values will be specified in Section 5.E.

C. REYNOLDS STRESSES AND EDDY DISSIPATION

The formally derived, unknown Reynolds stress tensor, τ (Sect. 3.A), may be specified by the original Boussinesq (1877) substitution (Sect's. 2(a) and 6.A); that is, by the laminar, viscous stress tensor (see, e.g., Schlichting, 1968; Whitaker, 1968):

$$\tau^{\lambda\lambda} = 2\rho A \left[\frac{1}{r\cos\phi} u_{\lambda} - \frac{v}{r}\tan\phi + \frac{w}{r} \right], \tag{5a}$$

$$\tau^{\phi\phi} = 2\rho A \left[\frac{1}{r} \mathbf{v}_{\phi} + \frac{w}{r} \right], \tag{5b}$$

$$\tau^{rr} = 2\rho A w_r \,. \tag{5c}$$

$$\tau^{\lambda\phi} = \tau^{\phi\lambda} = \rho A \left[\frac{\cos\phi}{r} \left(\frac{u}{\cos\phi} \right)_{\phi} + \frac{1}{r\cos\phi} v_{\lambda} \right] , \qquad (5d)$$

$$\tau^{\lambda r} = \tau^{r\lambda} = \rho A \left[r \left(\frac{u}{r} \right)_r + \frac{1}{r \cos \phi} w_{\lambda} \right]. \tag{5e}$$

$$\tau^{\phi r} = \tau^{r\phi} = \rho A \left[\left(\frac{\mathbf{v}}{r} \right)_r + \frac{1}{r} \mathbf{w}_{\phi} \right]. \tag{5f}$$

In this substitution, the ordinary kinetic molecular viscosity is replaced by the so-called "eddy viscosity" (momentum austausch – exchange, mixing – coefficient), which remains to be modeled to represent the true turbulent flow characteristics at hand as closely as possible.

In order to achieve a greater modeling flexibility, it is customary to divide the eddy viscosity into a "vertical" eddy viscosity associated with vertical shear and into a "horizontal" (lateral) eddy viscosity* associated with horizontal shear. In single-layer ocean models like the present one (Sect. 3.E), the separate treatment of the vertical eddy viscosity needs no explicit specification, because (Eq's. 28) the viscosity becomes an important part of the bottom-friction coefficient, B, defined by Equations 4a and 4b.

Based on the author's extensive computer experiments and on the physical arguments elaborated in Sections 2(a), 5.G, and 6.A the (horizontal) eddy viscosity, A, may now be specified by

$$A = \frac{a}{2}LH(\lambda, \phi)(1 + \mu\cos\phi). \tag{6a}$$

In this definition, L and μ denote the mesh size and grading parameters introduced in Section 3.B(1) and (m). Accordingly, the novel eddy viscosity depends on the mean lateral cross-section area $H \times L(1 + \mu\cos\phi)/2$ of the flow cell of depth H and average northeast mesh size $L(1 + \mu\cos\phi)/2$. The remaining reduced eddy-dissipation coefficient, a (in sec), must be subjected to trial-and-error computations in order to achieve best results uniformly** over

^{*}Since east and north dimensions are usually nearly equal, no need for two lateral eddy viscosities has ever been encountered.

^{**}Different a-values for the Pacific, North and South Atlantic, and Indian Oceans were also tested without significant effects.

all oceans. It may be recalled from Section 2(a) that some dependence of A on the mesh size L was noticed by several earlier researchers. Clearly, the strong variations of L and the depth H explain the huge magnitude differences of the "constant" eddy viscosity used by analysts in different problems. In the present case with H varying from 10 to 7 000 m (Eq's. 50, 51, 71, 103, and 123), the eddy viscosity varied between

$$1.3 \cdot 10^3 \frac{\text{m}^2}{\text{sec}} < A < 1.3 \cdot 10^6 \frac{\text{m}^2}{\text{sec}},\tag{6b}$$

which fits the customary range very well. It is perhaps fortuitous yet interesting to note that for a cubic cell of about 1-in. (2.54-cm.) mesh size the eddy viscosity, A (Eq's. 6a, b), reduces to the value of the molecular viscosity of water.

D. THE TOTAL TIDE-GENERATING POTENTIAL

The total tide-producing potential, q (Sect. 3.A), may be expressed in the form

$$q = G(\eta + \eta'), \tag{7}$$

where $G\eta$ is the "primary" astronomical potential directly proportional to Newton's equilibrium tide η (see Sect. 2). The remaining "secondary" potential $G\eta'$ can be expanded into its three major parts

$$G\eta' = G(\eta^o + \eta^e - \eta^{eo}),\tag{8}$$

which reveal their corresponding origins (Figure 4, Sect. 3.B(d), (e), (f), respectively);

 $G\eta^{o}$ = gravity potential perturbation due to the ocean tide ζ

 $G\eta^e$ = gravity potential perturbation due to the earth tide ζ^e

 $G\eta^{eo}$ = gravity potential perturbation due to the earth dip-response ζ^{eo} caused by the ocean load (tide) ζ

Attention to the significance of the earth-ocean tidal interactions manifested in the five quantities ξ^e , η^e , ξ^{eo} , η^{eo} , and η^o was called for by Farrell (1972 a, b; 1973). As was argued in Section 2(c), it is sufficient in ocean-tide models to use the following approximate relations:

$$\xi^e \approx 0.61\eta, \quad \eta^e \approx 0.30\eta,$$
 (9a)

$$\zeta^e - \eta^e \approx 0.31\eta. \tag{9b}$$

and

$$\zeta^{eo} - \eta^{eo} + \eta^o \approx 0.10\zeta. \tag{10}$$

The first relations (Eq's. 9) used by Hendershott (1972, 1975), Zahel (1975), Estes (1977), and others, can be justified by second-order Love number approximations. The second relation (Eq. 10) has been suggested by Pekeris (see Sect. 2(c)), who also recommended the factor 0.10 after evaluating the Green's function representation of the three oceanic tidal load effects (ξ^{eo} , η^{eo} , η^o) derived by Farrell. Apparently, the suggestion by Pekeris (Eq. 10) is physically just as plausible as the accepted approximation (Eq's. 9). In Equations 9a and 9b, it is assumed that the solid-earth tide, ξ^e , and its subsequent gravity perturbation, $G\eta^e$, are essentially instantaneous responses to the tide-generating potential, q. Analogously, in Equation 10 it is assumed that the solid-earth dip, ξ^{eo} , and the gravity perturbations, η^{eo} and η^o , are almost instantaneous responses to the ocean's tidal load, ξ . It may be mentioned that the author conducted extensive computer experiments using the factors 0.00, 0.08, and 0.12 instead

of 0.10 in Equation 10. The last two factors produced no noteworthy alterations of the results. The first factor (0.00) obviously deletes all oceanic tidal-load effects to which critically large significance has been attached by Hendershott (1972, 1975). The author's computations supported the marginal effects of oceanic tidal loading found by Estes (1977).

Following Thomson (Lord Kelvin, 1868), G. H. Darwin (1883), Doodson (1921), and Cartwright and Tayler (1970), the primary astronomical tide-generating potential, $G\eta$, or, equivalently, the equilibrium tide, η , may be expanded into a series (see, e.g., Dietrich, 1963) of "harmonic components" (constituents), η_{ν} , with a nonharmonic frequency spectrum

$$\eta = \sum_{\nu=0}^{\infty} \eta_{\nu}(\lambda, \phi, t). \tag{11}$$

In the final analysis of ocean tides, the following three major species will have to be considered:

(a) Semidiurnal Equilibrium Tides

$$\nu = 2: \eta_2 = K \cos^2 \phi \cos(\sigma t + 2\lambda + \chi) \tag{12}$$

(b) Diurnal Equilibrium Tides

$$\nu = 1: \eta_1 = K \sin 2\phi \cos(\sigma t + \lambda + \chi) \tag{13}$$

(c) Long-Period Equilibrium Tides

$$\nu = 0: \eta_0 = \frac{K}{2}(1 - 3\cos^2\phi)\cos(\sigma t + \chi)$$
 (14)

In Equations 12-14, the constants (K, σ, χ) denote

K = amplitude of equilibrium tide (in m)

 σ = frequency of equilibrium tide (in sec⁻¹)

 χ = astronomical argument of equilibrium tide (in rad)

In Table 1, these constants are listed for the major tidal modes with amplitudes larger than 4 percent of the leading semidiurnal moon (M_2) tide. The daily astronomical argument, χ , can be neglected in the following construction of the oceanic tidal modes

$$\zeta = \xi(\lambda, \phi) \cos(\sigma t + \chi - \delta(\lambda, \phi)) \tag{15a}$$

corresponding to the considered mode of the equilibrium tide, $\eta = \eta_{\nu}$. According to Equation 15a, only the "harmonic constants"

$$\xi = \xi(\lambda, \phi) = \text{tidal amplitude (in m)}$$
and
$$\delta = \delta(\lambda, \phi) = \text{Greenwich phase (in rad)}$$
(15b)

need to be found, provided the tide model is linear or almost linear (Sect. 2(g)).

E. SIMPLIFYING ASSUMPTIONS

In addition to the simplifying assumptions made concerning bottom friction (Sect. 3.B(k)-(n)), eddy dissipation (Sect. 3.C), and total tide-generating potential (Sect. 3.D), the following simplifications may be invoked:

- (a) Hydrostatic pressure assumption in Equations 1; i.e.,
- (α) Neglect all quadratic inertial accelerations (Stokes slow-motion assumption consistent with linear eddy dissipation and bottom friction (see, e.g., Schlichting, 1968))
 - (β) Neglect all centrifugal accelerations
 - (γ) Neglect vertical Coriolis force
 - (δ) Neglect vertical dissipation
 - (e) Neglect vertical motion
 - (b) Single-layer ocean assumption in Equations 1 and 2; i.e.,

(a)
$$r = R + z \approx R$$
, but $\partial r = \partial z$

$$(β)$$
 $ζ(λ, φ, t) ≪ H(λ, φ)$ $ζ^b(λ, φ, t) ≪ H(λ, φ)$

$$(\gamma) \ u(\lambda, \phi, z, t) \approx u(\lambda, \phi, t)$$

 $v(\lambda, \phi, z, t) \approx v(\lambda, \phi, t)$

It may be pointed out that all assumptions of (a) and (b) are well justified over most flow areas. In particular, the strong assumptions (b, γ) are realistic because the tide-generating potential is a body force. Moreover, the motion is fully turbulent and, hence, the averaged velocity profile exhibits only a very thin boundary layer and laminar sublayer (see, e.g., Schlichting, 1968). It may be mentioned that for the same reason the condition of free-slip at a boundary wall (Sect's. 2(d) and 5.E) seems more appropriate than the no-slip condition used in laminar-flow situations. The assumption (b, γ) is, of course, much less realistic in general ocean currents, which are driven by surface pressure and wind forces and/or by density variations confined to the upper ocean layers.

By applying assumption (a) to Equation 1c, one finds the hydrostatic pressure

$$p_z = -G\rho$$
,

i.e.,

$$p-p^{s} = \int_{\tau}^{\xi^{s}} G\rho dz = G\rho (\xi^{s} - z),$$

or, with (Sect. 3.B (a) - (g))

$$p^{s} = P,$$

$$\xi^{s} = \xi + \xi^{b} = \xi + \xi^{e} - \xi^{eo}.$$

$$P = G\rho (\xi + \xi^{e} - \xi^{eo} - z) + P.$$
(16)

With Equations 7-10 and 16, one has

$$q - \frac{p}{\rho} = G \left[\eta - (\zeta^e - \eta^e) - \zeta + (\zeta^{eo} - \eta^{eo} + \eta^o) \right] - P/\rho,$$

i.e.

$$q - \frac{p}{\rho} = G(\alpha \eta - \beta \zeta) - P/\rho. \tag{17a}$$

Where.

$$\alpha = 0.69,$$

$$\beta = 0.90. \tag{17b}$$

In the following equations, it is convenient to introduce the notations

$$\overline{H}_{\lambda} = H_{\lambda}/H, \quad \overline{H}_{\lambda\lambda} = H_{\lambda\lambda}/H,$$
 (18)

$$\overline{H}_{\phi} = H_{\phi}/H, \quad \overline{H}_{\phi\phi} = H_{\phi\phi}/H, \tag{19}$$

and

$$\widetilde{H}_{\phi} = \overline{H}_{\phi} - \mu \sin \phi / (1 + \mu \cos \phi) \tag{20}$$

so that (Eq. 6a)

$$A_{\lambda} = A \overline{H}_{\lambda} , \quad A_{\phi} = A \widetilde{H}_{\phi} . \tag{21}$$

Using the simplifications (a) and (b) and Equations 17-21 in Equations 1 and 2, one arrives at the reduced equations of motion:

$$u_t = \frac{G}{R\cos\phi} (\alpha\eta - \beta\xi)_{\lambda} + 2\Omega v \sin\phi + f^{\lambda}, \qquad (22a)$$

$$v_t = \frac{G}{R} (\alpha \eta - \beta \zeta)_{\phi} - 2\Omega u \sin \phi + f^{\phi}, \qquad (22b)$$

and

$$u_{\lambda} + (\mathbf{v}\cos\phi)_{\phi} + R\cos\phi w_{z} = 0, \tag{23}$$

where (Eq's. 3a, b; 5a-f; and 21)

$$f^{\lambda} = \frac{A}{R^{2} \cos^{2} \phi} \left[u_{\lambda \lambda} + 2\overline{H}_{\lambda} u_{\lambda} - u \right] + A u_{zz}$$

$$+ \frac{A}{R^{2}} \left[u_{\phi \phi} + (\widetilde{H}_{\phi} - \tan \phi) u_{\phi} + \widetilde{H}_{\phi} u \tan \phi \right]$$

$$+ \frac{A}{R^{2} \cos \phi} \left[(\widetilde{H}_{\phi} - 2 \tan \phi) v_{\lambda} - 2\overline{H}_{\lambda} v \tan \phi \right], \qquad (24a)$$

$$f^{\phi} = \frac{A}{R^{2} \cos^{2} \phi} \left[v_{\lambda \lambda} + \overline{H}_{\lambda} v_{\lambda} - v \right] + A v_{zz}$$

$$+ \frac{A}{R^{2}} \left[v_{\phi \phi} + (2\widetilde{H}_{\phi} - \tan \phi) v_{\phi} \right]$$

$$+ \frac{A}{R^{2} \cos \phi} \left[(2u_{\lambda} + \overline{H}_{\lambda} u) \tan \phi + \overline{H}_{\lambda} u_{\phi} \right]. \qquad (24b)$$

F. DERIVATION OF CONTINUOUS OCEAN TIDAL EQUATIONS

Under the single-layer ocean assumptions made in Section 3.E(b), the reduced Equations 22 and 23 may be integrated over the instantaneous ocean depth $(z^s - z^b = H + \zeta)$ while observing the surface- and bottom-boundary data specified in Section 3.B. For that purpose, it is useful to introduce the "integrated" velocity components:

$$U(\lambda, \phi, t) = \int_{zb}^{zs} udz \approx uH = \text{(integrated) east velocity}$$
 (25a)

and

$$V(\lambda, \phi, t) = \int_{zb}^{z^s} v dz \approx vH = \text{(integrated) north velocity,}$$
 (25b)

where the term "integrated" may be omitted when no confusion appears possible.

Using the simplifications and notations of Section 3.E, one finds the helpful approximations with the corresponding replacements $(u \leftrightarrow v, U \leftrightarrow V, \lambda \leftrightarrow \phi)$:

$$\int_{z^{b}}^{z^{s}} u_{t} dz = U_{t} - u^{s} z_{t}^{s} + u^{b} z_{t}^{b} \approx U_{t}, \tag{26a}$$

$$\int_{z^b}^{z^5} u_{\lambda} dz = U_{\lambda} - u^5 z_{\lambda}^5 + u^b z_{\lambda}^b \approx U_{\lambda} - \overline{H}_{\lambda} U, \tag{26b}$$

and

$$\int_{zb}^{z^{s}} u_{\lambda\lambda} dz \approx U_{\lambda\lambda} - \overline{H}_{\lambda} U_{\lambda} + (\overline{H}_{\lambda}^{2} - \overline{H}_{\lambda\lambda}) U - H_{\lambda} u_{\lambda}^{b}. \tag{26c}$$

The bottom-boundary conditions imposed in Section 3.B(j) and (k) assume the following approximate forms:

$$\zeta_t^b = \frac{H_\lambda}{R \cos \phi} u^b + \frac{H_\phi}{R} v^b + w^b , \qquad (27)$$

$$\frac{1}{\rho} \tau^{b\lambda} = \frac{B}{H} U \approx A \left(\frac{H_{\lambda} u_{\lambda}^b}{R^2 \cos^2 \phi} + \frac{H_{\phi} u_{\phi}^b}{R^2} + u_z^b \right), \tag{28a}$$

and

$$\frac{1}{\rho} \tau^{b\phi} = \frac{B}{H} V \approx A \left(\frac{H_{\lambda} v_{\lambda}^b}{R^2 \cos^2 \phi} + \frac{H_{\phi} v_{\phi}^b}{R^2} + v_{z}^b \right) . \tag{28b}$$

With Equations 25-28 and the surface-boundary conditions in Section 3.B(h) and (i), the proposed integration of Equations 22 and 23 is easily carried out and yields the following "continuous ocean tidal equations" (COTEs with $\theta = \pi/2 - \phi$ = colatitude):

$$U_{t} = \frac{GH}{R \sin \theta} (\alpha \eta - \beta \xi)_{\lambda} - U \left(\frac{B}{H} + \frac{A}{R^{2}} H^{\lambda} \right)$$

$$+ \frac{A}{R^{2}} \left[\frac{U_{\lambda \lambda} + \overline{H}_{\lambda} U_{\lambda}}{\sin^{2} \theta} + U_{\theta \theta} + (\cot \theta - \overline{H}_{\theta} + \widetilde{H}_{\theta}) U_{\theta} \right]$$

$$+ \frac{A}{R^{2} \sin \theta} \left[\overline{H}_{\lambda} \widetilde{H}_{\theta} V - (2 \cot \theta + \widetilde{H}_{\theta}) V_{\lambda} \right] + 2\Omega V \cos \theta, \tag{29a}$$

$$V_{t} = \frac{GH}{R} (\beta \zeta - \alpha \eta)_{\theta} - V \left(\frac{B}{H} + \frac{A}{R^{2}} H^{\theta} \right)$$

$$+ \frac{A}{R^{2}} \left[\frac{V_{\lambda \lambda}}{\sin^{2} \theta} + V_{\theta \theta} + (\cot \theta - \overline{H}_{\theta} + 2\widetilde{H}_{\theta})V_{\theta} \right]$$

$$+ \frac{A}{R^{2} \sin \theta} \left[2 \cot \theta U_{\lambda} - \overline{H}_{\lambda} U_{\theta} - (\cot \theta - \overline{H}_{\theta}) \overline{H}_{\lambda} U \right] - 2\Omega U \cos \theta, \tag{29b}$$

and

$$R\sin\theta\,\zeta_t + U_\lambda - (V\sin\theta)_\theta = 0,\tag{30}$$

where $\alpha = 0.69$, $\beta = 0.10$, and

$$H^{\lambda} = \frac{\overline{H}_{\lambda\lambda} + 1 + H_{\lambda}^{2}}{\sin^{2}\theta} + \overline{H}_{\theta\theta} - \overline{H}_{\theta}(\overline{H}_{\theta} - \widetilde{H}_{\theta}) + (\overline{H}_{\theta} + \widetilde{H}_{\theta}) \cot\theta, \tag{31a}$$

$$H^{\theta} = \frac{\bar{H}_{\lambda\lambda} + 1}{\sin^2 \theta} + \bar{H}_{\theta} + \bar{H}_{\theta} (\cot \theta - \bar{H}_{\theta} + 2\tilde{H}_{\theta}), \tag{31b}$$

$$\bar{H}_{\theta} = H_{\theta}/H, \quad \bar{H}_{\theta\theta} = H_{\theta\theta}/H,$$

$$\widetilde{H}_{\theta} = \overline{H}_{\theta} + \mu \cos \theta / (1 + \mu \sin \theta), \tag{32}$$

and

$$A = \frac{a}{2} LH(\lambda, \theta) (1 + \mu \sin \theta), \quad B = b \mu \sin \theta. \tag{33}$$

It may be mentioned that for $\alpha = \beta = 1$ and A = B = 0 the ocean-tide equations (Eq's. 29 and 30) reduce to the considerably simpler classical Laplace tidal equations. Evidently, the complete COTEs require second derivatives $(H_{\lambda\lambda}, H_{\theta\theta})$ of the bottom topography, which can be assumed to exist without placing major restrictions on the realistic features of the bathymetry. It is interesting to note that those second derivatives in H^{λ} and H^{θ} act as modifying bottom-friction terms. A ridge-like ocean floor (say, $H_{\lambda\lambda} > 0$) always adds to the bottom-friction terms (U, V) B/H, while a valley $(H_{\lambda\lambda} < 0)$ always diminishes bottom friction.

4. SECOND-ORDER ARCTIC TIDES

As was mentioned in Section 2(e), the COTEs (Eq's. 29 and 30) become singular at the North Pole. For the intended numerical procedure it is therefore advantageous to seek an approximate analytic solution around this singularity that can be matched together with the numerical solution at some appropriate colatitude. In fact, a unique "second-order arctic tide" solution can be determined for all three species of tide-generating potentials listed as Equations 12, 13, and 14, provided the ocean depth around the North Pole is assumed constant.

The COTEs (Eq's. 29 and 30) for constant depth, $H = H_0$, constant eddy viscosity, $A = A_0$, and constant bottom-friction coefficient, $B = B_0$, assume the simpler form

$$U_{t} = \frac{GH_{0}}{R \sin \theta} (\alpha \eta - \beta \zeta)_{\lambda} + 2\Omega V \cos \theta - \frac{B_{0}}{H_{0}} U$$

$$+ \frac{A_{0}}{R^{2} \sin^{2} \theta} \left[U_{\lambda \lambda} - U + \sin \theta (\sin \theta U_{\theta})_{\theta} - 2 \cos \theta V_{\lambda} \right], \tag{34a}$$

$$V_{t} = \frac{GH_{0}}{R} \left(\beta \zeta - \alpha \eta\right)_{\theta} - 2\Omega U \cos \theta - \frac{B_{0}}{H_{0}} V + \frac{A_{0}}{R^{2} \sin^{2} \theta} \left[V_{\lambda \lambda} - V + \sin \theta \left(\sin \theta V_{\theta}\right)_{\theta} + 2 \cos \theta U_{\lambda}\right], \tag{34b}$$

and

$$\xi_t + \frac{1}{R \sin \theta} \left[U_{\lambda} - (V \sin \theta)_{\theta} \right] = 0. \tag{35}$$

The forcing equilibrium tides, η (Eq's 12, 13, and 14), may be written in the unified complex form

$$\eta = \overline{\eta} (\lambda, \theta) e^{i\sigma t}, \tag{36}$$

where

$$\overline{\eta} = \overline{\eta}_{\nu} (\lambda, \theta) = \begin{cases}
K \sin^2 \theta \ e^{2i\lambda}, \ (\nu = 2) \\
K \sin 2\theta \ e^{i\lambda}, \ (\nu = 1)
\end{cases}$$
(37a)
(37b)
$$\frac{K}{2} (1 - 3 \sin^2 \theta), (\nu = 0).$$
(37c)

With the substitution

$$(\eta, \zeta, U, V) = (\overline{\eta}, \overline{\zeta}, \overline{U}, i\overline{V}) e^{i\sigma t}$$
(38)

one arrives at the following three complex differential equations in (λ, θ) :

$$\sigma \bar{U} = \frac{iGH_0}{R \sin \theta} (\beta \bar{\zeta} - \alpha \bar{\eta})_{\lambda} + 2\bar{\Omega} \bar{V} \cos \theta + i \frac{B_0}{H_0} \bar{U}$$

$$- \frac{iA_0}{R^2 \sin^2 \theta} [\bar{U}_{\lambda\lambda} - \bar{U} + \sin \theta (\sin \theta \bar{U}_{\theta})_{\theta} - 2i \cos \theta \bar{V}_{\lambda}], \qquad (39a)$$

$$\sigma \overline{V} = \frac{GH_0}{R} (\alpha \overline{\eta} - \beta \overline{\xi})_{\theta} + 2\Omega \overline{U} \cos \theta + i \frac{\overline{B}}{H_0} \overline{V}$$

$$- \frac{iA_0}{R^2 \sin^2 \theta} [\overline{V}_{\lambda\lambda} - \overline{V} + \sin \theta (\sin \theta V_{\theta})_{\theta} - 2i \cos \theta \overline{U}_{\lambda}], \qquad (39b)$$

and

$$\sigma \overline{\xi} = \frac{1}{R \sin \theta} \left[i \, \overline{U}_{\lambda} + (\overline{V} \sin \theta)_{\theta} \right]. \tag{40}$$

The reduced Equations 39 and 40 may be solved for regular solutions by power series in $\sin \theta$ (essentially polar distance) of the form:

For $\nu = 2$ and $\nu = 0$,

$$\overline{\zeta} = \zeta_0(\lambda) + \zeta_1(\lambda)\sin\theta + \zeta_2(\lambda)\sin^2\theta + \dots,$$

$$\overline{U} = U_0(\lambda) + U_1(\lambda)\sin\theta + U_2(\lambda)\sin^2\theta + \dots,$$
(41)

and

$$\bar{V} = \cos\theta \left[V_0(\lambda) + V_1(\lambda) \sin\theta + V_2(\lambda) \sin^2\theta + \ldots \right];$$

and for $\nu = 1$,

$$\overline{\xi} = \cos \theta \left[\xi_0(\lambda) + \xi_1(\lambda) \sin \theta + \xi_2(\lambda) \sin^2 \theta + \ldots \right],$$

$$\overline{U} = \cos \theta \left[U_0(\lambda) + U_1(\lambda) \sin \theta + U_2(\lambda) \sin^2 \theta + \ldots \right],$$
(42)

and

$$\overline{V} = V_0(\lambda) + V_1(\lambda) \sin \theta + V_2(\lambda) \sin^2 \theta + \dots$$

By truncating the series expansions (Eq's. 41 and 42) after the second order in $\sin \theta$ and substituting these truncations into Equations 39 and 40 up to the same quadratic power, one arrives after some lengthy but simple algebra at the following unique "second-order arctic tidal approximations."

For semidiurnal tides ($\nu = 2$),

$$\eta = K \sin^2 \theta \, e^{i(2\lambda + \sigma t)},
\xi = 6\overline{K} \sin^2 \theta \, e^{i(2\lambda + \sigma t)},
U = -2\overline{K} \sigma R \sin \theta \, e^{i(2\lambda + \sigma t)},$$
(43)

and

$$V = -i\bar{K}\sigma R \sin\theta \, e^{i(2\lambda + \sigma t)},$$

where

$$\bar{K} = \alpha G H_0 K / [6\beta G H_0 + \sigma R^2 (2\Omega - \sigma) + i\sigma (6A_0 + B_0 R^2 / H_0)]; \qquad (43a)$$

and for diurnal tides ($\nu = 1$),

$$\eta = K \sin 2\theta e^{i(\lambda + \sigma t)},$$

$$\zeta = 3\overline{K}_1 \sin 2\theta e^{i(\lambda + \sigma t)},$$

$$U = (\overline{K}_2 + 6\overline{K}_3 \sin^2 \theta) \cos \theta e^{i(\lambda + \sigma t)},$$
(44)

 $V = i \left[\overline{K}_2 + 2(\overline{K}_3 + \overline{K}_1 \sigma R) \sin^2 \theta \right] e^{i(\lambda + \sigma t)},$

where

$$\bar{K}_{3} = \frac{K_{31}K_{33}(K_{22} - K_{12})}{K_{22}(K_{13}K_{31} - K_{11}K_{32}) - K_{12}(K_{31}K_{23} - K_{21}K_{32})} ,$$

$$\bar{K}_{2} = \frac{1}{K_{12}} [K_{33} - \bar{K}_{3}(K_{13} - K_{11}K_{32}/K_{31})],$$

$$\bar{K}_{1} = -\bar{K}_{3}K_{32}/K_{31} ;$$
(44a)

and

$$K_{11} = (6\beta GH_0 - \sigma^2 R^2) + i\sigma(6A_0 + B_0 R^2/H_0),$$

$$K_{12} = -\Omega R,$$

$$K_{13} = R(6\Omega - \sigma) + \frac{i}{R}(12A_0 + B_0 R^2/H_0),$$

$$K_{21} = 6\beta GH_0 + 4i\sigma A_0,$$

$$K_{22} = R(\sigma - 2\Omega) - \frac{i}{R}(2A_0 + B_0 R^2/H_0),$$

$$K_{23} = 16iA_0/R,$$

$$K_{31} = 2\sigma\Omega R^2,$$

$$K_{32} = R(2\Omega - 3\sigma) + 3\frac{i}{R}(12A_0 + B_0 R^2/H_0),$$

$$K_{33} = 2\alpha GH_0 K.$$

$$(44b)$$

For long-period tides ($\nu = 0$),

$$\eta = \frac{K}{2} (1 - 3 \sin^2 \theta) e^{i\sigma t},$$

$$\zeta = 2\widetilde{K}_1 (2 - 3 \sin^2 \theta) e^{i\sigma t},$$

$$U = \widetilde{K}_2 \sin \theta e^{i\sigma t},$$
(45)

and

$$V = i\widetilde{K}_1 \sigma R \sin 2\theta \, e^{i\sigma t} \,,$$

where

and

$$\widetilde{K}_{3} = 4 \sigma \Omega R^{3} / [\sigma R^{2} - i(2A_{0} + B_{0} R^{2} / H_{0})],$$

$$\widetilde{K}_{2} = \widetilde{K}_{1} \widetilde{K}_{3},$$

$$\widetilde{K}_{1} = 3 \alpha G H_{0} K / 2 [(6 \beta G H_{0} - \sigma^{2} R^{2} + \Omega R \overline{K}_{3}) + i \sigma (6A_{0} + B_{0} R^{2} / H_{0})].$$
(45a)

It is important to note that the surprising uniqueness is achieved by requiring a regular integral at the North Pole (Eq's. 41 or 42) and by truncating the series of the solution and the differential equations after the second power of $\sin \theta$. Of course, without the second-order truncation, uniqueness can no longer exist. Undetermined coefficients become available to satisfy prescribed boundary data, say, at distant continental shorelines.

In the present global-tide model, the arctic tides (Eq's. 43, 44, and 45) will be considered valid up to 1° colatitude (the first land occurs at colatitude 7°; see Bathymetric Tables in Schwiderski (1978a)). For colatitudes 2° and 3°, a linear interpolation will be used to match the polar solution with the numerical solution computed south of 3° colatitude.

It is interesting to observe that if the Coriolis force and eddy dissipation are neglected ($\Omega = 0$, $A_0 = 0$), then the second-order arctic tides (Eq's. 43, 44, and 45) become exact global tides with the constants

$$\begin{split} \overline{K} &= \alpha G H_0 K / \left[(6\beta G H_0 - \sigma^2 R^2) + i \sigma B_0 R^2 / H_0 \right], \\ \overline{K}_1 &= 2 \overline{K}, \quad \overline{K}_2 = -2 \sigma R \overline{K}, \quad \overline{K}_3 = 0, \\ \widetilde{K}_1 &= \frac{3}{2} \overline{K}, \quad \widetilde{K}_2 = 0, \quad \widetilde{K}_3 = 0. \end{split}$$

From Equations 43, 44, and 45, one concludes that all second-order arctic tides, ζ , vanish at the North Pole with the same order as their corresponding driving equilibrium tides η . Hence, only long-period tides exist at the North Pole.

5. THE DISCRETE OCEAN-TIDE EQUATIONS (DOTEs)

A. THE 1° BY 1° GRADED GRID SYSTEM

With the exception of Antarctica (south of colatitude $\theta=168^\circ$), the entire (ocean and land) area of the globe is covered by a 1° by 1° grid system that is "graded" toward the poles. Each spherically rectangular mesh cell $S_{m,n}$ is bounded on the east and west by longitudes $\lambda_m=m^\circ$ and $\lambda_{m-\mu}=(m-\mu)^\circ$, respectively, and to the south and north by colatitudes $\theta_n=n^\circ$ and $\theta_{n-1}=(n-1)^\circ$, respectively, so that

$$S_{m,n} = \left\{ (m - \mu)^{\circ} \frac{(n-1)^{\circ}}{n^{\circ}} m^{0} \right\}, \tag{46}$$

where

$$m = \mu, 2\mu, \dots, (360 \to 0)$$

and $n = 1, 2, \dots, 168$.

The "grading" parameter $\mu = \mu_n$ is defined by

$$\mu = 1 \text{ for } n = 30 \text{ to } 150$$

$$\mu = 2 \text{ for } n = 15 \text{ to } 29 \text{ and } n = 151 \text{ to } 168$$

$$\mu = 4 \text{ for } n = 8 \text{ to } 14$$

$$\mu = 8 \text{ for } n = 1 \text{ to } 7$$
(48)

As was mentioned in Section 2(f), the grading of the network toward the poles is necessary, in order to maintain a more uniform mesh area for higher accuracy and stability (Sect. 5.G) of the discrete tide model. In fact, the grading equations (Eq's. 48) have been chosen in such a way that

$$\mu \sin n^{\circ} \ge \frac{1}{2} \text{ for } n = 4, 8, 15, 30, 150, 165;$$
 (49)

i.e., the southern mesh size remains larger than half the equator mesh size. This important condition is violated for $n = 1^{\circ}$, 2° , 3° and $n = 166^{\circ}$, 167° , 168° . However, in Section 4 it was pointed out that the numerically discrete tide model begins at colatitude $\theta = 4^{\circ}$. For colatitudes $\theta = 3^{\circ}$ and 2° , the numerical solution will be matched by linear interpolation to the second-order arctic-tide approximations derived in Section 4, which are assumed to be useful up to colatitude $\theta = 1^{\circ}$. As will be shown in Section 5.G, the slight violation of the condition in Equation 49 at the three southern colatitudes is not so severe as to affect the stability characteristics of the model. In any case, no need was apparent to justify an additional grading step, which is accompanied by an unnecessary, extra computational effort (Sect. 5.D).

In the global network defined above, land and ocean mesh cells are distinguished by zero or nonzero depth data which will be assigned to each cell (Sect. 5.B). The "mathematical" boundaries of the oceans follow in an obviously zigzagging fashion the mesh lines of boundary cells. However, empirical tide data known at continental and island stations will be utilized to specify indirectly more physically true boundaries (Sect's. 2(f) and 5.F).

B. HYDRODYNAMICAL OCEAN BATHYMETRY

The ocean-depth data collected by S. M. Smith et al. (1966) were rearranged and linearly interpolated to fit the new 1° by 1° graded grid system described in Section 5.A. The original data bank had to be corrected for obvious errors in continental and oceanic labeling, ocean and land signs, shorelines, and some exponents. All land elevations were set to zero including all depth data less than 50 m. Furthermore, the following meshwise-disconnected border seas were excluded from consideration by assigning zero-depth values to their corresponding mesh cells: the Baltic, Kattegat, Irish, Mediterranean, Red, Japan, Sulu, and Ceram Seas; the Hudson and Korean Bays; and the Chihlian, Persian, and Californian Gulfs.

After some negative computational experience, the depth data, which were originally defined by applying a "hydrostatical" averaging principle (see Sect. 2(f)), were revised using the following "hydrodynamical" principles:

- (a) Boundary cells at or near continental shorelines consisting of more than half oceanic areas of depths larger than 5 m were designated ocean cells, and the average oceanic depth values were assigned as the "hydrodynamically" averaged depths of the entire cells. The new depth value is preferable to the "hydrostatically" averaged depth, which preserves the actual water mass but ascribes artificially a shallow shelf character to the cell.
- (b) Island cells were declared terrestrial cells with depths zero if either the island areas were larger than half the mesh areas or the (elongated) island lengths exceeded the mesh diameters.
- (c) Island cells that remained oceanic cells were assigned depth values less than the hydrostatically averaged values. In this case and in situations of submerged seamounts or narrow ocean ridges (e.g., Aleutian, Marianas, and Carribbean), the hydrodynamical depths depended on the assessed "barrier" effects of the current obstacles: the longer and/or higher the barrier, the lesser the depth.
 - (d) The assigned minimum depth was

$$H_m = \min H(\lambda, \theta) = 20 \text{ m}, \tag{50}$$

which is lowered to 10 m by the averaging equations (Eq's. 71). The maximum depth was set at

$$H_{\mathbf{M}} = \operatorname{Max} H(\lambda, \theta) = 7000 \,\mathrm{m},\tag{51}$$

which eliminated a few isolated deeper values that unnecessarily lessened the stability of the tide model (Sect. 5.G). The averaged North Pole depth was found to be

$$H_0 = 3600 \,\mathrm{m},\tag{52}$$

which is needed to compute second-order arctic tides (Sect. 4) for colatitude line n = 1.

(e) All depth data $H_{m,n}$ are considered representative for the center of the cell $S_{m,n}$; i.e., for $m = \mu, 2\mu, \ldots$, 360 and $n = 1, 2, \ldots, 168$,

$$H_{m,n} = H(\lambda_m, \theta_n), \tag{53}$$

where $\lambda_m = (m - \mu/2)^\circ$ and $\theta_n = (n - \frac{1}{2})^\circ$.

As was already pointed out in Section 2(f), the hydrodynamically justified principles discussed above in (a)—(e) are, naturally, quite subjective and by no means free of any error. Nevertheless, some computational experiments indicated only very minor effects of isolated depth data changes. More than 3 000 depth values were changed but only very few of those required additional readjustments in order to keep some limitation on the first and second derivatives of $H(\lambda, \theta)$; i.e., on the relative differences given by Equations 71. Furthermore, the hydrodynamical interpolation of empirical tidal data (Sect. 5.F) known at continental and island stations greatly diminishes the need for precise boundary-depth data. The revised depth data bank used in the present tidal computations will be published in Schwiderski (1978 a).

C. EMPIRICAL TIDE DATA

The present tide model incorporates, by a unique hydrodynamical interpolation procedure (Sect. 5.F), empirical tidal data observed and harmonically analyzed at numerous continental and island stations. These data were taken from publications by the National Ocean Survey (1942), the International Hydrographic Bureau (1966), the British Admiralty (1977), and by Pekeris and Accad (1969), Zahel (1970, 1973), Cartwright (1971), and Luther and Wunsch (1974). Unfortunately, the most recent publication by the British Admiralty (1977) lists harmonic constants only for the four major tide components M_2 , S_2 , K_1 , and O_1 and excludes the European waters completely.

The voluminous data banks had to be screened in order to eliminate observations that are meaningless or unreliable for the present ocean-tide investigations. For example, tidal constants were excluded that were listed for stations deep inside estuaries or narrow bays (e.g., Hudson River, Bay of Fundy), at the mouth of large rivers (e.g., Amazon), between sheltering islands (e.g., Alexander Archipelago, Solomon Islands), and inside sheltering reefs (e.g., Great Barrier Reef).

About 2 500 stations were selected for further examination of their data concerning locally restricted distortions. For instance, some data taken over short distances along a coastline displayed rather drastically alternating times of high water, which are obviously meaningless for oceanic tidal studies. At many stations, different tables gave different tidal constants. Some of those discrepancies at island stations are shown for the M₂ tide in Table 2. Similarly, for some mesh cells, several different station data were available, and only one representative average had to be chosen. This situation is illustrated in Table 3 for the M₂ tide around Bermuda. Many of those differences can probably be explained as simple errors in printing or computing. For instance, the phase difference of about 1 hr at Port Galets on La Reunion Island (Table 2) seems to be due to some error in observing the correct reference time, which varies from listing to listing. Most differences, such as those shown for Bermuda stations in Table 3, are definitely true local variations. In this connection, the important tidal measurements by Gallagher et al. (1971) at Fanning Atoll in the Central Pacific may be mentioned. Tides outside and inside the small atoll's lagoon differed by about 50% (20 cm) in amplitude and by a phase lag of about 50° (1 hr, 40 min.).

In general, the most recent listings in the British Admiralty Tide Tables (1977) were chosen over older tabulations as most reliable. The selection of the data was further aided by the earlier and subsequent tidal computations. Altogether, some 1 700 M₂-tide data were selected and assigned to the centers of their respective mesh cells. Using linear interpolation and tidal computations, the total number of prescribed tide data used in the M₂-tide construction was increased to more than 2 000. Essentially all continental boundary cells carry empirically supported tide data. The empirical coverage is only marginal at Arctic and Antarctic shorelines. Most empirical tide data known at island stations are also included in the tide model. All empirical tide data will be distinguished from computed values by underlining in the printed tide tables, which will be published (see Tables 5-8) as supplemental parts of this report (Schwiderski, 1978 b).

Table 2. Empirical M₂-Tide Differences

Station		B.A.T. (77) ^a		N.O.S. (42) ^d		Others Initialed		
Latitude, Longitude		ξ ^b (m)	δ ^c (°)	- ξ(m)	δ(°)	ξ(m)	δ(°)	
Tenerife, Canary Island (A) 28°29'N 16°14'W	#	0.67	18			0.69	30	Ze
Port Praia, Cape Verde Island (A) 14°55'N 23°31'W		0.42	244			0.43	220	Ze
Ascension Island (A) 7°55'S 14°25'W		0.33	177			0.51	174	P,¹Z
St. Helena Island (A) 15°55'S 5°42'W		0.32	81	i zalitar		0.34	87	P,Z
Tristan da Cunha Island (A) 37°02'S 12°18'W		0.23	12	,		0.34	354	P,Z
Agalega Island (1) 10°26'S 56°40'E		0.29	350			0.29	290	Z
Port Galets, La Reunion Isla. (1) 20°55'S 55°17'E		0.16	302	0.14	328	0.14	328	Z
Mawson, Antarctica (1) 67°36'S 62°53'E		0.04	232			0.04	155	Z
Wilkes Station, Antarctica (I) 66°15'S 110°31'E		0.28	162			0.38	140	Z
Welles Harbor, Midway Island (P) 28°12'N 117°22'W		0.11	82			0.11	91	P,Z
Eniwetok Atoll, Marshall Is. (P) 11°21'N 162°21'E		-0.36	127	0.36	137			
lles Wallis, Fiji Island (P) 13°22'S 176°11'W		0.53	178	0.52	154			
Suva Harbor, Viti Levu, Fiji I.(P) 18°09'S 178°26'E		0.56	195	0.50	212			

^aB.A.T. (77) = British Admiralty Tables (1977).

 $b\xi$ = tidal amplitude.

 $^{^{}c}\delta$ = tidal phase relative to Greenwich.

^dN.O.S. (42) = National Ocean Survey (1942).

 $^{^{}e}Z = Zahel (1970).$

^fP = Pekeris and Accad (1969).

Table 3. Bermuda M₂-Tide Observations

36		
	359	British Admiralty (1977)
34		
	355	British Admiralty (1977)
38		
	6	British Admiralty (1977)
35		
	0	U.S. Nat. Ocean Survey (1942)
37		
	0	Pekeris & Accad (1969)
36		
	359	Zahel (1970)
36		
	358	Zettler et al. (1975)
38		
	11	J. T. Kuo Letter (1977)
	38 35 37 36 36	34 355 38 6 35 0 37 0 36 359 36 359 36 358

 $a\xi$ = tidal amplitude.

Table 4a. Deep-Sea M2-Tide Data for the Gulf of Mexico and Caribbean Sea

Station	Observed		Model		Error	
Latitude, Longitude	ξ ^a (cm)	δ ^b (°)	ξ(cm)	δ(°)	Δξ(cm)	Δδ(°)
W Florida Shelf St	7		7		0	
26.71N, 84.25W		97		92		-5
Deep Gulf St	1.3		1.6		+0.3	
24.77N, 89.65W		226		225		-1
Misteriosa Bank	8		9		+1	
18.88N, 83.81W		84	1	89		+5
Rosalind Bank	7		8		+1	
16.61N, 80.34W		107		102		-5
East Carib. St. (6-month)	0.5		1.6		+1	
16.54N, 64.88W		156		151		-5
East Carib. St. (1-month)	0.6		1.5		0.9	
16.52N, 64.91W		153		148		-5

 $a\xi$ = tidal amplitude.

 $^{{}^{}b}\delta$ = tidal phase relative to Greenwich.

 $^{^{}b}\delta$ = tidal phase relative to Greenwich.

Table 4b. Deep-Sea $\mathbf{M_2}$ -Tide Data for the Pacific and Atlantic Oceans

Station		Observed		Model		Error	
Latitude, Longitude	ξ ^a (cm)	δ ^b (°)	ξ(cm)	δ(°)	Δξ(cm)	$\Delta\delta(^{\circ})$	
Pacific St. 1 (Middleton)	110		includ	ed	NA THE		
58.76N, 145.71W		284					
Pacific St. 3 (Tofino)	99		included				
48.97N, 127.29W		239					
Pacific St. (San Francisco)	54		includ	ed			
38.16N, 124.91W		227					
Pacific St. (Josie II)	27		27		0		
34.00N, 144.99W		267		273		+6	
Pacific St. (Flicki)	43		includ	ed			
32.24N, 120.86W		149					
Pacific St. (Josie I)	43		includ	led			
31.03N, 119.80W		142					
Pacific St. (Kathy)	29		27		-2		
27.75N, 124.37W		128		130		+2	
Pacific St. (Filloux)	19		18		-1		
24.78N, 129.02W		107		105		2	
Atlantic St. 1 (N.Y. Bight)	44		includ	led			
39.32N, 64.36W		350					
Atlantic St. (N.C. St. 1)	48		46		-2		
32.69N, 75.62W		356		358		+2	
Atlantic St. (Savannah B)	88		includ	ed			
31.95N, 80.68W		15					
Atlantic St. (Scope)	45		46		+2		
30.43N, 76.42W		358		3		+5	
Atlantic St. (AOML 1)	34		35		+1		
28.14N, 69.75W		1		6		+5	
Atlantic St. (AOML 3)	34		34		0		
28.24N, 67.54W		359		4		+5	
Atlantic St. (MERT)	34		- 34		0		
27.99N, 69.67W		360		6		+6	
Atlantic St. (REIKO)	35		34		-1		
27.97N, 69.67W		1		6		+5	
Atlantic St. (EDIE-May)	32		32		0		
26.46N, 69.33W		3		7		+4	
Atlantic St. (EDIE-March)	31		32		+1		
26.45N, 69.32W		1		7		+6	

 $^{^{}a}\xi$ = tidal amplitude. $^{b}\delta$ = tidal phase relative to Greenwich.

Naturally, it must be remembered that the selection of representative empirical tidal data (compare depth data, Sect. 5.B) is not at all free of subjective judgment and may be somewhat erratic. Obviously, only future additional tidal measurements can improve this model in this respect. Nevertheless, according to the instruction notes accompanying the British Admiralty Tide Tables (1977), it can probably be assumed that almost all important tide data selected carry an accuracy that is at least as high as the desired 10 cm specified in Section 1. In any case, computational experiments showed that isolated reasonable variations of the boundary-tide data do not affect significantly the adjacent oceanic tides. It was also found insignificant to the overall quality of the tide model whether the empirical data were assigned to the centers or to the shore boundaries of the respective cells (see Part II).

Attempts were made to incorporate also recent deep-sea tidal measurements (Sect. 2) into the present model. Since the hydrodynamical interpolation of empirical data is essentially based on bottom and boundary irregularities (see Sect. 5.F(a)-(d), (\overline{a})), no physically valid justification was found to include distant offshore deep-sea measurements into the model. However, some deep-sea measurements near rough shore and bottom areas were included. Fortunately, without exception, all excluded offshore deep-sea measurements known to the author agree very well with the computed M_2 -tide data (see Table 4).

D. DERIVATION OF DISCRETE OCEAN-TIDE EQUATIONS

Following essentially the hydrodynamical numerical method of Hansen (1966) and Zahel (1970), the COTEs (Eq's. 29 and 30) may now be converted to an explicit finite-difference analog called "discrete ocean-tide equations" (DOTEs). In a first step, all spacial derivatives will be replaced by divided central finite differences using the Richardson (1922) staggered scheme (Sect. 2(g)) illustrated in Figure 2c. In agreement with the graded grid system defined in Section 5.A, it is convenient to introduce the following notations:

$$\Delta \theta = \pi/360 = 1/2^{\circ} \text{ mesh size}, \tag{54}$$

$$\Delta \lambda = \mu \Delta \theta , \quad L = 2R\Delta \theta, \tag{55}$$

 Δt = time step to be specified;

and for a fixed, oceanic mesh cell $S_{m,n}$ (Eq's. 46, 47, 48) at u-points (Figure 2c)

$$\lambda_{m}(u) = 2(m - \mu)\Delta\lambda,$$

$$\theta_{n}(u) = (2n - 1)\Delta\theta,$$

$$U_{m,n}^{(t)} = U(\lambda_{m}(u), \theta_{n}(u), t),$$

$$\eta(t,u) = \eta(\lambda_{m}(u), \theta_{n}(u), t),$$

$$\Gamma_{n}(u) = 1/\mu \sin\theta_{n}(u);$$
(56)

at v-points (Figure 2c)

$$\lambda_{m}(\mathbf{v}) = (2m - \mu)\Delta\lambda,$$

$$\theta_{n}(\mathbf{v}) = 2n\Delta\theta,$$

$$V(t) = V(\lambda_{m}(\mathbf{v}), \theta_{n}(\mathbf{v}), t),$$

$$\eta(t, \mathbf{v}) = \eta(\lambda_{m}(\mathbf{v}), \theta_{n}(\mathbf{v}), t),$$

$$\Gamma_{n}(\mathbf{v}) = 1/\mu \sin\theta_{n}(\mathbf{v});$$

$$(57)$$

and at 5-points (Figure 2c)

$$\lambda_{m}(\zeta) = \lambda_{m}(v),$$

$$\theta_{n}(\zeta) = \theta_{n}(u),$$

$$\xi(t) = \xi \left(\lambda_{m}(v), \theta_{n}(u), t\right)$$

$$(58)$$

For the first COTE (Eq. 29a) at u-points one has the following differential-difference replacements:

$$4\Delta\lambda U(t) \to U(t) - U(t) \\ \lambda, m, n \to U(t) - 2U(t) \\ \lambda\lambda, m, n \to W+\mu, n \to W+\mu$$

$$2\Delta\lambda\zeta(t;u) \to \zeta(t) - \zeta(t) \atop \lambda,m,n} \to \zeta(t) - \zeta(t) \atop m-\mu,n}.$$

For the second COTE (Eq. 29b) at v-points one has the analogous differential-difference substitutions

$$4\Delta\lambda^{2}V(t) \to V(t) = 2V(t) + V(t) \\ \lambda\lambda,m,n \to V(t) = -2V(t) + V(t) \\ \theta,m,n \to V(t) = -V(t) \\ m,n-1 \to V(t) = -2V(t) + V(t) \\ \theta\theta,m,n \to m,n+1 \to m,n-1,$$

$$4\Delta\theta^{2}V(t) \to V(t) = -2V(t) + V(t) \\ m,n \to m,n+1 \to m,n-1,$$

$$4U(t;v) \to [U(t) + U(t) + U(t) + U(t) \\ m,n \to m+\mu,n+1 \to m,n \to m,n+1,$$

$$4\Delta\lambda U(t;v) \to [U(t) + U(t) + U(t) + U(t) \\ \lambda,m,n \to m+\mu,n+1 \to m,n \to m,n+1,$$

$$4\Delta\theta U(t;v) \to [U(t) + U(t) + U(t) + U(t) + U(t) \\ \theta,m,n \to m+\mu,n+1 \to m,n \to m,n+1,$$

$$4\Delta\theta U(t;v) \to [U(t) + U(t) +$$

and

$$2\Delta\theta \xi(t;\mathbf{v}) \rightarrow \xi(t) - \xi(t)$$

 $\theta,m,n \rightarrow m,n+1$

The third COTE (Eq. 30) at ζ-points becomes directly

$$R \sin \theta_n(u) \, \xi(t) = \frac{1}{2\Delta \lambda} \left[U(t) - U(t) \atop m,n} + \frac{1}{2\Delta \theta} \left[\sin \theta_n(v) V(t) - \sin \theta(v) V(t) \atop n-1 \ m,n-1} \right] \, . \tag{61}$$

It may be noted that in Equations 59 and 60 V(t; u) at u-points and U(t; v) at v-points, are defined as linear averages of adjacent V(t) at v-points and U(t) at u-points, respectively. After defining obvious analogous substitutions for the depth function, $H(\lambda, \theta)$ (given at ξ -points, Sect. 5.B), and its first and second partial derivatives at u-and v-points, one arrives at the following spacially differenced equations:

$$\Delta t U_{t,m,n}^{(t)} = \alpha \Delta t \frac{GH(u)}{R \sin \theta_n(u)} \eta_{t,u}^{(t,u)} + A_{m,n}^{(3)} \left[\xi_{m-\mu,n}^{(t)} - \xi_{m,n}^{(t)} \right] - A_{m,n}^4 U_{t}^{(t)}$$

$$+ A_{m,n}^5 U_{m+\mu,n}^{(t)} + A_{m,n}^6 U_{m,n}^{(t)} + A_{m,n}^7 U_{m,n}^{(t)} + A_{m,n}^8 U_{m,n}^{(t)}$$

$$+ A_{m,n}^9 \left[(V(t) + V(t)) - (V(t) + V(t)) \right]$$

$$+ A_{m,n}^{(t)} \left[(V(t) + V(t)) + (V(t)) + V(t) \right]$$

$$+ A_{m,n}^{(t)} \left[(V(t) + V(t)) + (V(t)) + V(t) \right] , \qquad (62a)$$

$$\Delta t V(t) = \alpha \Delta t \frac{GH(v)}{R} \eta(t,v) + B_{m,n}^{3} \left[\xi(t) - \xi(t) \right] - B_{m,n}^{4} V(t)$$

$$+ B_{m,n}^{5} \left[V(t) + V(t) \right] + B_{m,n}^{6} V(t) + B_{m,n-m,n-1}^{7} V(t)$$

$$+ B_{m,n}^{8} \left[U(t) - U(t) \right] + B_{m,n}^{9} \left[U(t) - U(t) \right]$$

$$+ B_{m,n}^{10} \left[U(t) + U(t) + U(t) + U(t) \right]$$

$$+ B_{m,n}^{10} \left[U(t) + U(t) + U(t) + U(t) \right], \qquad (62b)$$

and

$$\Delta t \zeta(t) = C_n^1 \left[U(t) - U(t) \atop m, n} + C_n^2 V(t) - C_n^3 V(t) \atop m, n-1, \qquad (63)$$

where all A, B, C, and H's are defined by Equations 69, 70, and 71.

In the second part of the differencing process, a regular, finite-difference scheme (Sect. 2(g), Figure 3a) in time is defined by integrating Equations 62 and 63 over a single time step, Δt , using an average integration rule of the form $(U \to V \to \zeta)$

$$\int_{t_j}^{t_{j+1}} U(t)dt = \Delta t \left[\kappa U^{j+1} + (1-\kappa)U^j \right], \tag{64}$$

where

$$U^{j} = U(t_{j}),$$

 $t_{j} = (j-1)\Delta t, j = 1, 2, ...,$ (65)

and κ = some differencing parameters, which usually satisfy the restriction

$$0 \le \kappa \le 1. \tag{66}$$

As will be seen, the resulting discrete ocean-tide equations (DOTEs) are very sensitive to the choice of the differencing parameters, κ , and the time step, Δt . In fact, depending on the chosen values of κ and Δt , the DOTEs may be stable or unstable (Sect. 5.G) for any specified values of eddy viscosity, A, and bottom-friction coefficient, B. Moreover, different values of κ for Equations 62a, 62b, and 63 and for the various point values of U, V, and ζ can be chosen so that the resulting DOTEs may become explicit or implicit. Considering the complexity of the ocean basin and the large number of oceanic mesh cells, an explicit form for the DOTEs is here preferred.

After carrying out the integration of Equations 62 and 63 with the three (obvious) differencing parameters $\kappa = 0$, κ , and $\overline{\kappa}$, one arrives at the following explicit DOTEs:

$$[1 + \kappa A_{m,n}^{4}] U_{m,n}^{j+1} = A_{m,n}^{1} \sin \frac{\sigma \Delta t}{2} (2j-1) + A_{m,n}^{2} \cos \frac{\sigma \Delta t}{2} (2j-1)$$

$$+ A_{m,n}^{3} [\xi_{m-\mu,n}^{j} - \xi_{m,n}^{j}] + [1 - (1 - \kappa) A_{m,n}^{4}] U_{m,n}^{j}$$

$$+ A_{m,n}^{5} U_{m+\mu,n}^{j} + A_{m,n}^{6} U_{m-\mu,n}^{j} + A_{m,n}^{7} U_{m,n+1}^{j} + A_{m,n}^{8} U_{m,n-1}^{j}$$

$$+ A_{m,n}^{9} [(V_{m,n}^{j} + V_{m,n-1}^{j}) - (V_{m-\mu,n}^{j} + V_{m-\mu,n-1}^{j})]$$

$$+ A_{m,n}^{10} [(V_{m,n}^{j} + V_{m,n-1}^{j}) + (V_{m-\mu,n}^{j} + V_{m-\mu,n-1}^{j})] ,$$

$$[1 + \kappa B_{m,n}^{4}] V_{m,n}^{j+1} = B_{m,n}^{1} \cos \frac{\sigma \Delta t}{2} (2j-1) + B_{m,n}^{2} \sin \frac{\sigma \Delta t}{2} (2j-1)$$

$$+ B_{m,n}^{3} [\xi_{m,n+1}^{j} - \xi_{m,n}^{j}] + [1 - (1 - \kappa) B_{m,n}^{4}] V_{m,n}^{j}$$

$$+ B_{m,n}^{5} [V_{m+\mu,n}^{j} + V_{m-\mu,n}^{j}] + B_{m,n}^{6} V_{m,n+1}^{j} + B_{m,n}^{7} V_{m,n-1}^{j}$$

$$+ B_{m,n}^{8} [U_{m+\mu,n+1}^{j} - U_{m,n}^{j}] + B_{m,n}^{9} [U_{m+\mu,n}^{j} - U_{m,n+1}^{j}]$$

$$+ B_{m,n}^{10} [U_{m+\mu,n+1}^{j} + U_{m,n}^{j} + U_{m+\mu,n}^{j} + U_{m,n+1}^{j}] ,$$

$$(67b)$$

$$\xi_{m,n}^{j+1} = \xi_{m,n}^{j} + \overline{\kappa} \left[C_{n}^{1} \left(U_{m,n}^{j+1} - U_{m+\mu,n}^{j+1} \right) + C_{n}^{2} V_{m,n}^{j+1} - C_{n}^{3} V_{m,n-1}^{j+1} \right] \\
+ \left(1 - \overline{\kappa} \right) \left[C_{n}^{1} \left(U_{m,n}^{j} - U_{m+\mu,n}^{j} \right) + C_{n}^{2} V_{m,n}^{j} - C_{n}^{3} V_{m,n-1}^{j} \right].$$
(68)

The coefficients of the DOTEs (Eq's. 67 and 68) are

$$A_{m,n}^{1} = A_{m,n} \cos 2\nu (m - \mu) \Delta\theta,$$

$$A_{m,n}^{2} = A_{m,n} \sin 2\nu (n - \mu) \Delta\theta,$$

$$A_{m,n}^{3} = \beta \Delta t G H(u) \Gamma_{n}(u) / L,$$

$$A_{m,n}^{4} = 2a \frac{\Delta t}{L} \psi_{n}(u) \left[\omega_{n}(u) H(u) - \widetilde{H}(u) \right] + b \Delta t / \Gamma_{n}(u) H(u),$$

$$A_{m,n}^{5} = a \frac{\Delta t}{L} \psi_{n}(u) \Gamma_{n}^{2}(u) H(u) \left[1 + \overline{H}(u) \right],$$

$$A_{m,n}^{6} = a \frac{\Delta t}{L} \psi_{n}(u) \Gamma_{n}^{2}(u) H(u) \left[1 - \overline{H}(u) \right],$$

$$A_{m,n}^{7} = a \frac{\Delta t}{L} H(u) \left[\psi_{n}(u) + \frac{\mu \Delta \theta}{2} \left(2 + \Gamma_{n}(u) \right) \cos (2n - 1) \Delta \theta \right],$$

$$A_{m,n}^{8} = a \frac{\Delta t}{L} H(u) \left[\psi_{n}(u) - \frac{\mu \Delta \theta}{2} \left(2 + \Gamma_{n}(u) \right) \cos (2n - 1) \Delta \theta \right],$$

$$A_{m,n}^{9} = a \frac{\Delta t}{L} \Gamma_{n}(u) H(u) \left[\psi_{n}(u) \hat{H}(u) - \frac{\mu \Delta \theta}{2} \left(3 + 2\Gamma_{n}(u) \right) \cos (2n - 1) \Delta \theta \right],$$

$$A_{m,n}^{9} = a \frac{\Delta t}{L} \Gamma_{n}(u) H(u) \left[\psi_{n}(u) \hat{H}(u) - \frac{\mu \Delta \theta}{2} \left(3 + 2\Gamma_{n}(u) \right) \cos (2n - 1) \Delta \theta \right],$$

$$A_{m,n}^{10} = -a \frac{\Delta t}{L} \Gamma_{n}(u) H(u) \left[\psi_{n}(u) \hat{H}(u) - \frac{\mu \Delta \theta}{m} \cos (2n - 1) \Delta \theta \right],$$

and

$$\begin{split} A_{m,n}^{10} &= -a\frac{\Delta t}{L} \, \Gamma_n(u) \, H(u) \, \bar{H}(u) \, \left[\psi_n(u) \, \hat{H}(u) \, - \frac{\mu \Delta \theta}{2} \, \cos \left(2n \, - \, 1 \right) \Delta \theta \, \right] \\ &+ \frac{\Delta t}{2} \, \Omega \, \cos \left(2n \, - \, 1 \right) \Delta \theta \; , \end{split}$$

where

$$A_{m,n} = -4 \frac{\alpha G}{\sigma R} KH(u) \sin \frac{\sigma \Delta t}{2}$$

$$\begin{cases} \sin (2n-1) \Delta \theta, & \nu = 2 \\ \cos (2n-1) \Delta \theta, & \nu = 1 \\ 0, & \nu = 0 \end{cases}$$

$$\psi_n(u) = \frac{1}{2} (1 + 1/\Gamma_n(u)), \quad \omega_n(u) = 1 + \Gamma_n^2(u);$$
 (69a)

$$B_{m,n}^1 = B_{m,n} \cos \nu (2m - \mu) \Delta\theta,$$

$$B_{m,n}^2 = -B_{m,n} \sin \nu (2m - \mu) \Delta\theta,$$

$$B_{m,n}^3 = \beta \Delta t GH(\mathbf{v})/L,$$

$$B_{m,n}^4 = 2a \frac{\Delta t}{L} \psi_n(\mathbf{v}) [\omega_n(\mathbf{v}) H(\mathbf{v}) - \widetilde{H}(\mathbf{v})] + b \Delta t / \Gamma_n(\mathbf{v}) H(\mathbf{v}) ,$$

$$B_{m,n}^5 = a \frac{\Delta t}{L} \psi_n(\mathbf{v}) \Gamma_n^2(\mathbf{v}) H(\mathbf{v}) ,$$

$$B_{m,n}^6 = a \frac{\Delta t}{L} H(v) \left[\psi_n(v) (1 - \hat{H}(v)) + \frac{\mu \Delta \theta}{4} (3 + 2\Gamma_n(v)) \cos 2n \Delta \theta \right],$$

$$B_{m,n}^7 = a \frac{\Delta t}{L} H(v) \left[\psi_n(v) (1 + \hat{H}(v)) - \frac{\mu \Delta \theta}{4} (3 + 2\Gamma_n(v)) \cos 2n \Delta \theta \right],$$

(69b)

$$B_{m,n}^{8} = a \frac{\Delta t}{L} \Gamma_{n}(\mathbf{v}) \psi_{n}(\mathbf{v}) H(\mathbf{v}) \left\{ 2\mu \Delta \theta \Gamma_{n}(\mathbf{v}) \cos 2n \Delta \theta - \overline{H}(\mathbf{v}) \right\},$$

$$B_{m,n}^9 = a \frac{\Delta t}{L} \Gamma_n(\mathbf{v}) \psi_n(\mathbf{v}) H(\mathbf{v}) \left[2\mu \Delta \theta \Gamma_n(\mathbf{v}) \cos 2n \Delta \theta + \bar{H}(\mathbf{v}) \right],$$

and

$$B_{m,n}^{10} = -a \frac{\Delta t}{L} \Gamma_n(\mathbf{v}) \psi_n(\mathbf{v}) H(\mathbf{v}) \bar{H}(\mathbf{v}) \left[\mu \Delta \theta \Gamma_n(\mathbf{v}) \cos 2n \Delta \theta + \hat{H}(\mathbf{v}) \right]$$
$$- \frac{\Delta t}{2} \Omega \cos 2n \Delta \theta ,$$

where

$$B_{m,n} = -4 \frac{\alpha G}{\sigma R} KH(v) \sin \frac{\sigma \Delta t}{2} \begin{cases} 2 \sin 4n \Delta \theta, & v = 2 \\ \cos 4n \Delta \theta, & v = 1 \\ -3 \sin 4n \Delta \theta, & v = 0 \end{cases}$$

$$\psi_n(\mathbf{v}) = \frac{1}{2} (1 + 1/\Gamma_n(\mathbf{v})), \quad \omega_n(\mathbf{v}) = 1 + \Gamma_n^2(\mathbf{v});$$
 (69b)

$$C_n^{(1)} = \frac{\Delta t}{L} \Gamma_n(u) ,$$

$$C_n^{(2)} = \mu \frac{\Delta t}{L} \Gamma_n(u) \sin 2n\Delta\theta ,$$

and

$$C_n^3 = \mu \frac{\Delta t}{L} \Gamma_n(u) \sin 2(n-1) \Delta \theta$$
.

In these coefficients, the following depth functions were used at u-points:

$$H(u) = \frac{1}{2} [H_{m,n} + H_{m-\mu,n}],$$

$$\bar{H}(u) = \frac{1}{4H(u)} [H(u) - H(u)],$$

$$m,n = \frac{1}{4H(u)} [H(u) - H(u)],$$

$$m,n = \frac{1}{4H(u)} [H(u) - H(u)],$$

$$m,n = \frac{1}{m,n} [H(u) - 2\Gamma_n^2(u) [H^2(u) + (\mu\Delta\theta)^2]]$$

$$+ \mu\Delta\theta\psi_n(u) \cos(2n - 1)\Delta\theta [(3 + 2\Gamma_n(u)) \hat{H}(u) - \mu\Delta\theta\Gamma_n(u) \cos(2n - 1)\Delta\theta]$$

$$-\frac{1}{2} [\Gamma_n^2(u) (H(u) + H(u)) + (H(u) + H(u))];$$

$$-\frac{1}{2} [\Gamma_n^2(u) (H(u) + H(u)) + (H(u) + H(u))];$$

(70)

and at v-points

$$\begin{split} H_{(\mathbf{v})} &= \frac{1}{2} [H_{m,n} + H_{m,n+1}] \;, \\ \bar{H}_{(\mathbf{v})} &= \frac{1}{4H(\mathbf{v})} [H_{(\mathbf{v})} - H_{(\mathbf{v})}] \;, \\ \bar{H}_{(\mathbf{v})} &= \frac{1}{4H(\mathbf{v})} [H_{(\mathbf{v})} - H_{(\mathbf{v})}] \;, \\ \bar{H}_{(\mathbf{v})} &= \frac{1}{4H(\mathbf{v})} [H_{(\mathbf{v})} - H_{(\mathbf{v})}] \;, \\ \bar{H}_{m,n} &= H_{(\mathbf{v})} \left\{ \omega_n(\mathbf{v}) - 2 [\hat{H}_{m,n}^2(\mathbf{v}) + (\mu \Delta \theta)^2 \Gamma_n^2(\mathbf{v})] \right. \\ &+ \mu \Delta \theta \psi_n(\mathbf{v}) \hat{H}_{(\mathbf{v})} [3 + \Gamma_n(\mathbf{v})] \cos 2n \Delta \theta \right\} \\ &- \frac{1}{2} \left[\Gamma_n^2(\mathbf{v}) (H_{(\mathbf{v})} + H_{(\mathbf{v})}) + (H_{(\mathbf{v})} + H_{(\mathbf{v})}) \right] \;. \end{split}$$

It may be mentioned that for $\kappa = 0$ and $\overline{\kappa} = 1$, the finite-difference scheme coincides with the technique used, e.g., by Hansen (1966), Zahel (1970, 1973, 1975), and Estes (1975, 1977). Extensive exploratory computations were carried out by the author with numerous κ and $\overline{\kappa}$ values within the ranges $0 \le \kappa \le 1.5$ and $0.5 \le \overline{\kappa} \le 1$. The computations produced no drastic differences, provided the eddy and bottom-friction coefficients a and b and the time step Δt were suitably chosen within their respective stability constraints (Sect. 5.G). Finally, it was decided to use the values

$$\kappa = \overline{\kappa} = 1 \tag{72}$$

because they yielded a preferable stability and seemingly best results. Moreover, this deviation from the Hansen-Zahel method became most significant for the hydrodynamical interpolation of empirical tidal data described in Section 5.F. Indeed, this novel technique uses the special property that for $\kappa = 1$ the bottom-friction coefficient, b, which enters only in A^4 and B^4 (Eq's. 69), becomes essentially a scaling multiplier of $U_{m,n}^{j+1}$ and $V_{m,n}^{j+1}$ in Equations 67a and 67b. Thus, together with $\overline{\kappa} = 1$ in Equation 68, the bottom-friction coefficient can easily be adjusted locally to match more closely prescribed tidal data.

With κ and $\overline{\kappa}$ specified by Equation 72, the DOTEs (Eq's. 67 and 68) still contain the parameters a, b, and Δt , which remain at one's disposal within their respective stability ranges (Sect. 5.G). They will be utilized to achieve best results by trial-and-error computations. The DOTEs (Eq's. 67a, b, and 68) can be applied to all oceanic mesh cells $S_{m,n}$ with $m=\mu$, 2μ , ..., 360 and n=4, 5,..., 168 sweeping across the globe from n=4 to n=168. This procedure can be executed, provided suitable initial and lateral boundary data (Sect. 5.E) are prescribed. At colatitude line n=4, the numerical solution is matched to the second-order arctic solution (Sect. 4) by the linear interpolation ($m=\mu$, 2μ , ..., 360)

$$U_{m,3}^{j+1} = \frac{1}{3} \left[2U_{m,4}^{j+1} + U_{m,1}^{j+1} \right],$$

$$U_{m,2}^{j+1} = \frac{1}{3} \left[U_{m,4}^{j+1} + 2U_{m,1}^{j+1} \right],$$

$$(U \to V \to \xi)$$
(73)

where $U_{m,1}^{j+1}$, $V_{m,1}^{j+1}$, and $\zeta_{m,1}^{j+1}$ are computed by Equations 43, 44, or 45. For colatitude lines n=7,14,29,150, spacially corresponding data (U, V, ζ) on n=8,15,30,151 (see Eq's. 46, 47, and 48) are defined by linear interpolation. Vice versa for n=8,15,30,151, spacially corresponding values (U, V, ζ) on n=7,14,29,150 are also defined by linear interpolation.

E. LATERAL-BOUNDARY, INITIAL, AND FINAL DATA

In order to complete the ocean-tide model, the DOTEs (Eq's. 67 and 68) must be supplemented by suitable lateral-boundary and initial values. As was explained in Section 2(d), in turbulent-flow situations, the mathematical boundary conditions usually preferred are (a) no-flow across the ocean shorelines and (b) free-slip along the ocean shorelines. It is clearly at this point that the great attractiveness of Richardson's (1922) staggered, finite-difference scheme (Sect's. 2(g) and 5.D) manifests itself in the practical simplicity with which the no-flow and free-slip (or no-slip) boundary conditions can be worked into the model. In fact, if $S_{m,n}$ is an oceanic boundary cell, then, by

definition (Sect. 5.A and Figure 2c), the zig-zagging mathematical boundary lines follow only mesh lines and pass only through velocity (u and/or v) points. The no-flow condition (a) is implemented by declaring at those points

$$U_{m,n}^{j+1} = 0 \text{ and/or } V_{m,n}^{j+1} = 0,$$
 (74)

respectively. If, say, the v-point of $S_{m,n}$ is oceanic and the v-points of $S_{m-\mu,n}$ and/or $S_{m+\mu,n}$ are terrestrial, then the free-slip condition (b) is satisfied by reflectively setting*

$$V_{m-\mu,n}^{j+1} = +V_{m,n}^{j+1} \text{ and/or } V_{m+\mu,n}^{j+1} = +V_{m,n}^{j+1},$$
 (75)

respectively.

The mathematical boundary conditions (a) and (b) were applied by Hansen (1966), Zahel (1970, 1973, 1975), Estes (1975, 1977), and others. In the present model, these conditions are strictly enforced only at boundary points where no known empirical tidal data are available. However, at most oceanic boundary cells, empirical tidal constants are defined in Section 5.C. These data are worked into the present tide model by a unique hydrodynamical interpolation procedure (Sect. 5.F) that requires a controlled violation of the no-flow condition (a).

The construction of the M_2 tide was started at j=1 ($t_1=0$) with an ocean completely at rest; i.e., with the initial values ($m=\mu, 2\mu, \frac{1}{2}..., 360; n=1, 2, ..., 168$)

$$U_{m,n}^{1} = V_{m,n}^{1} = \zeta_{m,n}^{1} = 0. (76)$$

The computations were carried over a prescribed number of time steps, j = J (mostly a quarter period), and then printed for inspection of the results. With or without program and/or parameter changes, the computations were restarted using the latest or any earlier taped output instead of the initial values (Eq. 76). Occasionally, it was beneficial to speed up an unwanted slow decay of transient eigenmodes by "negatively" averaging the output data of half a period time difference; i.e., by setting

$$\overline{U}_{m,n}^{J} = \frac{1}{2} \left[U_{m,n}^{J} - U_{m,n}^{J-\overline{J}} \right], U \to V \to \zeta, \tag{77}$$

where $\sigma \Delta t \overline{J} = \pi$. This simple procedure diminishes all undesirable eigenmodes of lower frequency than the forced frequency, σ , and, similarly, most higher frequency modes. The negatively averaged data (Eq. 77) represent obviously improved initial data.

Two output tidal elevations 2J-time steps apart (mostly a quarter-period apart),

$$\zeta_{m,n}^{J} = \xi_{m,n} \cos \left[\sigma \Delta t J - \delta_{m,n}\right]$$
and
$$\zeta^{J-2\bar{J}} = \xi_{m,n} \cos \left[\sigma \Delta t \left(J - 2\bar{J}\right) \delta_{m,n}\right],$$

$$(78)$$

^{*}If the no-slip condition is imposed, then the + signs in Equations 75 must be changed to - signs.

were used to compute the tidal "amplitudes"

$$\xi_{m,n} = [\xi_1^2 + \xi_2^2]^{\frac{1}{2}} \tag{79}$$

and "phases"

$$\delta_{m,n} = \arctan y/x \ (= 0 \text{ for } x = y = 0),$$
 (80)

where

$$\xi_{1} = (\xi_{m,n}^{J-2\overline{J}} + \xi_{m,n}^{J})/2 \cos \sigma \Delta t \overline{J},$$

$$\xi_{2} = (\xi_{m,n}^{J-2\overline{J}} - \xi_{m,n}^{J})/2 \sin \sigma \Delta t \overline{J},$$
(81)

$$y = \xi_1 \sin \sigma \Delta t (J - \overline{J}) - \xi_2 \cos \sigma \Delta t (J - \overline{J}),$$

$$x = \xi_1 \cos \sigma \Delta t (J - \overline{J}) + \xi_2 \sin \sigma \Delta t (J - \overline{J}).$$
(82)

In the finished product, i.e., when all (unforced) transient eigenmodes have satisfactorily decayed, the amplitudes (Eq. 79) and phases (Eq. 80) become essentially independent of time and undergo no further variation of significance after continued computations.

In the final M₂-tide model presented in Part II, the convergence toward the steady state of the amplitudes and phases was found to be generally oscillating. So the integration process could safely be terminated when the amplitudes and phases over most ocean areas varied by less than 1 cm and 1°, respectively. This excellent convergence feature held true even at most coastal points. The convergence was slightly less complete only in some coastal cells where the tidal elevations are extremely large and vary rapidly from degree to degree. But even in those anomalous places (e.g., on the Patagonian Shelf; see the Tide Tables in Part II), the convergence error was well below the desired 10 cm specified in Section 1.

In order to follow the convergence of the computer program more closely, the squared tidal-amplitude sum

$$\xi_n^2 = \sum_{m=1}^{360} \xi_{m,n}^2 \tag{83}$$

was computed and printed for each fixed colatitude line $n = 4, 5, \ldots, 168$. To compute the amplitudes $\xi_{m, n}$ in Equation 83, Equations 79 and 81 were used for $2\overline{J} = 1$ and J = j + 1; i.e., for two consecutive time steps in connection with Equation 68. In this measure (Eq. 83), the convergence was carried to almost three significant figures for all n.

F. HYDRODYNAMICAL INTERPOLATION OF EMPIRICAL TIDE DATA

The discrete ocean-tide model developed so far is completely specified with the exception of the eddy and bottom-friction coefficients, a and b, and the time step, Δt (Eq's. 69 and 70). When those parameters are chosen uniformly for the entire ocean basin within certain stability limits (Sect. 5.G), then a unique solution can be computed, and any agreement or disagreement with observed tidal data must be accepted unless a compromising modification of the well-posed problem is feasible.

Preliminary computations carried out with this purely theoretical tide model produced results (Sect. 2(f)) that fulfilled the accuracy requirements posed in Section 1 over most ocean areas. Nevertheless, a brief glance at the many significant variations of ocean tides observed at continental and island stations (see, e.g., the underlined empirical M₂-tide data in Tables 5-8, and Part II) leaves one convinced that such local tidal features could not be accommodated by a mathematical model that ignores local boundary details.

In order to enhance the quality of the purely hydrodynamical tide model described above, it appeared desirable to seek a physically acceptable way to incorporate the empirical tidal data themselves into the program. In a simple exploratory experiment, the arranged empirical tidal amplitudes, $\widetilde{\xi}_{m,n}$, and phases, $\widetilde{\delta}_{m,n}$ (Sect. 5.C), were introduced into the computer program and used (when available) in place of Equation 68 by setting

$$\widetilde{\xi}_{m,n}^{j+1} = \widetilde{\xi}_{m,n} \cos(\sigma \Delta t j - \widetilde{\delta}_{m,n}). \tag{84}$$

The results represented a rather spectacular detailed improvement of the tidal model, very close to the final tide data presented in Tables 5-8 and Part II.

Unfortunately, the substitution of Equation 84 for the sensitive equation of conservation of mass (Eq. 68) constitutes a "continuity gap"

$$\Delta \zeta_{m,n}^{j+1} = \widetilde{\zeta}_{m,n}^{j+1} - \zeta_{m,n}^{j+1} , \qquad (85)$$

which remained bounded but grew unrealistically large in some areas after continued integration in time. Also, the velocity field near such observed points became unacceptably distorted.

Of course, the continuity gap (Eq. 85) can be attributed to the following major causes which are physically plausible and have been briefly mentioned before:

- (a) The bottom-friction coefficient, b (in A^4 and B^4 of Eq's. 69), which is most effective in boundary cells, depends on local shore features such as true cell size and bottom slope and roughness (Sect. 3.B(k)-(n)).
- (b) The boundary cells are idealized by definition of strictly mathematical boundaries (see Sect's. 2(f) and 5.A, and Figure 1).
 - (c) The depth data of boundary cells are subjectively defined and, hence, faulty (Sect's. 2(f) and 5.B).
- (d) The empirical tidal constants in Equation 84 are also faulty to some degree because of inaccurate measurements, harmonic analyses, and subjective selections and assignments to the centers of the boundary cells (Sect. 5.C).
- (e) The discrete ocean-tide model is certainly not an exact description of the true oceanic tide; e.g., at boundaries, nonlinear inertial terms assume significance.

Obviously, the last two (hopefully minor) faults can be reduced only through continued future observations and modeling. However, the first two faults, (a) and (b), can be weakened by "hydrodynamically interpolating" the empirical tidal elevations (Eq. 84) into the tidal model and narrowing the continuity gap (Eq. 85) to an acceptable level as follows:

- (\overline{a}) Adjusting the velocity field by a locally controlled implicit variation of the bottom-friction coefficient, b, in Equations 69.
- (b) Lifting the strict condition of no-flow across the mathematical ocean boundary and allowing for a monitored in- or out-flow by implicitly defining a more physical ocean boundary (Figure 1).

Since the finite-differencing parameters κ and $\overline{\kappa}$ in Equations 67 and 68 have been chosen by Equation 72, the bottom-friction coefficient, b, in A^4 and B^4 of Equations 69 can be considered "implicitly" varied by directly replacing the velocity components in Equation 68 as follows:

$$U_{m,n}^{j+1} \to U_{m,n}^{j+1} + \left| U_{m,n}^{j+1} \right| (wu_1 + \overline{w}\overline{u}_1),$$

$$U_{m+\mu,n}^{j+1} \to U_{m+\mu,n}^{j+1} + \left| U_{m+\mu,n}^{j+1} \right| (wu_2 + \overline{w}\overline{u}_2),$$

$$V_{m,n}^{j+1} \to V_{m,n}^{j+1} + \left| V_{m,n}^{j+1} \right| (wv_1 + \overline{w}\overline{v}_1),$$

$$(86)$$

and

$$V_{m,n-1}^{j+1} \to V_{m,n-1}^{j+1} + \left| V_{m,n-1}^{j+1} \right| (wv_2 + \overline{w}\overline{v}_2),$$

provided $\widetilde{\xi}_{m,n} \neq 0$; i.e., provided an empirical tidal amplitude is available for the considered mesh cell $S_{m,n}$. In Equations 86, the consistency and scale parameters (u, \overline{u}) and (v, \overline{v}) are defined by

$$\begin{cases} u_1 = 1, & \overline{u}_1 = 0 \\ u_1 = 0, & \overline{u}_1 = A_{m,n}^4 \end{cases} & \text{otherwise, but}$$

$$u_1 = 0, & \overline{u}_1 = 0 \\ u_1 = 0, & \overline{u}_1 = 0 \end{cases} & \text{if } \widetilde{\xi}_{m-\mu,n} \neq 0;$$

$$\begin{cases} u_2 = 1, & \overline{u}_2 = 0 \\ u_2 = 0, & \overline{u}_2 = A_{m+\mu,n}^4 \end{cases} & \text{otherwise;}$$

$$(87a)$$

$$\begin{cases} v_1 = 1, & \overline{v}_1 = 0 \\ v_1 = 0, & \overline{v}_1 = B^4 \\ m, n \end{cases} \quad \text{otherwise};$$
 (87c)

and

$$\begin{cases} v_2 = 1, & \overline{v}_2 = 0 \\ v_2 = 0, & \overline{v}_2 = B_{m,n-1}^4 \end{cases} & \text{for } \Delta \xi_{m,n}^{j+1} \cdot V_{m,n-1}^{j+1} > 0, \\ v_2 = 0, & \overline{v}_2 = B_{m,n-1}^4 \\ v_2 = 0, & \overline{v}_2 = 0 \end{cases} & \text{otherwise, but}$$

$$(87d)$$

The continuity gap (Eq. 85) will be narrowed when the "control parameters" w and \bar{w} are determined successively by

$$w = \begin{cases} \Delta \zeta_{m,n}^{j+1} / \zeta & \text{for } \zeta \neq 0, \\ 0 & \text{for } \zeta = 0 \end{cases}$$
(88a)

with the first "control limit"

$$|w| \le k_1 \tag{88b}$$

and

$$\overline{w} = \begin{cases} \left[\Delta \xi_{m,n}^{j+1} - w \xi \right] / \overline{\xi} & \text{for } \overline{\xi} \neq 0, \\ 0 & \text{for } \overline{\xi} = 0 \end{cases}$$
(89a)

with the second control limit

$$|\bar{w}| \leq k_2, \tag{89b}$$

where (see Eq. 68)

$$\zeta = C_n^1 \left[u_1 \left| U_{m,n}^{j+1} \right| + u_2 \left| U_{m+\mu,n}^{j+1} \right| \right] + v_1 C_n^2 \left| V_{m,n}^{j+1} \right| + v_2 C_n^3 \left| V_{m,n-1}^{j+1} \right|$$
and
$$\overline{\zeta} = C_n^1 \left[\overline{u}_1 \left| U_{m,n}^{j+1} \right| + \overline{u}_2 \left| U_{m+\mu,n}^{j+1} \right| \right] + \overline{v}_1 C_n^2 \left| V_{m,n}^{j+1} \right| + \overline{v}_1 C_n^3 \left| V_{m,n-1}^{j+1} \right|.$$

$$(90)$$

It is important to note that $u_i \cdot \overline{u}_i = 0$ and $v_i \cdot \overline{v}_i = 0$ for i = 1, 2. Accordingly, both control limits, k_1 and k_2 , which are at one's disposal, regulate the allowed decrease or, respectively, increase of the velocity components in Equations 86; i.e., the implicitly permitted corresponding increase or decrease of the local bottom-friction coefficients in Equations 67. Since the integration sweeps across the ocean from $m = \mu$ to 360 and n = 4 to 168, the special choice of $u_1 = \overline{u}_1 = 0$ and $v_2 = \overline{v}_2 = 0$ in Equations 87a and 87d excludes possible double adjustments of the velocity components. Also, if $u_1 \neq \overline{u}_1$ and/or $v_2 \neq \overline{v}_2$, backward adjustments of the tidal elevations via the corresponding Equation 68 must be made. This requires the replacements

$$\zeta_{m-\mu,n}^{j+1} \to \zeta_{m-\mu,n}^{j+1} - C_n^1 \left| U_{m,n}^{j+1} \right| (wu_1 + \overline{w}\overline{u}_1)$$
and
$$\zeta_{m,n-1}^{j+1} \to \zeta_{m,n-1}^{j+1} - C_n^2 \left| V_{m,n-1}^{j+1} \right| (wv_2 + \overline{w}\overline{v}_2).$$
(91)

Analogous substitutions in the forward directions of m and n follow automatically in the integration process.

The velocity replacements in Equations 86 may be illustrated by the example

$$U_{m,n}^{j+1} > 0, \quad U_{m+\mu,n}^{j+1} > 0, \quad V_{m,n}^{j+1} > 0, \quad V_{m,n-1}^{j+1} \ge 0,$$

$$\Delta \xi_{m,n}^{j+1} > 0, \quad \widetilde{\xi}_{m-\mu,n} = 0, \quad \widetilde{\xi}_{m,n-1} \neq 0. \tag{92}$$

One finds w > 0, $\overline{w} \ge 0$, and

$$U_{m,n}^{j+1} \to U_{m,n}^{j+1} (1 + \overline{w} A_{m,n}^4),$$

$$U_{m+\mu,n}^{j+1} \to U_{m+\mu,n}^{j+1} (1 - w),$$

$$V_{m,n}^{j+1} \to V_{m,n}^{j+1} (1 + \overline{w} B_{m,n}^4),$$

$$(93)$$

and

$$\zeta_{m-\mu,n}^{j+1} \to \zeta_{m-\mu,n}^{j+1} - C_n^1 U_{m,n}^{j+1} \, \overline{w} A_{m,n}^4$$
.

At this point, it must be mentioned that attempts were explored to lift the control limits prescribed by k_1 and k_2 in Equations 88b and 89b in an effort to close the continuity gap completely. However, since the bottom-friction coefficient, b, in A^4 and B^4 of Equations 69 is rather small (Eq. 4b), the control limits, k_1 and k_2 , had to be kept small to achieve best results. Computations conducted with large control limits k_1 (excessive bottom friction) seemed to close the continuity gap, but the tidal and velocity fields in the open oceans assumed unrealistically small values. Large control limits k_2 (insufficient bottom friction) produced strong instabilities as anticipated from the analysis in Section 5.G. To safely check the possible instability, the second control parameter \overline{w} (Eq's. 86, 87, and 93) was defined, respectively, in units of $\overline{u} = A^4$ and $\overline{v} = B^4$ in contrast to u = 1 and v = 1, used for the first control parameter w.

After some trial-and-error computations, the following control limits were chosen for the M2-tide model:

$$k_1 = .03, \quad k_2 = .06.$$
 (94)

These moderate values reflect the well-known fact that the magnitude of bottom friction has a strong effect on the motions considered. Indeed, with some minor improvements of the tidal field, significant improvements of the continuity gap, velocity field, and convergence of the integration were achieved. This procedure was applied to all oceanic cells with known empirical tide data (Eq. 84), provided these cells bordered terrestrial cells or contained small islands or other bottom irregularities. No meaningful reason was seen to apply the same bottom-friction adjustment procedure to distant offshore oceanic cells with available deep-sea tide measurements. A comparison of computed tide data with those empirical deep-sea values is given in Table 4.

In order to implement the second step (\overline{b}) of the hydrodynamical interpolation procedure, the following velocity replacements in oceanic mesh cells bordering terrestrial cells were defined:

$$U_{m,n}^{j+1} \to \widetilde{wu}_{1} U_{m+\mu,n}^{j+1} ,$$

$$U_{m+\mu,n}^{j+1} \to \widetilde{w} \widetilde{u}_{2} U_{m,n}^{j+1} ,$$

$$V_{m,n}^{j+1} \to \widetilde{w} \widetilde{v}_{1} V_{m,n-1}^{j+1}$$
(95)

and

$$V_{m,n-1}^{j+1} \to \widetilde{w} \, \widetilde{\mathbf{v}}_2 \, V_{m,n}^{j+1} \ , \label{eq:vj+1}$$

provided $\xi_{m,n} = 0$ in Equation 84. The parameters $(\widetilde{u} \widetilde{v})$ are mutually consistent by definition:

$$\widetilde{u}_1 = 1 \quad \text{if } U_{m,n}^{j+1} = 0, \quad \text{otherwise } \widetilde{u}_1 = 0,$$

$$\widetilde{u}_2 = 1 \quad \text{if } U_{m+\mu,n}^{j+1} = 0, \quad \text{otherwise } \widetilde{u}_2 = 0,$$

$$\widetilde{V}_1 = 1 \quad \text{if } V_{m,n}^{j+1} = 0, \quad \text{otherwise } \widetilde{v}_1 = 0,$$

$$(96)$$

and

$$\widetilde{\mathbf{V}}_2 = 1$$
 if $V_{m,n-1}^{j+1} = 0$, otherwise $\widetilde{\mathbf{v}}_2 = 0$.

The remaining continuity gap will be further narrowed when the control parameter \widetilde{w} is determined to be in agreement with Equations 68, 85, 88, 89, and 90 by

$$\widetilde{w} = \begin{cases} \left[\Delta \zeta_{m,n}^{j+1} - w \zeta - \overline{w} \overline{\zeta} \right] / \widetilde{\zeta} & \text{for } \widetilde{\zeta} \neq 0 \\ 0 & \text{for } \widetilde{\zeta} = 0 \end{cases}$$

$$(97a)$$

with the third control limit

$$|\widetilde{w}| \leq k_3$$
, (97b)

where

$$\widetilde{\zeta} = C_n^1 \left[\widetilde{u}_1 U_{m+\mu,n}^{j+1} - \widetilde{u}_2 U_{m,n}^{j+1} \right] + \widetilde{V}_1 C_n^2 V_{m,n-1}^{j+1} - \widetilde{V}_2 C_n^3 V_{m,n}^{j+1} . \tag{98}$$

Obviously, the substitutions (Eq's. 95) specify consistent in- or out-flows across the mathematical boundaries of oceanic coastal cells as illustrated in Figure 1 without explicitly fixing the physical boundary line. Again, no complete removal of the continuity gap was possible. The most satisfactory results for the M₂ tide (Part II) were achieved by setting the third control limit (Eq. 79b) at

$$k_3 = 0.5.$$
 (99)

While the improvement of the tidal field was again moderate, the remaining continuity gaps and near-shore velocity distortions assumed uniformly satisfactory levels. These remaining small shortcomings of the model can easily be attributed to the boundary inaccuracies (c), (d) and (e) listed above, but for which no simple remedies were found.

It may be emphasized that the rather significant change in the near-shore velocity field permitted by the inand out-flow specifications (Eq's. 95) affected the tidal field only in a minor fashion. This important phenomenon is in agreement with the well-known fact that the pressure distribution in a fluid motion is very insensitive to large but local velocity variations. For instance, it is perhaps the most important postulate in Prandtl's boundary-layer theory (see, e.g., Schlichting, 1968), and it is the basis of the hydrostatic-pressure assumption invoked here in Section 3.E for the present tidal model.

G. STABILITY ANALYSIS

A rigorous stability analysis of the homogeneous DOTEs (Eq's. 67 and 68) is, of course, not possible. However, under the assumption of constant coefficients A, B, and C, the simplified DOTEs possess Fourier-type eigensolutions (Eq's. 109) that permit a local stability analysis of the difference system. As is well known (see, e.g., Richtmyer, 1957), such a local stability analysis produces stability limits that are usually sufficient for computational purposes. Indeed, computer experiments showed that the stability limits so derived below were scrupulously binding for the success of the integration. The following analysis is an expanded version of the investigation presented by Zahel (1970).

In detail, the following simplifications may be assumed:

- (a) b = 0; i.e., no bottom friction
- (b) $\Omega = 0$; i.e., no Coriolis force
- (c) For an arbitrary but fixed mesh cell (see Eq's. 56, 57, and $69\overline{a}$, \overline{b}),

$$\Gamma = \Gamma_n(u) = \Gamma_n(v) = 1/\mu \sin \theta = \text{locally constant},$$

$$\psi = \psi_n(u) = \psi_n(v) = \frac{1}{2}(1 + 1/\Gamma),$$

$$\omega = \omega_n(u) = \omega_n(v) = 1 + \Gamma^2;$$

$$H = H(u) = H(v) = \text{locally constant},$$

$$m,n = m,n$$

$$H(u) = \hat{H}(u) = \hat{H}(u) = 0, (u \to v).$$

$$m,n = m,n = m,n$$

$$m,n = m,n$$

$$m,n = m,n$$

$$m,n = m,n$$
(100)

It may be mentioned that the assumptions (a) and (b) have been made in order to display more clearly the most important stability characteristics of the DOTEs that are due to eddy dissipation and to the differencing parameters, κ and $\overline{\kappa}$. It is relatively easy to show that the bottom friction is always stabilizing, while the Coriolis force (in the present differencing scheme) is slightly destabilizing.

For the following derivations, it is helpful to introduce some specified reference values θ_r and H_r (consistent

with Eq's. 100),
$$\Gamma_r$$
, ψ_r , and ω_r , and
$$a_r = \frac{1}{\psi_r} \left(\frac{G}{\omega_r H_r} \right)^{\frac{1}{2}}, \quad \Delta t_r = L/(\omega_r H_r G)^{\frac{1}{2}}. \tag{101}$$

Finally, the following relative quantities may be introduced:

$$\overline{\psi} = \psi/\psi_r, \quad \overline{\omega} = \omega/\omega_r, \quad h = H/H_r,$$

$$\tau = \Delta t/\Delta t_r, \quad \epsilon = a/a_r.$$
(102)

For example, in this notation the eddy viscosity, A (Eq. 6), assumes the form

$$A = \epsilon a_r \psi H L, \tag{103}$$

where ϵ is the "dimensionless eddy coefficient."

The M2-tide Tables 5-8 (and Part II) are computed with the reference values

$$\theta_r = 30^{\circ}, H_r = 7259.84 \text{ m},$$
 (104)

so that

$$\psi_r = 3/4,$$
 $a_r = 0.021\,919\,2\,\text{sec}^{-1},$
 $\omega_r = 5,$ $\Delta t_r = 186.309\,\text{sec},$

$$180\sigma\Delta t_r/\pi = 1.5^\circ,$$
 $J_p = 360^\circ/1.5^\circ = 240,$

where J_p is the number of time steps required to integrate through one tidal period with $\Delta t = \Delta t_r$. Due to the grading of the grid system (Eq. 49), one has almost everywhere $(n \neq 1, 2, 3, \text{ and } 166, 167, 168)$

$$1 \le \overline{\psi} \le \frac{4}{3}, \quad 1 \ge \overline{\omega} \ge \frac{2}{5}. \tag{106}$$

Similarly, due to the cutoff depth data (Eq's. 50, 51, 71), one has the limits

$$0.001\ 3 < h < 1. \tag{107}$$

Several other reference values within the stability limits have been explored. Particularly extensive computations were carried out with $\Delta t_r = 248.412$ sec (so that $J_p = 180$), but the above reference values appeared to yield the best results.

With the simplifications and notations above, the coefficients (Eq's. 69 and 70) of the homogeneous DOTEs (Eq's. 67 and 68) become

$$A^{1} = A^{2} = B^{1} = B^{2} = 0,$$

$$A^{3} = \Gamma B^{3} = \beta G H \Gamma \Delta t / L,$$

$$A^{4} = B^{4} = 2 h \epsilon \tau \overline{\psi} \overline{\omega},$$

$$A^{5} = A^{6} = B^{5} = B^{6} = h \epsilon \tau \overline{\psi} / \omega_{r},$$

$$A^{7} = A^{8} = B^{7} = \Gamma^{2} A^{5},$$

$$A^{9} = A^{10} = B^{8} = B^{9} = B^{10} = 0,$$

$$C^{1} = \Gamma \Delta t / L, \quad C^{2} = C^{3} = \Delta t / L.$$
(108)

The reduced DOTEs (with constant coefficients) yield the Fourier-type eigensolutions

$$U_{m,n}^{j} = U_{0} d^{j} e^{i \left[\gamma_{1} (2m - 2\mu) \Delta \lambda + \gamma_{2} (2n - 1) \Delta \theta \right]},$$

$$V_{m,n}^{j} = V_{0} d^{j} e^{i \left[\gamma_{1} (2m - \mu) \Delta \lambda + \gamma_{2} (2n - 1) \Delta \theta \right]},$$
(109)

and

$$\xi_{m,n}^{j} = \xi_0 d^{j} e^{i \left[\gamma_1 (2m - \mu) \Delta \lambda + \gamma_2 (2n - 1) \Delta \theta \right]}$$

with an arbitrary wave vector (γ_1, γ_2) and some nonzero amplitude vectors (U_0, V_0, ζ_0) , provided the eingenvalue d satisfies the cubic characteristic equation

$$\begin{pmatrix} (A_{11} - dA_{10}) & 0 & A_{13} \\ 0 & (A_{22} - dA_{10}) & A_{23} \\ (1 - \overline{\kappa} + \overline{\kappa}d)A_{31} & (1 - \overline{\kappa} + \overline{\kappa}d)A_{32} & (1-d) \end{pmatrix} = 0,$$
 (110)

where (after some algebra)

$$A_{10} = 1 + \kappa A^4 = 1 + 2\kappa he\tau \psi \bar{\omega},$$

$$A_{11} = A_{22} = A_{10} - 2A^4 s^2,$$
and
$$A_{13}A_{31} + A_{32}A_{23} = -4\beta h\bar{\omega}\tau^2 s^2$$
(111a)

with

$$0 \le s^2 \le \frac{\Gamma^2 \sin^2 \gamma_1 \Delta \theta + \sin^2 \gamma_2 \Delta \theta}{\Gamma^2 + 1} \le 1. \tag{111b}$$

The cubic characteristic in Equation 110 yields the three eigenvalues $(d_0, d = d_1, d_2)$

$$d_0 = 1 - 4h\epsilon\tau \overline{\psi} \overline{\omega} s^2 / A_{11} \tag{112a}$$

and

$$dA_{10} = A_{10} - 2h\tau \overline{\omega}s^2 A_{01} \pm 2i\tau s \left[h\overline{\omega}(\beta A_{10} - h\overline{\omega}s^2 A_{01}^2) \right]^{1/2}, \tag{112b}$$

where

$$A_{01} = \epsilon \overline{\psi} + \beta \overline{\kappa} \tau. \tag{112c}$$

The DOTEs will be stable, provided

$$|d_k| \le 1 \text{ for } k = 0, 1, 2.$$
 (113)

Under the strict inequality of Equation 113, the three eigenvalues d_0 , d_1 , and d_2 define three decaying eigenwaves represented by Equations 109. Since d_0 is real, the corresponding eigenwave is a standing wave with no phase shift if $d_0 \ge 0$ (see Eq. 119). The other two eigenvalues d_1 and d_2 define a pair of eigenwaves progressing in opposite directions with the same decay and dispersion rates, provided

$$\beta A_{10} \geqslant h \bar{\omega} s^2 A_{01}^2. \tag{114}$$

This condition holds true for all $0 \le s \le 1$ (Eq. 111b); i.e., for all wave vectors (γ_1, γ_2) if and only if

$$\beta A_{10} \ge h \overline{\omega} A_{01}^2. \tag{115}$$

If this condition fails, then there exist some short waves with large wave numbers, γ_1 and γ_2 , which become standing waves of different decay rates. However, all sufficiently long waves remain dispersively progressing and decaying at the same rates.

It seems physically plausible to treat all long and short waves equally and, hence, to impose the conditions of Equations 113 and 115 on the free parameters, ϵ , τ , κ , and $\overline{\kappa}$. Using Equations 112, 113, and 115, one finds

$$|d_1|^2 = |d_2|^2 = 1 - 4h\tau \overline{\omega} s^2 \left[\epsilon \overline{\psi} - (1 - \overline{\kappa}) \beta \tau \right] / A_{10}$$
 (116)

with

$$|d_1|^2 = |d_2|^2 = d_0 \text{ for } \overline{\kappa} = 1.$$
 (117)

For $\overline{\kappa} < 1$, the stability condition requires

$$\epsilon \overline{\psi} \geqslant (1 - \overline{\kappa})\beta \tau;$$
 (118)

i.e., a minimum of eddy viscosity is necessary for stability. However, for $\bar{\kappa} = 1$, no minimum eddy viscosity is required, which explains the choice made here (Eq. 72) and by Hansen (1966), Zahel (1970, 1973, 1975), and Estes (1975, 1977).

For the chosen value $\bar{\kappa} = 1$, the stability condition is satisfied for all s and all eigenvalues d_0 , d_1 , and d_2 when

$$4h\epsilon\tau\overline{\psi}\overline{\omega} \leqslant A_{10}; \tag{119}$$

i.e., when

$$\tau = \frac{\Delta t}{\Delta t_r} \le 1/2 (2 - \kappa) \epsilon h \bar{\psi} \bar{\omega} , \qquad (120)$$

provided (Eq. 115 in explicit form) also

$$\frac{\beta}{h\bar{\omega}} \geqslant \beta^2 \tau^2 + 2\beta \epsilon \tau \bar{\psi} (1 - \kappa) + \epsilon^2 \bar{\psi}^2. \tag{121}$$

The obviously increased stability limits imposed by both Equations 120 and 121 explain the choice of $\kappa = 1$ made here for the present tide model (Eq. 72) in deviation from the value $\kappa = 0$ used by Hansen, Zahel, and Estes. It may also be recalled that as an important by-product, $\kappa = 1$ facilitated the simple hydrodynamical interpolation of empirical tidal data into the model described in Section 5.F.

With $\beta = 0.90$ (Eq's. 17) and the possible extreme values of h = 1 and, simultaneously, $\overline{\psi} = \overline{\omega} = 1$ (Eq's. 106 and 107), one finds from Equations 120 and 121 for $\kappa = 1$ and $\tau = 1$ the allowable range for the dimensionless eddy coefficient, ϵ :

$$0 \le \epsilon \le 0.3 \ (\Delta t = \Delta t_{\star}). \tag{122}$$

The same range holds also for the southern three colatitude lines n = 166, 167, 168, which violate the condition of Equation 49, but for which the relative depth, h, falls sufficiently below unity. The upper limit on ϵ could be raised somewhat by considering the simultaneous values of ψ , $\overline{\omega}$, and h on each colatitude $n = 4, 5, \ldots, 168$ separately. In order to obtain the best possible tidal field, extensive trial-and-error computations led to the choice

$$\Delta t = \Delta t_{\rm p} = 186.309 \text{ sec and } \epsilon = 0.075.$$
 (123)

This final choice completes the detailed parameter specifications of the present tide model.

At this point, it may be noticed that the stability requirement for the DOTEs restricts the possible amount of eddy dissipation. As is physically plausible, the finite-differencing parameters, κ and $\overline{\kappa}$, the mesh size, L, the depth, H, the time step, Δt , and the dimensionless eddy coefficient, ϵ , are intimately related to each other. Trial-and-error computations are needed to select those parameters for best results. It is particularly important to observe that (especially for $\overline{\kappa} = 1$) the rate of decay (Eq's. 112a, 116, and 117) of all eigenwaves depends directly on the product $h\epsilon$. Accordingly, for fixed ϵ , waves in deep ($h\approx 1$) ocean basins decay faster than in shallow ($h \leq 1$) regions if bottom friction is negligible.

It is obviously this physically realistic phenomenon that led to the introduction of the novel depth-dependent eddy viscosity, A, defined by Equations 6 or 103. For a depth-independent eddy viscosity, one has $h\epsilon = \text{constant}$, in which case waves would decay at the same rate in deep or shallow (see Sect. 6.A) oceans, even though no bottom friction is present. Following Hansen (1966), Zahel (1970, 1973, 1975), and Estes (1975, 1977), the present tide model also used at first a constant eddy viscosity with rather disappointing results caused by the strongly varying bathymetry.

It is interesting to note that for the limiting case of Equation 115, i.e., for

$$\beta A_{10} = h\widetilde{\omega} A_{01}^2 \,, \tag{124}$$

Equation 112 assumes the simple form

$$dA_{01} = A_{01} - 2\beta\tau s^2 \pm 2i\beta\tau s (1 - s^2)^{1/2}.$$
 (125)

Hence, for the northeast waves under 45° with wave numbers (Eq. 111b)

$$\gamma_1 = \gamma_2 = \gamma, \ s^2 = \sin^2 \gamma \Delta \theta, \tag{126}$$

and $(\kappa = \overline{\kappa} = 1)$

$$d = \frac{\epsilon \overline{\psi} + \beta \tau e^{\pm 2i\gamma \Delta \theta}}{\epsilon \overline{\psi} + \beta \tau} = d_c + d_r e^{\pm 2i\gamma \Delta \theta}$$
(127)

Hence, in this case, the eigenvalues d_1 and d_2 lie on the circle

$$|d - d_c| = d_r, \quad d_c + d_r = 1$$
 (128)

as illustrated in Figure 5.

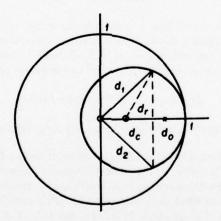


Figure 5. Illustration of Eigenvalues d_0 , d_1 , and $d_2 = \overline{d_1}$, in Circle $|d - d_c| \le d_r$

6. DISCUSSION OF THE TIDE MODEL

A. DISCRETE VERSUS CONTINUOUS OCEAN-TIDE EQUATIONS

In Section 2(g), it was contended that discrete ocean-tide equations (DOTEs, Eq's. 67 and 68) reflect the physical reality of ocean tidal currents more perfectly and perceptibly than the corresponding continuous equations (COTEs, Eq's. 29 and 30). Although the latter follow from the former by a "formal" limit process, it is neither technically feasible nor theoretically desirable to seek convergence of the discrete solution to the continuous integral. In fact, in fluid-flow problems of global dimensions, it is not possible to approximate the conditions of the continuous case to any reasonable degree. Even with future computer technology, it will not be meaningful to refine significantly, for instance, the 1° by 1° grid system defined in Sect. 5.A, which, with its 100-km mesh size, is far from being anywhere near a continuous description. Any attempt to refine the grid system would have to be matched by an improved bathymetry (Sect. 5.B), which requires worldwide in situ measurements.

From the theoretical point of view, it is equally superfluous to seek an e-approximation of the continuous situation, which in fact is only vaguely defined (Sect. 3.A). In laminar viscous-flow theory, such notions as particle (point) velocity and pressure, as well as the Navier-Stokes equations, are all derived from physically sound discrete (finite-difference) definitions by a formal limit procedure; i.e., by simply assuming the existence of the limit values (see, e.g., Schlichting, 1968; and Whitaker, 1968). While this assumption is well justified in most laminar-flow problems, it is well known today (see, e.g., Ladyzhenskaya, 1969) that, in general, even laminar motions must be sought in the class of generalized (distribution) functions. Hence, velocities, pressure, and their derivatives do not exist in the ordinary sense (pointwise); only their "functionals" (effects such as mass fluxes, forces, momenta) are physically defined.

The ambiguity of continuous-flow models becomes much more apparent in the critical laminar regime. Experiments (e.g., Busse and Whitehead, 1971) and theory (e.g., Schwiderski, 1972) have clearly established that, when given characteristic flow parameters (dimensions, velocities, etc.) exceed certain critical values, the corresponding, uniquely existent laminar motions become unstable and bifurcate into laminar flows of infinitely many different shapes. The classical laminar boundary and initial conditions (see, e.g., Sect's. 3.B and 5.E) are no longer sufficient to specify a unique motion. The situation seems to be governed by hysteresis and pure chance rather than rigorous physical selection principles.

Critical laminar motions become still more microscopically undefinable when the corresponding characteristic flow parameters (as the global dimensions of ocean currents) exceed further supercritical points and the motions go turbulent. The statistical approach underlying the "time-averaging" process to derive the so-called Navier-Stokes equations of mean turbulent flow (Eq's. 1 and 2) is entirely formal and vague (see, e.g., Schlichting, 1968). For example, what velocity (particle, point, etc.) is averaged over what time interval? In this respect, the present periodic tidal motions are clearly most illuminating of the problems at hand. If one averages (as usually meaningful) over a "sufficiently" long time span (say, longer than the tidal periods), then the averaged velocity should approach zero, which is obviously not of interest.

Evidently, turbulent particle velocities manifest themselves statistically through their integrated (macroscopic) physical effects, such as mass fluxes. Hence, the proper mathematical representation of turbulent motions should be sought in the class of generalized functions. Since the product of generalized functions has no mathematical meaning (see, e.g., Shilov, 1968), it appears understandable that there is no way to define the Reynolds stress tensor of turbulent motion (Sect. 3.C) by a meaningful ordinary or generalized function, because it contains quadratic products of the so-called fluctuating velocity residuals (see, e.g., Schlichting, 1968). However, its energy-dissipating (stress-like) effect is physically quite apparent and must be modeled in some macroscopic sense.

To avoid all conceptual difficulties of microscopic turbulent motions, it seems natural to fall back to the discrete (macroscopic) description of laminar flows that leads formally to the Navier-Stokes equations (see, e.g., Schlichting, 1968; and Whitaker, 1968). In fact, by proper "generalized" interpretation, virtually every notion used in the laminar regime retains its physical meaning in the discrete turbulent domain. For example, the "mean" x-velocity, u, of a "flow parcel" contained in a rectangular test (grid) cell of mesh lengths Δx , Δy , and Δz (Figure 6a) at some time, t, of a time interval, Δt , is defined as the mass flux (ΔM^x) crossing, say, the central ($x = x_1$) surface element (Δy , Δz) of the cell during the time span Δt divided by the fluid density, ρ , the area, $\Delta y \cdot \Delta z$, and the time, Δt , so that

 $u = \Delta M^{x} / \rho \Delta y \Delta x \Delta t. \tag{129}$

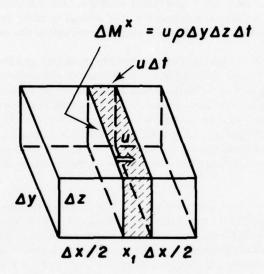


Figure 6a. Illustration of Average Velocity: ΔM^{x} = Mass Flux in x-Direction, u = Mean Velocity in x-Direction, and ρ = Fluid Density

This (generalized) definition of the mean-flow velocity, u, is uniformly valid for laminar and turbulent motions, since ΔM^x is a physically realistic (measurable) quantity for any $\rho > 0$, $\Delta x > 0$, $\Delta y > 0$, $\Delta z > 0$, and $\Delta t > 0$. If u exists in the limit $\Delta x \to 0$, $\Delta y \to 0$, $\Delta z \to 0$, and $\Delta t \to 0$, then u becomes the ordinary particle (or point) velocity. This assumption is well-justified in most (see above) laminar motions, where neighboring particles flow on smooth neighboring pathlines. In contrast to laminar flow, the limit assumption is certainly not justified in turbulent flow, where the fluid particles move in an indeterministically "turbulent," "fluctuating," or "eddying" way. In the sense of generalized functions, u could be defined in the limit by prescribing its mass flux (ΔM^x) for every nonvanishing ρ , Δx , Δy , Δz , and Δt .

Using similarly generalized definitions of all mean velocity components and pressure, and retaining their finite differences in place of (generally non-existent) ordinary derivatives, one arrives as usually at the discrete Navier-Stokes equations (analog of Eq's. 1 and 2) without requiring the existence of any limit values (see, e.g., Schlichting, 1968;

and Whitaker, 1968). They express simply the physical laws of conservation of momentum and mass for every test (mesh) cell. This is particularly tangible for the DOTE (Eq. 68), which conserves mass by balancing the excess of mass flux into a mesh cell with a corresponding increment in tidal height.

In the discrete form, the Navier-Stokes equations look formally the same for laminar and turbulent motions. Yet, when in the laminar case velocity and pressure exist pointwise in the limit, one has a unique microscopic (continuous) description of the flow subject only to specified boundary and initial conditions. Hence, the quality of a solution of the discrete model can be measured against the so-called exact integral of the continuous equations. In sharp contrast to the laminar case, no unique microscopic description of turbulent flow exists against which the quality of the discrete model could be measured. Therefore, the numerical analyst must seek an optimum size of the mesh cell $(\Delta x, \Delta y, \Delta z)$ and the time span (Δt) in order to model the undetermined microscopic turbulent motion so that their macroscopic effects match the expected or observed features.

The strong dependence of a discrete turbulent-flow model on the size of the mesh cell $(\Delta x, \Delta y, \Delta z)$ and the time span (Δt) can be assessed, for instance, from the definition of the average velocity, u, by Equation 129. With increasing mesh area $(\Delta y \cdot \Delta z)$, more and more as well as larger and larger fluctuating or eddying motions are filtered out and remain unaccounted for in the average value of u. Hence, the maximum mesh lengths Δx , Δy , and Δz must be sufficiently smaller than the smallest wave length one wishes to resolve. On the other side, if the area $(\Delta y \cdot \Delta z)$ in Equation 129 is chosen smaller and smaller, then u becomes more and more undetermined (fluctuating).

Similar arguments determine an optimum time step, Δt . In the present discrete tide model, the cell size was reasonably limited by the available bathymetric tables. The time step Δt (Eq. 123) was determined by trial-and-error computations so that 60 time points represent one-quarter of the M_2 -tide period.

Another significant distinction between the discrete Navier-Stokes equations of laminar and turbulent flow becomes apparent in the average-stress tensor. As is well known, the turbulent fluctuations neglected in the mean velocity and pressure manifest themselves as stress-like (energy-dissipating) forces that affect the mean motion. Unfortunately, no exact and unique constitutive equation is known today that relates those turbulent Reynolds stresses to the mean rate of strain (deformation of the flow parcel) determined by the average velocity (see, e.g., Schlichting, 1968; and Whitaker, 1968). In the Boussinesq (1877) substitution used in the present tide model, the mean turbulent-stress tensor is directly related to the average rate of strain (analog of Eq's. 5). Hence, the macroscopic stress effects of the turbulent fluctuations on the mean velocity are assumed to follow a similar simple law as the viscous laminar stresses in Newtonian fluids. Only the coefficient of viscosity is replaced by the so-called eddy viscosity which remains to be modeled to account for the otherwise neglected eddying motions in some best sense.

In the absence of better approximations, it seems idle to argue endlessly about the physical justification of the Boussinesq substitution; the fact remains that it represents the simplest possible constitutive equation, including zero used by some researchers. Moreover, it provides for considerable flexibility to model the microscopically undetermined but macroscopically apparent eddy dissipation by choosing suitable velocity-dependent or velocity-independent eddy viscosities either uniformly or separately for all three stress directions. Evidently, if the velocity field were known, a priori, then one could determine exact eddy viscosities, a posteriori, in many ways. In this connection, it is of interest to know that the mean flow is quite insensitive to fairly large variations (say 25 percent) of the eddy viscosity. This observation by Munk and Palmén (1951) was confirmed for oceanic tidal motions by the author's extensive computer experiments. It is probably related to the well-known fact that potential motions satisfy the complete Navier-Stokes equations of laminar flow with any constant viscosity. Above all, as with any other physical law, the Boussinesq substitution has successfully passed its crucial test in many practical applications in hydrodynamics, oceanography, and meteorology. The present ocean-tide model is no exception (see Sect. 6.B).

In order to illustrate the Boussinesq substitution in the discrete case, one may consider, for example, the average normal stress τ_1^{xx} produced by the filtered out (see the remarks to Eq. 129) fluctuating motions on the surface $(\Delta y, \Delta z)$ at $x = x_1$ shown in Figure 6b. Following Boussinesq, one has (see the analogous Eq. 5c)

$$\tau_1^{xx} = 2A\rho \frac{u_2 - u_0}{2\Delta x},\tag{130a}$$

where u_0 and u_2 are the corresponding mean velocities at $x = x_0 = x_1 - \Delta x$ and $x = x_2 = x_1 + \Delta x$. Hence, the turbulent stress τ_1^{xx} grows linearly with the rate of change of average velocity. Analogous to the corresponding laminar stress, this linear law appears physically acceptable, since the expected mean tidal velocities are small. One concludes from Equation 130a that a large change of mean flow produces a large turbulent stress which plausibly must be due to strong fluctuating motions.

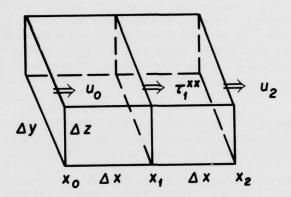


Figure 6b. Illustration of Mean Normal Stress: u_0 , u_2 = Average x-Velocities at $x = x_0$, x_2 , and τ_1^{xx} = Average Normal Stress

If one assumes a constant eddy viscosity, A, in Equation 130a, then the analogy between turbulent and laminar stress becomes complete. However, as was pointed out above, the strength of the fluctuating motions under consideration depends on the size of the surface area $(\Delta y \cdot \Delta z)$, which justifies the assumption

$$A = a\Delta y \Delta z \tag{130b}$$

in Equation 130a, where a may be held constant or used for further modeling. Similar arguments can be developed for all six turbulent stress components (see Eq's. 5) with eddy viscosities equivalent to Equation 130b.

The novel eddy-viscosity law expressed by Equation 130b is obviously equivalent to the eddy viscosity introduced in the present tide model by Equation 6a. It also explains Equation 4a, which specifies the bottom-friction coefficient, B, of the discrete tide model. The need for a mesh-area-dependent eddy viscosity became apparent when initial tidal computations with a constant value failed to yield realistic results (see Sect's. 2(a) and 5.G). It must be emphasized that the microscopically indeterministic nature of turbulent motions is not completely removed in the discrete flow model. Its specific macroscopic effects on the mean flow are apparent in the required optimum choice of the grid system, time step, differencing parameters, and, most of all, the eddy-viscosity law and scaling coefficient. No information on the fine structure of the turbulence can be expected from such a model.

B. QUALITY OF THE OCEAN-TIDE MODEL

Because of the absence of an exact (continuous) ocean-tide model, the degree of reality achieved by any approximate (discrete) model must be measured against empirically known features of ocean tides. In contrast to other turbulent-flow problems, for instance in general ocean currents, a large number of tidal observations (Sect. 5.C) around the world are available for comparison. Since the present tide model incorporates essentially all known empirical data by hydrodynamical interpolation (Sect. 5.F), no direct comparison of observed and computed data is feasible.

Nevertheless, a rather vivid and comprehensive appraisal of the reality of the present tide model is possible by inspecting the quality of hydrodynamical interpolation; i.e., by evaluating the "smoothness" with which the computed tide "accepts or rejects" (Figure 7) the empirical tidal data. In fact, the smoothness characteristics of the novel hydrodynamical interpolation technique are distinctly different from those of other direct interpolation procedures using power or trigonometric polynomials. In the latter case, smoothness of the interpolation can be carried up to any desired degree by simple intend. The adjustment of hydrodynamical parameters (Sect. 5.F) in the former method does not imply any smoothness of the interpolation, unless both the empirical input data and the hydrodynamical tide model are compatible with each other. As is well known (see Sect's. 5.B and 5.C), local tidal distortions, caused by an isolated roughness (seamount or small island) in the bottom relief, affect the surrounding ocean tide very little. The major level of ocean tides is shaped by continental shorelines and large (in area and/or length) islands and ridges. In contrast to ordinary polynomial interpolations, an important feature of the new hydrodynamical interpolation method is that it preserves those significant properties of ocean tidal currents without any essential alterations.

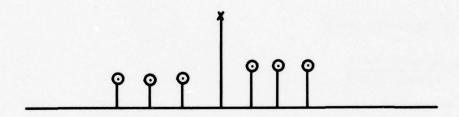


Figure 7. Illustration of "Rejected" (Protruding) Empirical Tidal Amplitude or Phase Value: x = Rejected Input Tide Values, and

• = Computed Tide Values

Extensive computer experiments were conducted to test the important smoothness characteristics of the hydrodynamical interpolation procedure. Faulty input data were deliberately inserted and quickly recognized as rejected by (protruding out of; Figure 7) the computed surrounding tide. Indeed, the first computations, which included empirical tidal data, revealed immediately several input errors in the data. Vice versa, smoothly accepted empirical tidal data were randomly deleted to test their backlash reaction on the computed tide. As anticipated, no significant modifications were detected. Consequently, the hydrodynamical interpolation technique permits a check of the reality of both the tide model and the empirical tidal input data. If an input value is rejected by the computed tide, then one or the other or both are defective. Fortunately, only very few discrepancies between the different sources of observed M₂-tide data (see Sect. 5.C) have been discovered that way.

The present discrete tide model has been applied to compute the global M2 ocean tide. A complete discussion and tabulation of all amplitudes and phases will be presented in Part II of this report. In order to display the quality of the tidal model, the computed amplitudes (in cm) and phases (in degrees) along with their adjacent empirical values have been tabulated in "30° by 50° map form" for four typical ocean areas (Tables 5-8). All empirically supported input data along continental shores and at island stations are underlined in the tables. All near-shore deep-sea measurements included in the model are labeled by subbrackets. As was explained in Section 5.F, all distant offshore deep-sea measurements are not included in the tide model. However, their approximate locations are marked by wavy underlines and their corresponding observed data are listed in Table 4. Land points are left blank.

In the evaluation of the tidal accuracy, one must remember that the ocean tide at any fixed location is determined by two harmonic constants. If (ξ_0, δ_0) and (ξ, δ) denote the respective local amplitudes and phases of the "true" and "computed" tides

$$\xi_0 = \xi_0 \cos(\sigma t - \delta_0), \quad \xi = \xi \cos(\sigma t - \delta), \tag{131}$$

then their time-dependent error is

$$\widetilde{\xi} = \xi_0 - \xi = \widetilde{\xi} \cos(\sigma t - \widetilde{\delta}) \tag{132a}$$

with the standard deviation

$$\operatorname{rms}(\widetilde{\xi}) = \frac{1}{2}\sqrt{2}\,\widetilde{\xi},\tag{132b}$$

$$\tilde{\xi}^2 = \xi_0^2 - 2\xi_0 \xi \cos(\delta_0 - \delta) + \xi^2$$
 (132c)

$$\tan \widetilde{\delta} = \frac{\xi_0 \sin \delta_0 - \xi \sin \delta}{\xi_0 \cos \delta_0 - \xi \cos \delta}.$$
 (132d)

Some maximum erros are

$$\widetilde{\xi}_{M} = \widetilde{\xi}_{M} = \xi_{0} + \xi \quad \text{for} \quad \delta_{0} - \delta = 180^{\circ}, \tag{133a}$$

$$\widetilde{\xi}_{M} = \widetilde{\xi}_{M} = \xi_{0} - \xi \quad \text{for} \quad \delta_{0} - \delta = 0^{\circ}, \tag{133b}$$

$$\widetilde{\xi}_{M} = \widetilde{\xi}_{M} = 2\xi \sin \frac{1}{2} (\delta_{0} - \delta) \quad \text{for} \quad \xi = \xi_{0}, \tag{134a}$$

$$\widetilde{\xi}_{M} = \widetilde{\xi}_{M} = \xi_{0} - \xi \quad \text{for} \quad \delta_{0} - \delta = 0^{\circ},$$
(133b)

$$\widetilde{\xi}_{L} = \widetilde{\xi}_{L} = 2\xi \sin \frac{1}{2} (\delta_{0} - \delta) \quad \text{for} \quad \xi = \xi_{0}, \tag{134a}$$

$$\widetilde{\xi}_{M} = \widetilde{\xi}_{M} = \xi$$
 for $\xi = \xi_{0}$ and $\delta_{0} - \delta = 60^{\circ}$. (134b)

Equation 134b expresses the important fact that a 60° phase error results in an amplitude error equal to the tidal amplitude and, hence, renders the computed tidal prediction completely useless. Of course, in regions of sufficiently small amplitudes, any phase error is acceptable.

Tables 5a and 5b depict the tidal amplitudes and phases, respectively, of the northwestern Atlantic Ocean including the eastern Caribbean Sea. As can be verified by earlier tide models, this entire area was very difficult to model, because its rough bottom topography has a strong effect on the tidal currents that sweep over or across various barriers with rapidly changing water levels. There is the broad and shallow continental shelf along the whole North American shoreline with Cape Hatteras, Long Island, Cape Cod, Nova Scotia, and Newfoundland all protruding into the ocean basin. Furthermore, there are the Grand Banks, the Bahama Banks, and the long and narrow

Caribbean Ridge. Obviously, all of the corresponding local tidal features could not be realistically captured by the tide model without a proper representation of the bathymetry (Sect. 5.B) and without the hydrodynamical interpolation (Sect. 5.F) of the locally collected tidal observations.

Now, if one scans the tidal amplitudes and phases (Tables 5a and 5b) from the north to the south, one gathers the vivid impression that the whole computed ocean tide is completely locked into the array of empirical (underlined) tidal data everywhere along the continental coast and along the many aligned islands separating the Atlantic Ocean from the Gulf of Mexico and the Caribbean Sea. It is particularly impressive to see the observed tide data at the offshore islands (Sable—SI, Barbados—BB, and even as far as Bermuda—BI) and at the included near-shore (subbrackets), deep-sea stations all realistically well-accepted by the computed surrounding tide. Moreover, one finds the excluded offshore deep-sea measurements (locations marked by wavy underlines) in the Atlantic and Caribbean Sea fully verified by the independent tide model.

As can be seen in the special listing of Table 4, the measured and computed amplitudes and phases at the Atlantic stations agree within 2 cm and 6°, respectively. The remaining discrepancy is probably within the experimental error due to short observation times and the use of the distant reference station Bermuda (Zettler et al., 1975), which exhibits even larger gaps between the various tidal observations listed in Table 3.

Attention may be drawn to the existence of considerable slopes between the empirical boundary data and the computed ocean-tide values in the high-amplitude ranges from Nova Scotia to Cape Cod and from Cape Hatteras to Florida's coast. Yet, these rapid tidal variations can be considered as realistic because throughout the same sections the empirical data, amongst themselves, display exactly the same roughness. This only substantiates clearly the fundamental difference between polynomial and hydrodynamical interpolation techniques pointed out above.

In Part II of this report, the same tidal roughness will be recognized in several similar coastal places around the world. From this typical phenomenon, one can draw the fortunate conclusion that, while some empirical data may be lacking high accuracy (see Table 2 and the British Admiralty Tide Tables, 1977), the computed adjacent ocean tide may retain its high quality.

In order to gain a deeper insight into the detailed tidal phenomena from the enclosed table charts (say, Tables 5a and 5b), it is helpful to recall the physical meaning of the tabulated tidal constants. The local tidal amplitude, ξ , is defined as half the tidal "range," which measu. s the total variation of the water level from high to low. Lines of constant amplitudes are called "corange lines." As can be seen from Equations 12 through 15, with $\lambda = 0$, the local phase, δ , specifies the tidal cresting time (in degrees) after the moon's (or sun's) passage over the Greenwich meridian ($\lambda \approx 0$). For the present M₂ tide (see Table 1), one has the following time conversions:

Lines of constant phases (simultaneous cresting times) are called "cotidal lines." In particular, at the $0^{\circ} = 360^{\circ}$ cotidal lines, which are conspicuously visible in the phase charts (Tables 5b to 8b), the tide crests simultaneously with the moon's passage over the Greenwich meridian. The tidal crest advances with time normal to the cotidal lines toward larger phases. A point of zero amplitude ($\xi = 0$) around which the tidal crest rotates from 0° to 360° is called an "amphidromic point;" it is marked in the tables by a circled star.

In the area of Tables 5a and 5b, a major amphidromic point is visible in the Caribbean Sea southeast of the island of Puerto Pico (PRI) near the marked deep-sea gauge station. The loosely connected Caribbean and Atlantic tides rotate counterclockwise around this point with the 0° = 360° cotidal line running northeastward. As a result

Table 5a. M₂ Tidal Amplitudes, ξ (cm), of the Northwestern Atlantic Ocean

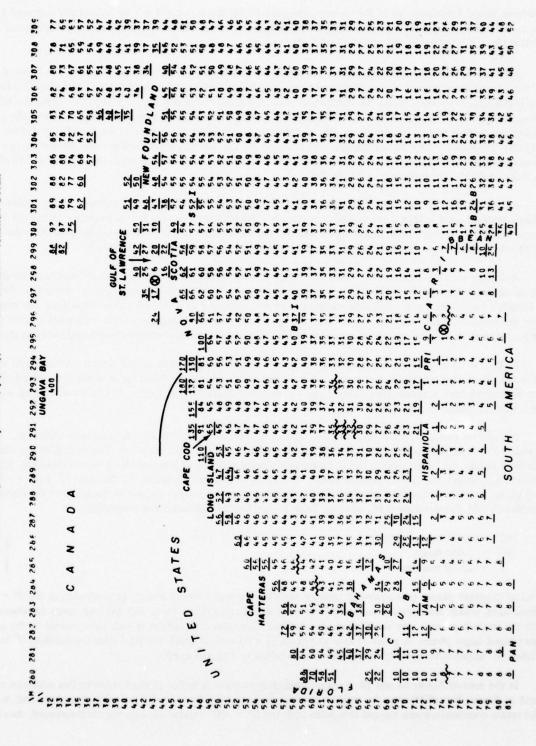


Table 5b. M₂ Tidal Phases, δ (°), of the Northwestern Atlantic Ocean

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of this rotation, the whole Caribbean Sea appears to be trapped and unable to develop any significant  $M_2$  tide. In agreement with the observations, the  $M_2$  tidal crest sweeps across the Caribbean Sea essentially from north to south with very little variation in water level.

If one follows the tidal crest around the amphidromic point from the Atlantic Ocean to the Caribbean Sea and back to the Atlantic, one recognizes a major tidal distortion caused by ocean ridges, which has long been discovered by practical tidalists (see, e.g., Harris, 1904; Bogdanov, 1961; Defant, 1961; and Luther and Wunsch, 1974). As the tide crosses the ridge between the islands, it suffers a distinct amplitude jump and a significant phase shift. For example, north of Puerto Rico (PRI) and Hispaniola and in the southeast around Barbados (BB), the computed and empirical Atlantic tide data display a higher water level and an earlier or, respectively, a delayed cresting time than the adjacent tide data on the Caribbean side. In particular, in full agreement with the observations, the tidal retardation time can easily exceed  $30^{\circ}$  ( $\approx 1$  hr). The distortion seems to depend on the angle with which the tidal crest spills over the ridge. Maximum distortion appears to be associated with a normal crossing. It may be pointed out that the realistic resolution of tidal distortions by ocean ridges (see below and Part II) constitutes probably the most significant improvement of the present model over all earlier hydrodynamical models.

The Atlantic portion of the Caribbean-Atlantic amphidromic rotation is opposed by a southward advancing tide from about Newfoundland in the north and by an eastward progressing tide from about Cape Cod to Cape Hatteras in the west. As a result of this interaction of three opposing tidal waves, the middle latitudes (around  $n = 60^{\circ}$ ) of the Atlantic display very few variations in tidal amplitudes and phases. In the high-amplitude sections between Nova Scotia and Cape Cod and between Cape Hatteras and Florida's coast, the Caribbean-Atlantic rotation system seems to be less affected by the opposing tidal waves. The extreme tidal amplitudes occurring at the former shore section are probably caused by the frontally advancing tidal crest splashing against the shallow coastal corner.

Although the computed tide in the Gulf of St. Lawrence displays the well-known amphidromic point (Defant, 1961), the grid system is much too crude to attach a high accuracy to the tidal constants in this border sea. For the same reason, the tidal data listed between Florida, Cuba, and the Bahamas are naturally less accurate than those in the open oceans.

Tables 6a and 6b illustrate the smoothness with which the computed tide of the northeastern Pacific Ocean attaches itself to the empirical tide data along the North American west coast. The tidal constants observed at the islands of Guadalupe (GI) and Farallon (FI), at the Cobb Seamount (CS), and at the included near-shore deep-sea stations fit realistically well into the computed surrounding tide. The amplitudes and phases of the excluded off-shore deep-sea measurements in the Pacific agree within 2 cm and 6°, respectively, with the computed data (Table 4), which is just the same accuracy as in the Atlantic.

Perhaps the most prominent feature of this area is the amphidromic point  $\bigotimes$ , around which the  $M_2$  tide rotates counterclockwise. This amphidromic system was predicted by Munk et al. (1970) and Irish et al. (1971) in almost identical geographical position. Earlier hydrodynamical tide models failed to resolve this system on proper location, although several models matched the empirical data along the coast quite well. Since the northeastern Pacific falls short in major bottom and coastal irregularities when compared to the northwestern Atlantic, the indicated rapid loss of quality in westerly direction seemed disappointing. Yet, as will be demonstrated below, this shortcoming could have been concluded from the obvious failure of those models to reasonably reproduce the tide over most of the north and central Pacific Ocean.

As was mentioned before, the author's preliminary tide model (Schwiderski, 1976) used a bathymetry that failed to represent the hydrodynamical barrier effects of the Marianas, Nampo, Kuril, Aleutian, and Hawaiian Ridges, as well as of other seamount chains. Consequently, the M₂ tide of almost the whole central, western, and northern Pacific area was modeled as a single huge amphidromic system, as pictured by the similar maps of other

numerical tidalists such as Zahel (1971) and Estes (1975, 1977). The clockwise-rotating Pacific tide was free to sweep undisturbed into the Philippine, Okhotsk, and Bering Seas. By the time the computed tidal crest reached the Aleutian Islands, it was just about 180° out of phase. When the original bathymetry was replaced by hydrodynamically defined depth data (Sect. 5.B), the entire Pacific Ocean resembled a whirlpool after some continued computations over several quarter periods. The amphidromic system weakened and its center slipped slowly southward, but drastically improved phases appeared gradually along the Aleutian Ridge confirming the anticipated effect of ocean ridges.

The complete turnaround of the Pacific M₂ tide near the Aleutian Islands was speeded up when the empirical tidal constants were introduced into the model. In fact, a repeat of the same computations settled the Pacific Ocean tide into its final position in a rather dramatic fashion. Striking improvements were registered over the whole Pacific and, of course, also over the Atlantic and Indian Oceans.

As is depicted in Tables 7a and 7b for the north-central Pacific, the amphidromic system is replaced by a low-amplitude tide. It appears to be locked in between the Aleutian and Hawaiian Ridges in the north and south and also between the Emperor Seamount chain in the west and the high-amplitude tide in the east, which progresses in a westerly direction from the west coast of North America (Tables 6a and 6b). The amplitude topography of this area resembles the low-amplitude tide in the Caribbean Sea (Table 5a). When the westward-advancing tidal wave enters the region between the Aleutian and Hawaiian Ridges, it suffers a remarkable, almost symmetric retardation at both ridges. In fact, as the visible  $(0^{\circ} = 360^{\circ})$  cotidal line in Table 7b reveals, the crest front of the tidal wave assumes the shape of an almost symmetric wedge. If one traces the  $0^{\circ}$  phase line westward beginning at both ridges, one can infer a definite idea about the realistic reproduction of the tide in this region. At both ends, the  $0^{\circ}$  phase is in full agreement with the empirical data. As the observed phases grow westward along both ridges, so grow proportionally the distances of the  $0^{\circ}$  phase line from the ridges.

The present computed M₂-tide model indicates no longer any symptoms of the original phase problems (see above) at the Aleutian and Hawaiian Ridges. The computed amplitudes and phases approach the empirical tidal constants from both sides of the ridges as smoothly as could be desired. As the tidal wave spills over both ridges in northwestward or southwestward directions, respectively, it suffers a tidal distortion similar to that found before at the Caribbean Ridge. Amplitude jumps and major phase shifts are again in complete agreement with observations (see the remarks of Luther and Wunsch, 1974). It is particularly gratifying to find the phase shift well developed along the whole length of the Hawaiian Ridge from the Island of Hawaii to Midway, even though only few empirical data were used at both ends. Also, it may be noticed that the observed tidal constants at the distant and isolated island stations of Pribilof (PF), Midway (MW), and Johnston (JI) are all realistically well integrated by the surrounding computed tide.

Ironically, the old and new M₂-tide maps constructed by Bogdanov (1961) and Luther and Wunsch (1974) by pure intuition and simple rules of thumb from empirical data came closest to the present charts. Indeed, both maps display no amphidromic system in the north-central Pacific. As will be verified in Part II, the computed amphidromic points between the Cook and Society Islands and near the southern edge of the Solomon Islands are both in almost identical positions with those charted by the same authors. Nevertheless, their detailed distribution of amplitudes and phases is still significantly different from the present one.

Perhaps the most spectacular display of the high quality of both the computed and the observed tidal data is brought out by Tables 8a and 8b depicting the high-amplitude tide of the central Pacific. Indeed, unlike any other open ocean area, the tabulated region is dotted with numerous tide gauge stations at island groups and at scattered isolated islands. In addition to the fully listed island chains, there are the isolated islands: Johnston (JI), Wake (WI), Kusaie (KI), Ocean (OI), Funafuti (FI), Wallis (IW), Niue (NI), and Norfolk (NF). The corresponding observed

Table 6a. M₂ Tidal Amplitudes, \(\xi\) (cm), of the Northeastern Pacific Ocean

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Table 6b. M2 Tidal Phases, 8 (°), of the Northeastern Pacific Ocean

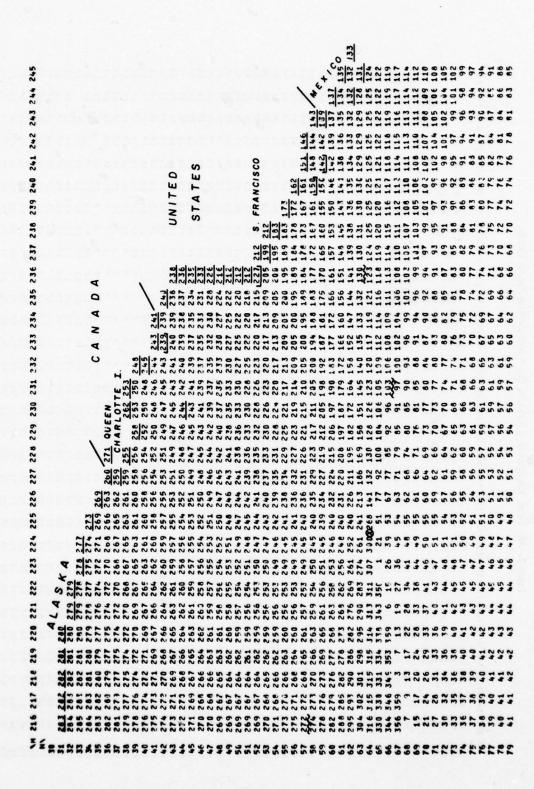


Table 7a.  $M_2$  Tidal Amplitudes,  $\xi$  (cm), of the North-Central Pacific Ocean

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Table 7b.  $M_2$  Tidal Phases,  $\delta$  (  $^{\circ}$  ), of the North-Central Pacific Ocean

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13 1		40495
2 18	11000000000000000000000000000000000000	12112
1 18		22222
1.9		22222
180	44440000000000000000000000000000000000	123 123 124 125 126 127
179	0 10 10 10 10 10 10 10 10 10 10 10 10 10	123 123 124 125 126 126
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111		
3	. W. W. C.	
	миничерововоровия и правительной правительной предостивной правительной предостивной предостивной предостивной п	

Table 8a. M2 Tidal Amplitudes, § (cm), of the Central Pacific Ocean

					H																																													
191		11	-	21	121	24	25	26	27	2.8	50		31	32	7 2		34	35	35	35	3.	36	35	36	35	35	35	32	35	35	35	34	36	34	3 4	33	72	35	35	32	36	3.8	39	36	5	,	41	25	23	42
161		11	20	25	25	52	56	22	28	30	11	32	*	72	2	35	36	36	33	37	-	3.8	3 6	3.8	37	37	37	37	37	37	37	36	36	36	36	2	4.5	37	**	37	2	49	39	0,	4.1	42	24		<b>t</b> 3	43
641		18	20	25	23	3	2	28	30	=	32	33	34	35	3	2	37		6	30	0,	07	07	75	5	6	39	39	39	39	46	39	39	3.5	0	47	2	38	38 N		39	39	04	4.1	24	27	4.3	43	11	77
8 8		1 9	21	23	52	26	20	50	31	32	13	15	36	11			19	0 7	9	1,1	1,1	42×	421	422	42 L	430	¥ 0 4	419	41	4.1	41	41	1,1	27	1	× 8		8 2	3.8	62	61	0 7	41	24	24	6.3	11	*	5.5	t t
181	-	20	22	54	26	58	52	11	42	33	15	36	37	3.8	5	67	41	3	42	43	4.3	.,	4.3	77	13	43	14	14	.,	£ 3	43	. 4	4.3	5	5	454	32	35		39	52	64	1,	42	1 1	**	77	4.5	45	46
9 6		21	23	56	22	58	30	35	33	15	36	37	39	39	10	;	42		1 1	11	1 1	45	45	45	64	45	6.5	45	54	45	54	45	45	45	3 .	3 3		43	77	4.3	20	0 4	41	24	43	**	4.5	94	46	46
185		22	54	27	58	30	35	33	35	46	3.8	39	0 7	41	62	43	1,1	45	45	9 7	94	97	47	15	47	14	94	46	99	14	11	14	47	14	6	60	45	45	4.5	₹94	994	258	110	123	1 1	5 7	94	44	14	1.7
184		53	52	28	30	11	3.4	35	36	3.8	30	0 %	41	42	43	11	54	94	4.7	47	14	8 7	84	24	8 7	9 5	0 5	84	6.8	8 1	84	0 1	64	5.0	25		99	44	94	14	14	44	45	57	54	94	4	14	47	4 6
183		72	56	28	31	33	7.	36	18	39	04	1,1	24	11	4	46	47	85	84	64	64	0 7	64	64	5 7	64	64	64	64	64	0.	20	20	51	7 15	200	8 7	14	84	6 4	8 7	47	17	47	4 8	8 7	6 4	4	2	6 4
		5	1	53	11	13	45	37	30	6.0	;	. 4	3 3	45	4	14	4.3	64	64	0 5	20	20	60	20	20	20	20	20	20	6.0	20	51	21	25	200	2.5	61	8 5	64	64	64	64	61	61	20	20	20	20	20	20
181		52	2	30	32	34	4	38	9	,	24	1 4	45	99	6.3	0,	64	20	51	51	51	25	25	51	25	25	25	25	51	51	51	25	24	25	*	22	51	14		6 5	20	20	20	51	25	25	22	25	51	51
180		92	2 8	31	33	35	3.7	39	1,	25	43	65	94	84	64	20	51	51	25	25	25	25	24	53	53	53	53	53	23	56.1	51	25	23	24	22	28	18	25	20	64	20	51	51	25	25	2 3	53	53	53	53
179		2	2	32	34	36	3.6	40	24	,	4,5	99	47	84	50	51	25	25	23	25	23	53	53	25	24	53	24	24	53	52 F	25	23	24	52		56	1	-	56	64	20	51	51	23	25	24	53	24	24	25
178	:		20	23	3.5	4.5	36	7	43	2 2	4	147	8 3	64	20	51	25	53	24	36	24	2¢	24	24	24	24	24	24	24	۲.	2 4	24	22	610	10	3 4	55	4	65	5	20	20	51	51	25	25	2 .	* 1	50	26
111		3	31	•	2	3.8	3	24	<b>7</b>	42	4		64	20	51	25	53	24	55	55	56	155	956	55	2 4	24	24	24	24	24	24	24	1	24		55	26	19	25	51	20	20	20	21	51	25	٠.	20	52	22
176		2	25	36	37	36	;	42	3 3	46	2	8 5	64	20	51	25	53	24	25	ď	57	3.5	52 0	53	53	24	24	24	24	24	24	24	24	53	20	24	24	54	25	51	20	0 1	6	20	20	21	25	26	2,	26
175	,	000	2	7	2	3.5	;	4.3	3	46	47	8 5	20	51	25	53	24	55	26	26	26	9 5 8 E	23	2	24	24	24	24	24	24	75	24	53	2 4	2	230	53	25	20	20	0	5	25	6 7	20	05	25	55	22	21
174	30	2 .	2 .	60	25	3.5	3	7	3	46	4	64	20	51	25	5 3	23	56	25	26	27	28	24	24	24	24	24	24	23	23	53	23	5	25	20	25	51	23	6	10		,	4	3	50	20	51	55	5,5	21
173		,	2 6	5	25	65	;	5 4	3 3	9	3	64	20	51	55	53	52	25	550	2	26	25	65	24	24	24	23	53	23	2	25	25	25	1,1		25	20	64	5	47	3	9	9	9	25	5 1	200	1,5	5.6	20
112		2 .		22	5	3.6	5	2	2	10	64	64	51	95	57	51	20	51	53	24	25	25	52	24	24	24	23	25	25	25	21	21	21	200	0 1	6	64	15	99	9	5 .	3	3 1	5.	4	3	3 0	2 0	25	20
171	3		,		9	9	,	2	3	9	0 4	9	64	667	20	20	20	51	23	53	24	24	124	24	24	53	25	51	51	51	20	20	5 .	2 4	4	9	6	1 5	1	43	? :	,	7.	3 .	3 .		3	2 0	25	22
170	2	::	2		6	20			,	9	3	9	4.8	784	8	14	•	20	23	53	24	20	53	23	23	23	21	20	20	6	6 4	3		0 4	14	. 5	6 5	5	0 1	2	1		*	3.	9 :	3 .	9 2	2 0	0 11	22
169	3.0			200	6	5		7	9	42	61	47	147	64	94	45	4.0	20	21	51	25	25	536	53	25	21	20	64	6	9	*	9 1		1 1	4.5	2	45	30	2	34	200	00	55	5	;	3 .		200	20	22
158	7 20		2 6	0	2 .	200	0 .	7	3	20	9	47	12	4.5	45	46	3	5 3	20	20	51	21	25	25	51	20	40		3	. 42		3	,	3 6	10	الم	8	7	25	34	35	00	200	20	0 0	3 .	7 3	2 0	576	1
167	200	10			2	*	0		,	1	42	68	3775	2 2	45	46	14	9	6 5	63	20	51	51	51	20	20	4	3	4	3	7 .	3	4	1 2	328	30.4	34	50	-	33	22	1	1	3		3 4		, r	5 2 N	76
16€	261				;	2			5	3 0	24	42	643	43	2 2	45	46	13	4	48	5	20	20	20	4 3	1 0	4	40	3 3	2,5	5	3	2 .	29	27	26	56	28	20	33	2	0.0	3	3 5	7 4	<b>7</b> 4	) e	2 0	2,5	31
16 5	25	, ,		9 6	,	,			,	5	*	0,	145	42	£ 3	5 5 I	45	9		4	9	4	5 3	5 4	5	5	9		3	3;	2	4 .	100	2 2	24	23	54	50		12	36,	4:		;	7 4		3 3		2 5	
164	24	30	22			2 6		2 .	55	5		39	0 4	4.1	42	× 52	3	42	4	94	2 5	3	3	4	3	*	9		5	40	200	35	000	23	21	20	21	53	87	C		0 .	3	7 .	3 4		ט פ	2 2	2 2	2
163	2.2	36	36	3 6	,	0 0		25	*	4:	2	38	38	39	9	;	45	43	3	42	42	9	9	40	9	42	42	:	3	39	4	213		19	11	11	18	20	3		22		,	÷ ;		, 0	, ,	23	55	
162	22	2	36	30		,,	0 0	2 .	7	25	*	36	37	37	38	39	0	;		5	3	5	42	5.		3	3 !	2	5	3	070	SNO	1:	1 1	14	14	16	13	22	17	36	0 .		* 4	6 9	, ,	21	2,0	26	2
* 2	7.1	12		3.5		22		:			9		28	83	40	85	96	18		58	06	1	26	56	*	35	36	,	96	66	2 00	100		70	0.5	90	20	90	5	01	::	21	21	**	12	1			20.	
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Table 8b. M2 Tidal Phases,  $\delta$  (  $^{\circ}$  ), of the Central Pacific Ocean

161	96
1 90	00000000000000000000000000000000000000
189	11111111111111111111111111111111111111
1 86	11111111111111111111111111111111111111
167	00000000000000000000000000000000000000
186	1000
185	20000000000000000000000000000000000000
184	11117 11117 11117 11117 1117 1117 1117
183	11111111111111111111111111111111111111
182	\$\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\
191	13 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1
1 80	11111111111111111111111111111111111111
179	11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 11120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120
178	20000000000000000000000000000000000000
111	00000000000000000000000000000000000000
92	2 4 0 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5
175	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
174	
173	
172	00000000000000000000000000000000000000
171	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
170	
169	11120 11120 11120 11120 11120 11120 11120 11120 11120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 1120 120
168	600007446000000000000000000000000000000
167	10000000000000000000000000000000000000
166	9 110 1115   110 1115   110 1115   110 1115   110 1115   110 1115   110 1115   110 1115   110 1115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 115   110 1
165	10000000000000000000000000000000000000
164	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
163	5 2 5 5 6 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7
291	000
3:	. 172 173 173 173 173 173 173 173 173 173 173

tidal constants listed in nongeographical arrangement appeared incoherent and, hence, uncorrelated, giving rise to doubt their true value. Yet, the computed tidal wave sweeps across the whole area in a southwesterly direction with little variation of its high amplitude. As the wave crest passes through the many checkpoints, it integrates and correlates without a single exception all the empirical data into one coherent unity.

### C. CONCLUSIONS

The hydrodynamical modeling of global ocean tides, described and tested above, shows that the original question posed by contemporary researchers (Section 1) can be successfully answered. In fact, one infers from the evaluation of the constructed  $M_2$  tide (Sect. 6.B) that it is now possible to predict the  $M_2$ -tide elevation of the ocean surface over the geoid anywhere in the open oceans with an accuracy of better than 5 cm. This accuracy goes considerably beyond the originally desired error bounds. It leaves ample room for superposable errors due to the additional tidal constituents listed in Table 1, which must be constructed with equivalent relative accuracy. The computation of the leading three components,  $S_2$ ,  $K_1$ , and  $O_1$ , is presently in progress.

Naturally, the achieved high accuracy of the M₂ tide in the open oceans drops off somewhat near continental or island stations where empirical data are missing or are less accurate themselves (see the introduction to the British Admiralty Tide Tables, 1977). Also, less accurate predictions must be anticipated in small border seas, bays, estuaries, and channels where the 1° by 1° grid system precludes a sufficient resolution. To improve the present tide model in those areas, significantly improved observations will be needed along with a locally refined network and corresponding bathymetric data.

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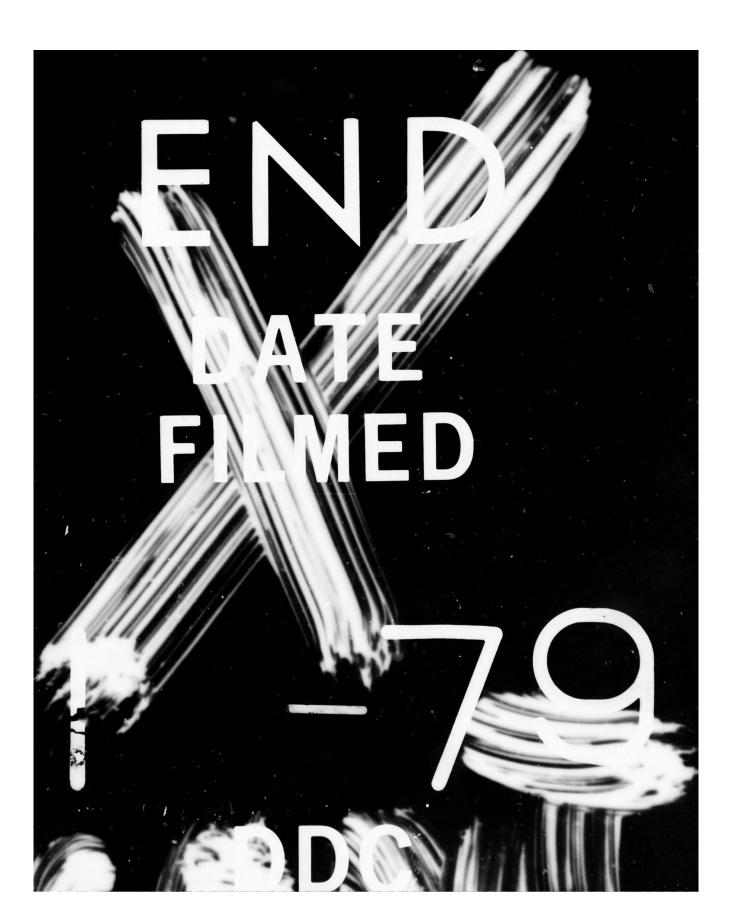
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NAVAL SURFACE WEAPONS CENTER DAHLGREN LAB VA GLOBAL OCEAN TIDES. PART I. A DETAILED HYDRODYNAMICAL INTERPOLA-ETC(U) AD-A060 913 SEP 78 E W SCHWIDERSKI NSWC/DL-TR-3866 UNCLASSIFIED GIDEP-E125-1971 NL







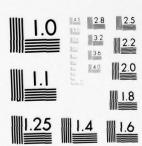
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# JPPLEMENTARY

# INFORMATION

# LIST OF CORRECTIONS TO NSWC/DL TR-3866

Page 3: In Table 1, change amplitude K of  $S_2$  to:  $K(S_2) = 0.112841$ and D to: D = d + 365(y - 1975) + Int[(y-1973)/4]

Page 22: In Eq. 14 change  $\eta_0$  to:  $\eta_0 = \frac{1}{2}K(3\cos^2 \theta - 2) \cos (\sigma t + \chi)$ 

Page 27: Top line change  $\beta$  to:  $\beta$  = 0.90 In first equation of Eq. 32 add:  $\overline{H}_{\lambda}$ =  $H_{\lambda}/H$ ,  $\overline{H}_{\lambda\lambda}$  =  $H_{\lambda\lambda}/H$ .

Page 29: Change Eq. 37c to:  $\frac{1}{2}K(3 \sin^2 \theta - 2)$ .

Page 31: In V-equation of Eq. (43) change  $\sin \theta$  to  $\sin 2\theta$ .

Page 32: Change Eq. 45 to:

 $\eta = \frac{1}{2}K(3\sin^2\theta - 2) e^{i\sigma t}$   $= K_1(3\sin^2\theta - 2) e^{i\sigma t}$   $U = \frac{1}{2}K_2 \sin \theta e^{i\sigma t}$   $V = \frac{-i}{2}K_1\sigma R \sin 2\theta e^{i\sigma t}$ 

In Equation 45a change  $\overset{\sim}{\textbf{k}}_2$  and  $\overset{\sim}{\textbf{k}}_1$  to:

Note: A higher order approximation of improved accuracy is in preparation.

Page 33: Change  $\hat{K}_1$  to:  $\hat{K}_1 = -3\overline{K}$ .

Page 47: In  $B_{m,n}$  change for  $\gamma = 0$ :

-3  $\sin 4n\Delta\theta$  to  $+6\sin 4n\Delta\theta$ 

Page 52: In Eq. 84 change the left-side  $\xi$  to  $\zeta$ .

Page 56: In Eq. 94 change  $k_1$  and  $k_2$  to:  $k_1 = .06$ ,  $k_2 = .03$ .

